

Analysing the PDF of density fluctuations - can it work in real data?

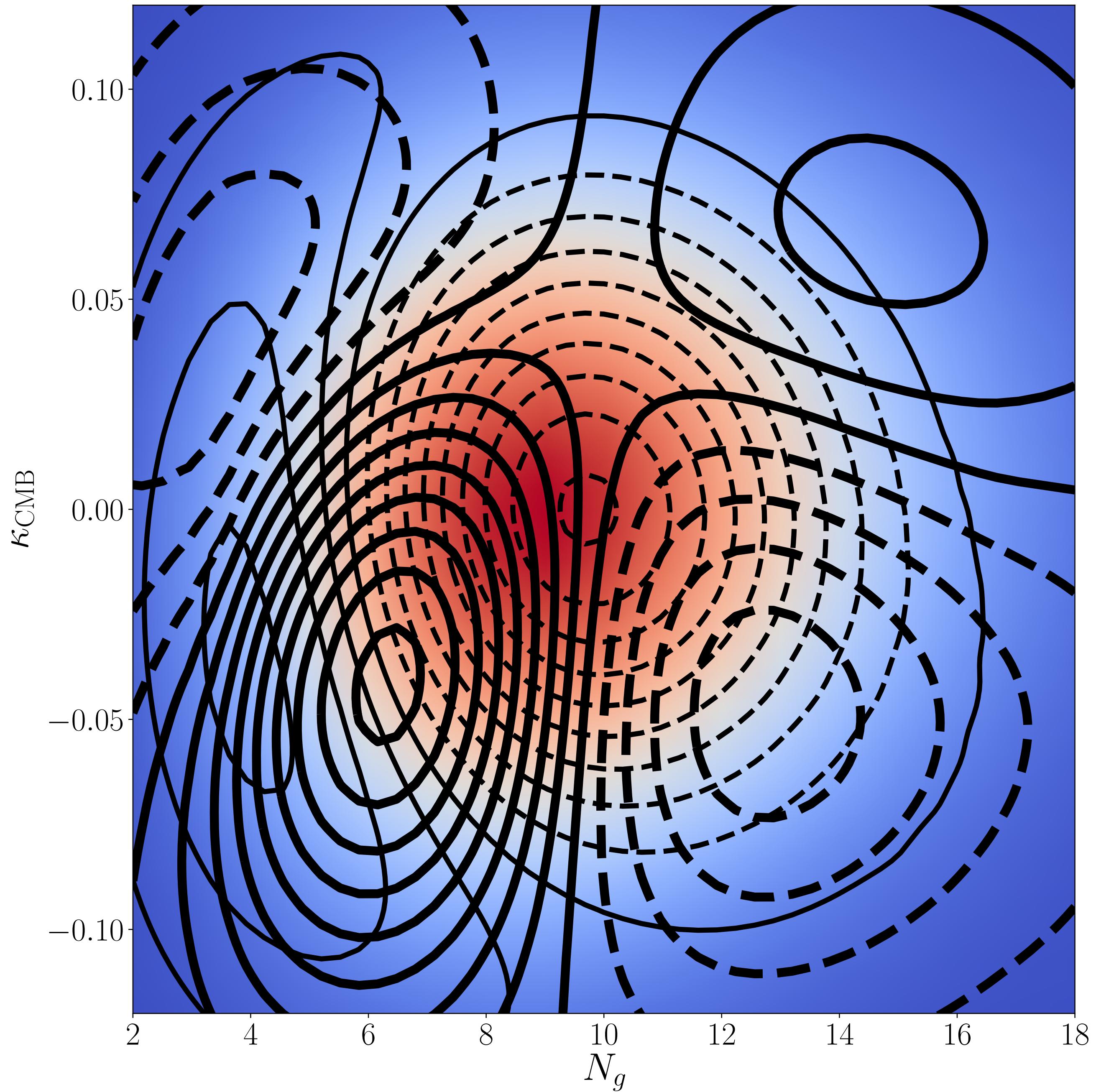
Focus week

“Interpretable higher-order statistics”

Trieste, 30 June 2022

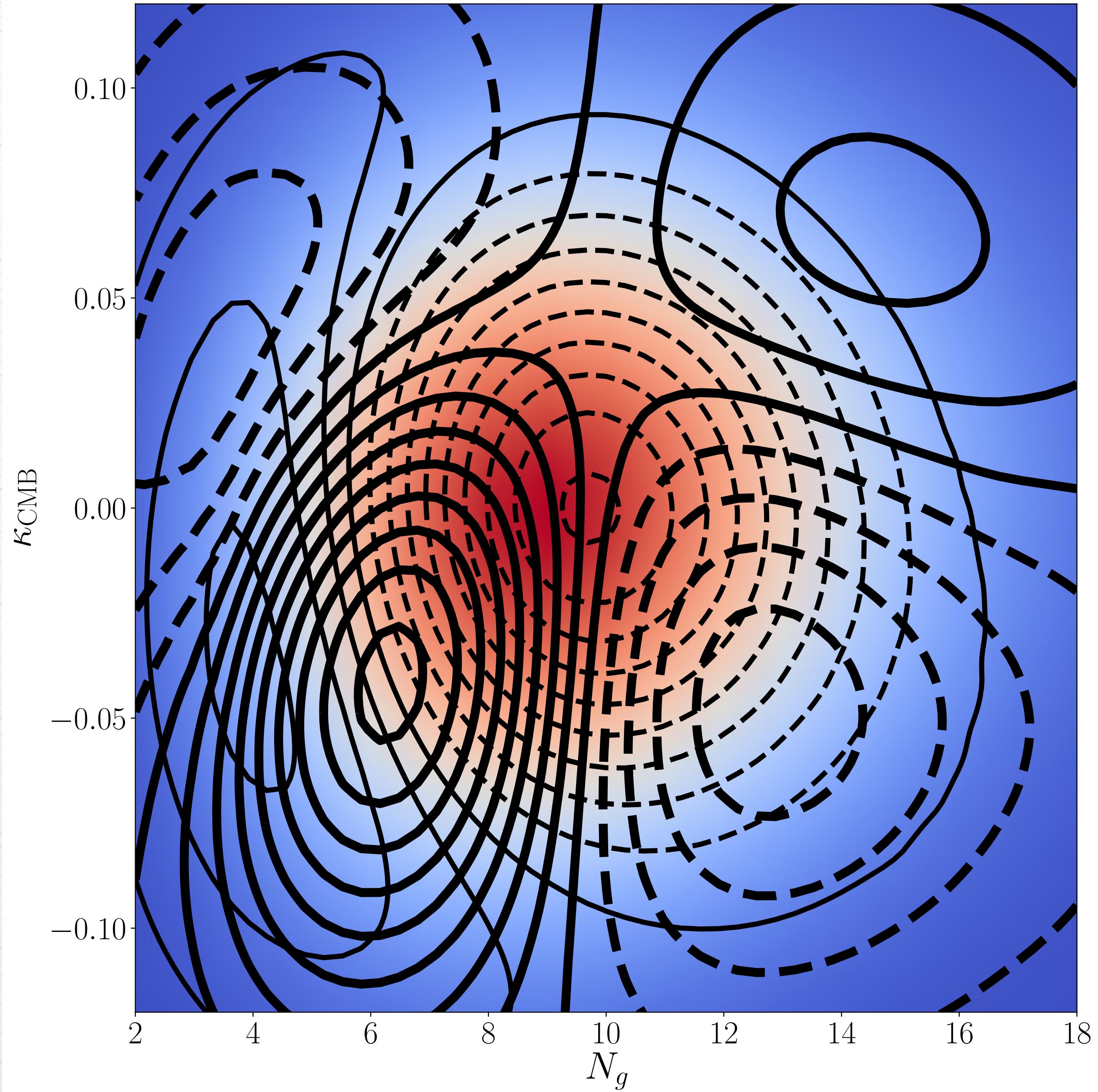
Oliver Friedrich
LMU Munich

*work with Anik Halder, Cora Uhlemann,
Aoife Boyle, Alexandre Barthelemy,
Sandrine Codis, Daniel Gruen,
Joe DeRose, Elisabeth Krause,
Tobias Baldauf, Francisco
Villaescusa-Navarro, Marc Manera,
Takahiro Nishimichi and more!*



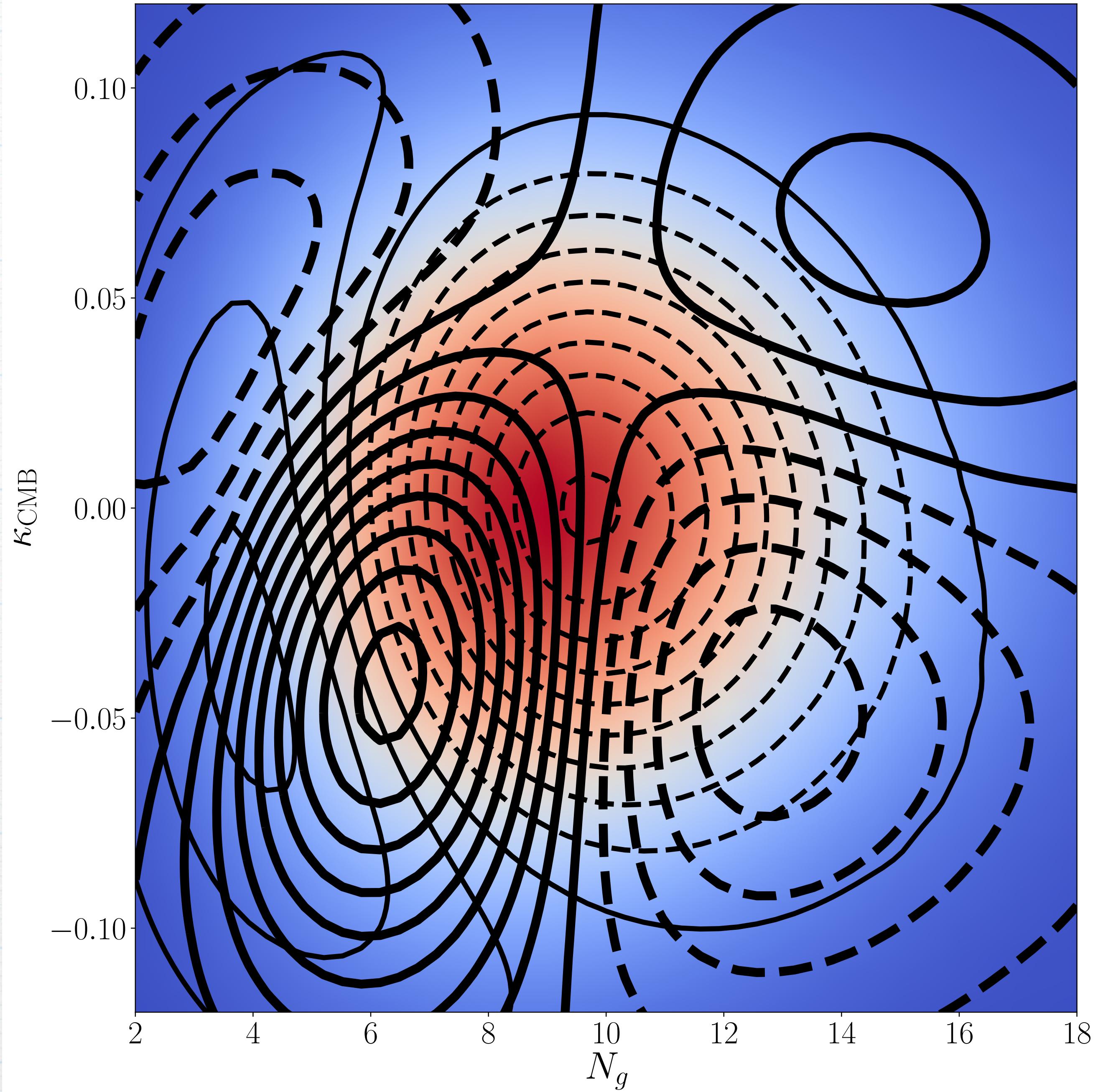
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- Why the PDF?
- PDF vs. $P(k)$ — the details
- Let's interpret the PDF
- My view of PDF theory

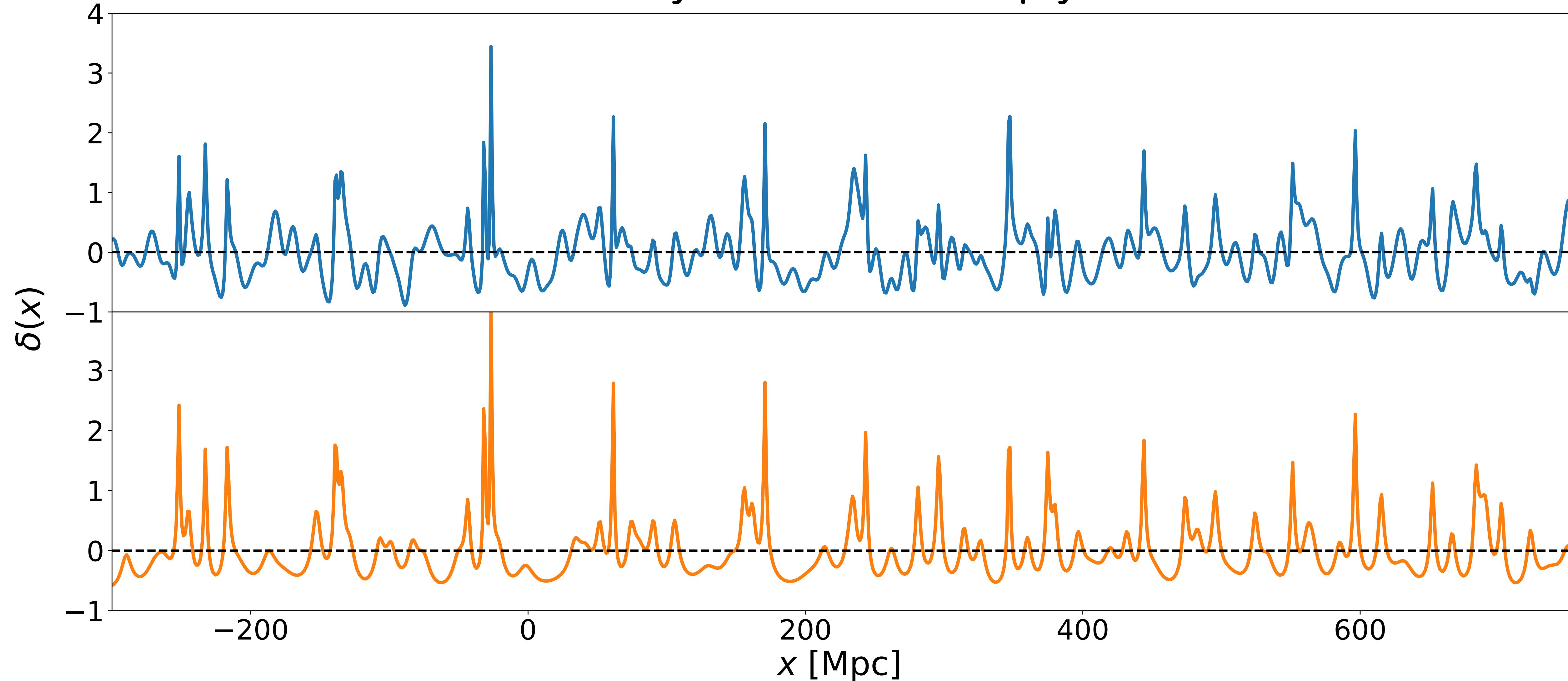


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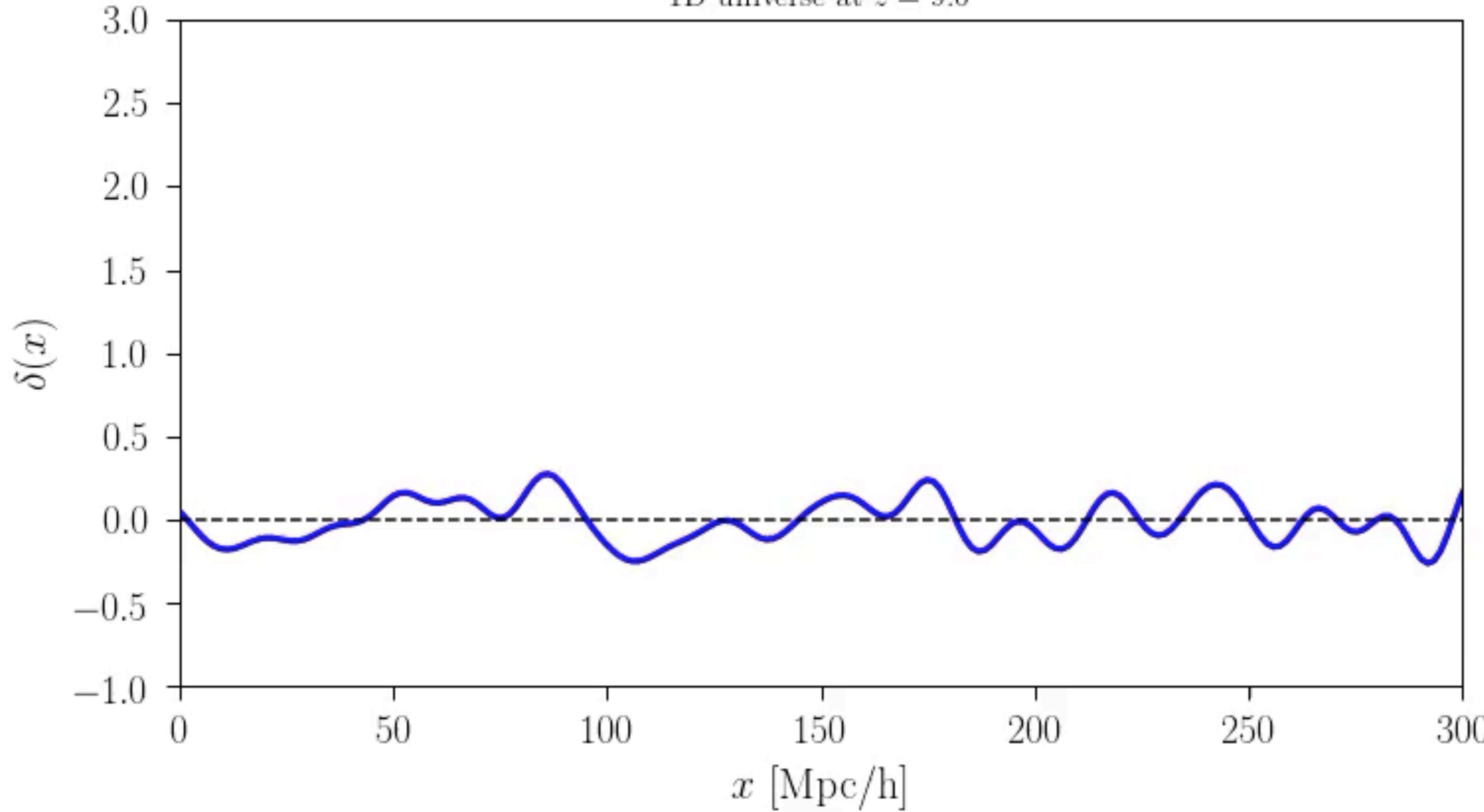


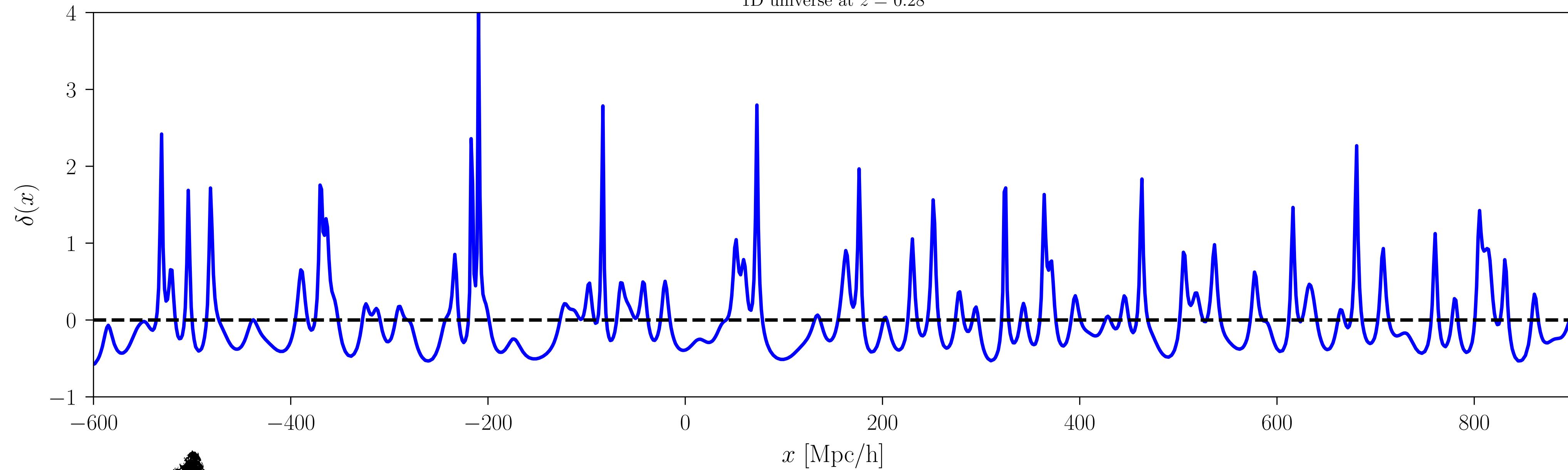
two “density fields” - which one is physical?



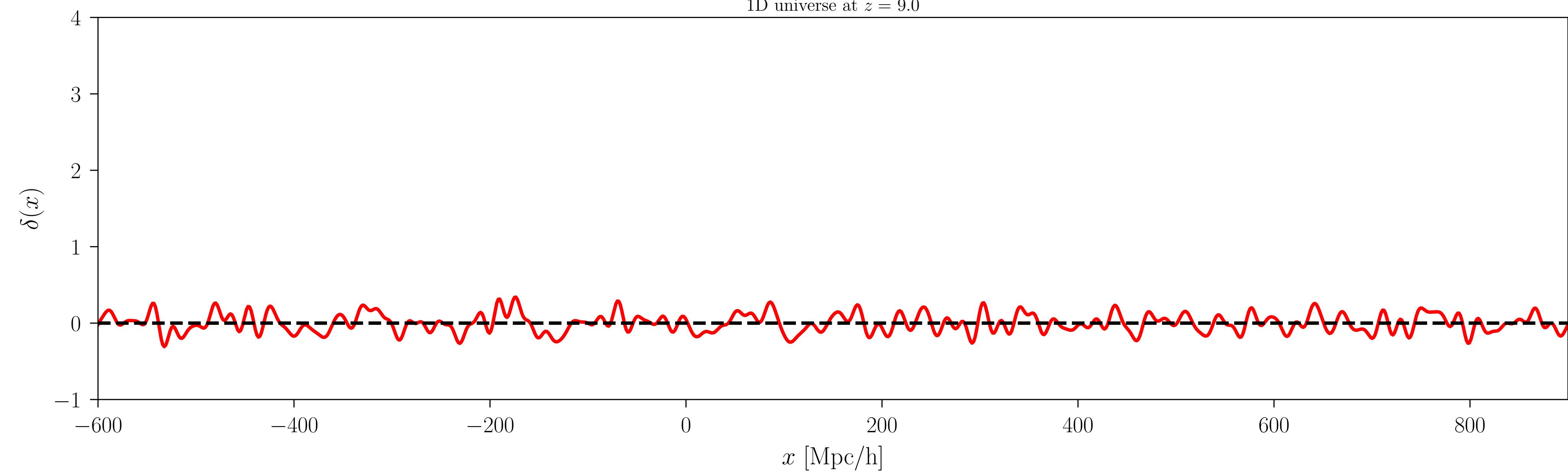
$$\delta_m = (\rho_m - \bar{\rho}_m)/\bar{\rho}_m$$

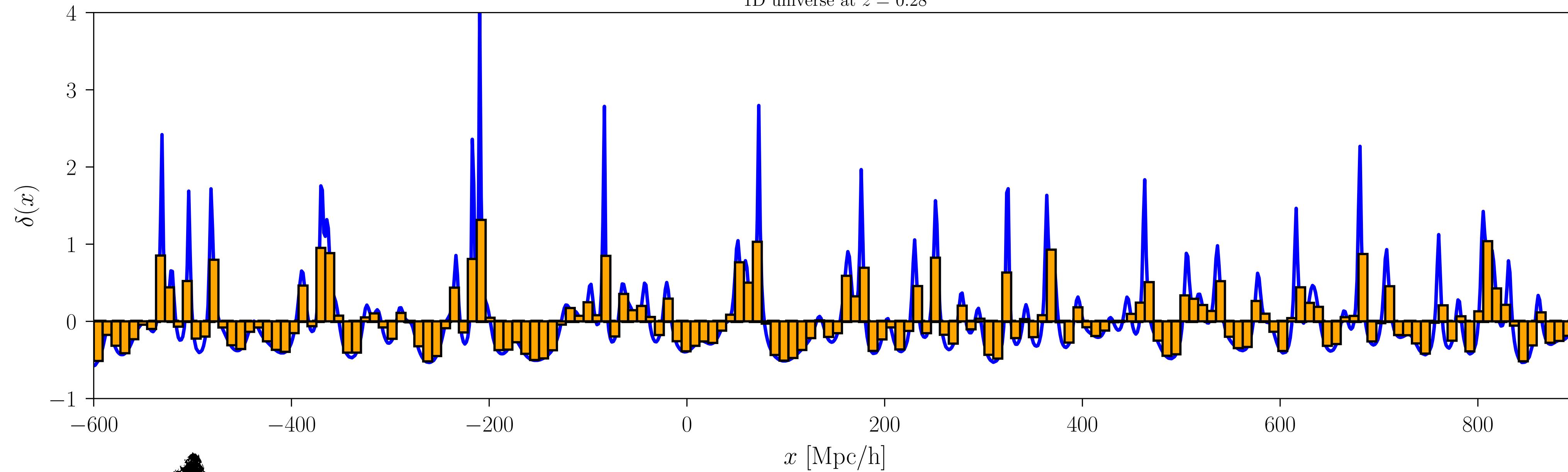
1D universe at $z = 9.0$



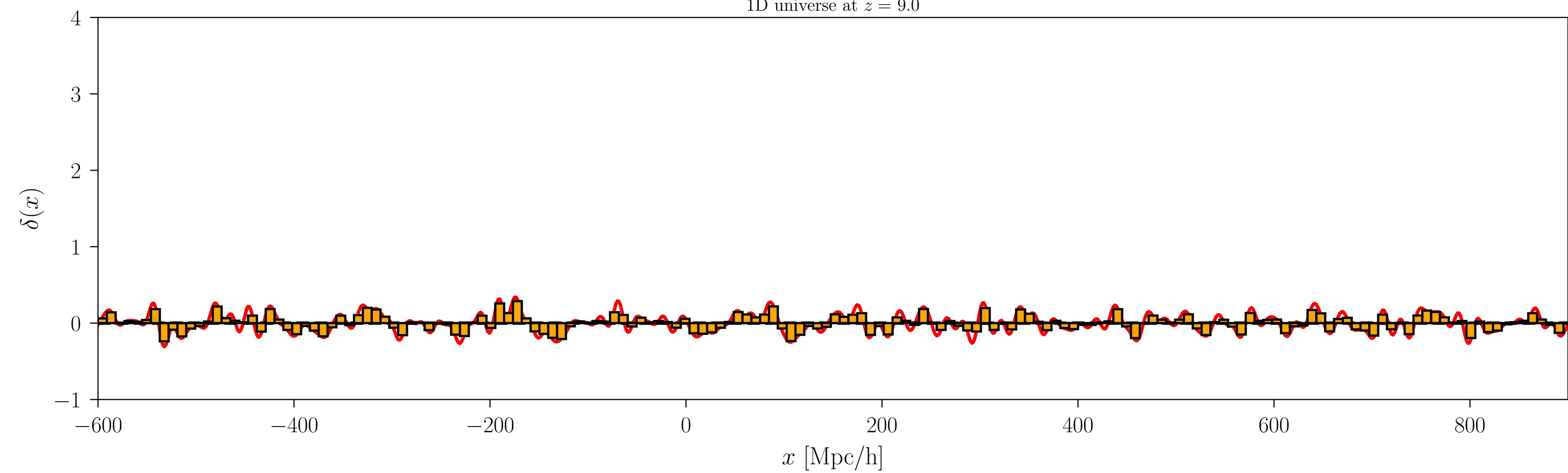
1D universe at $z = 0.28$ 

gravitational collapse

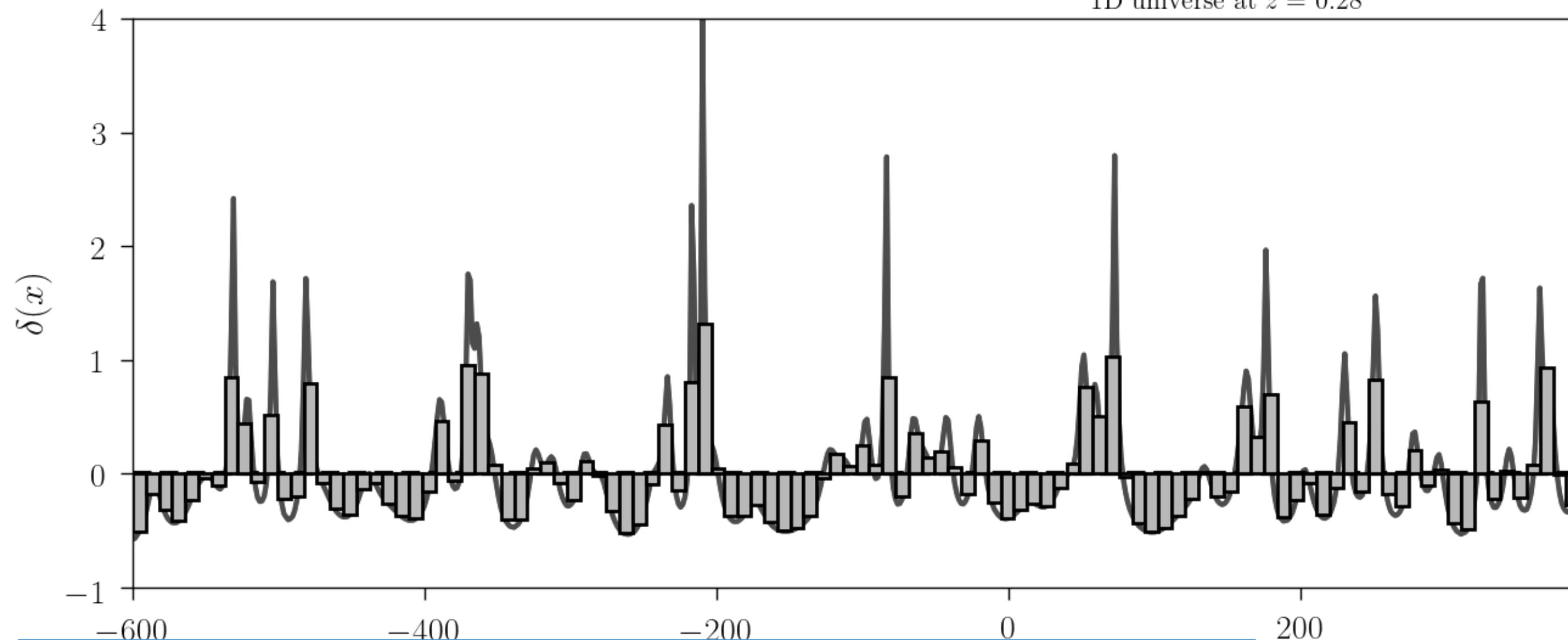
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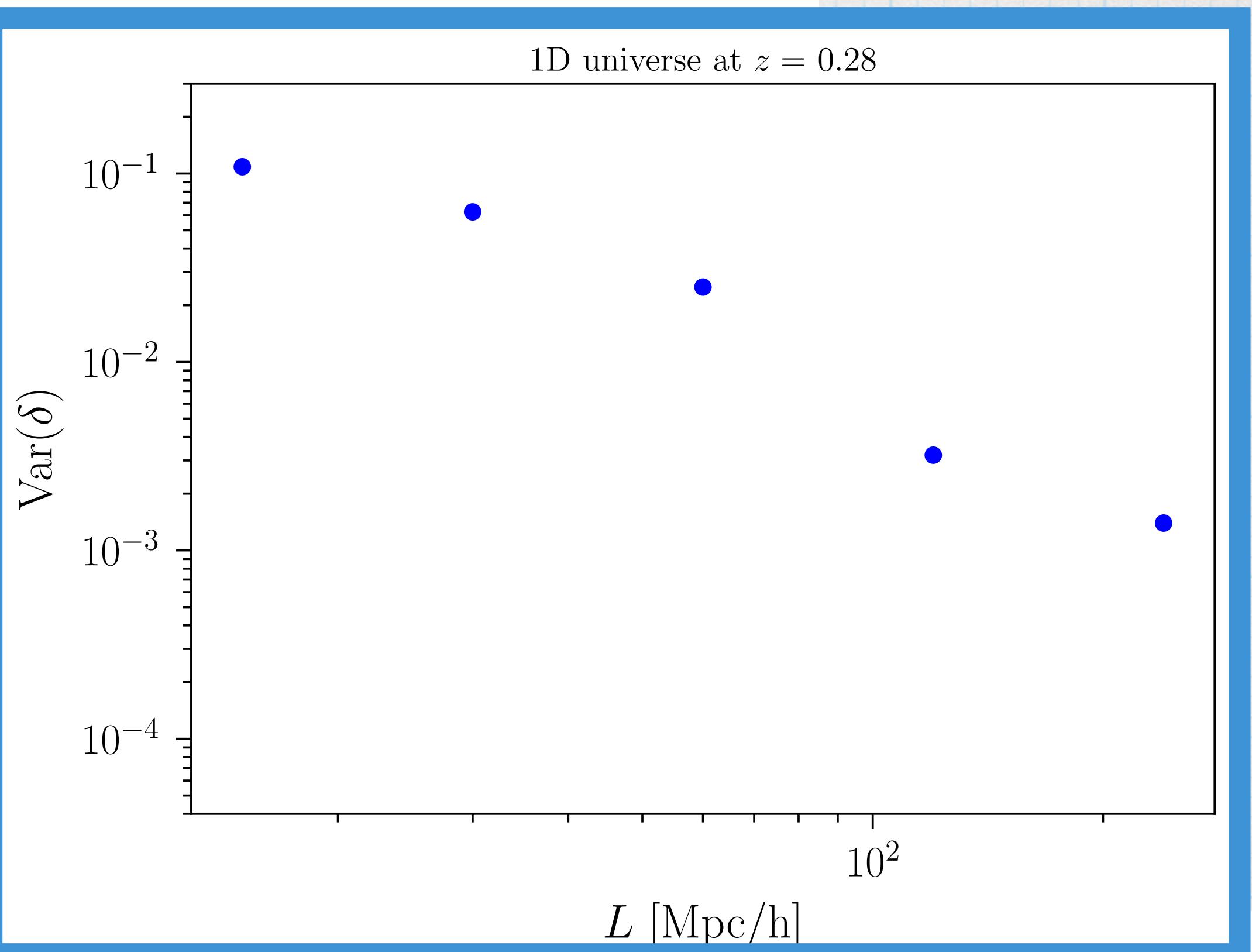
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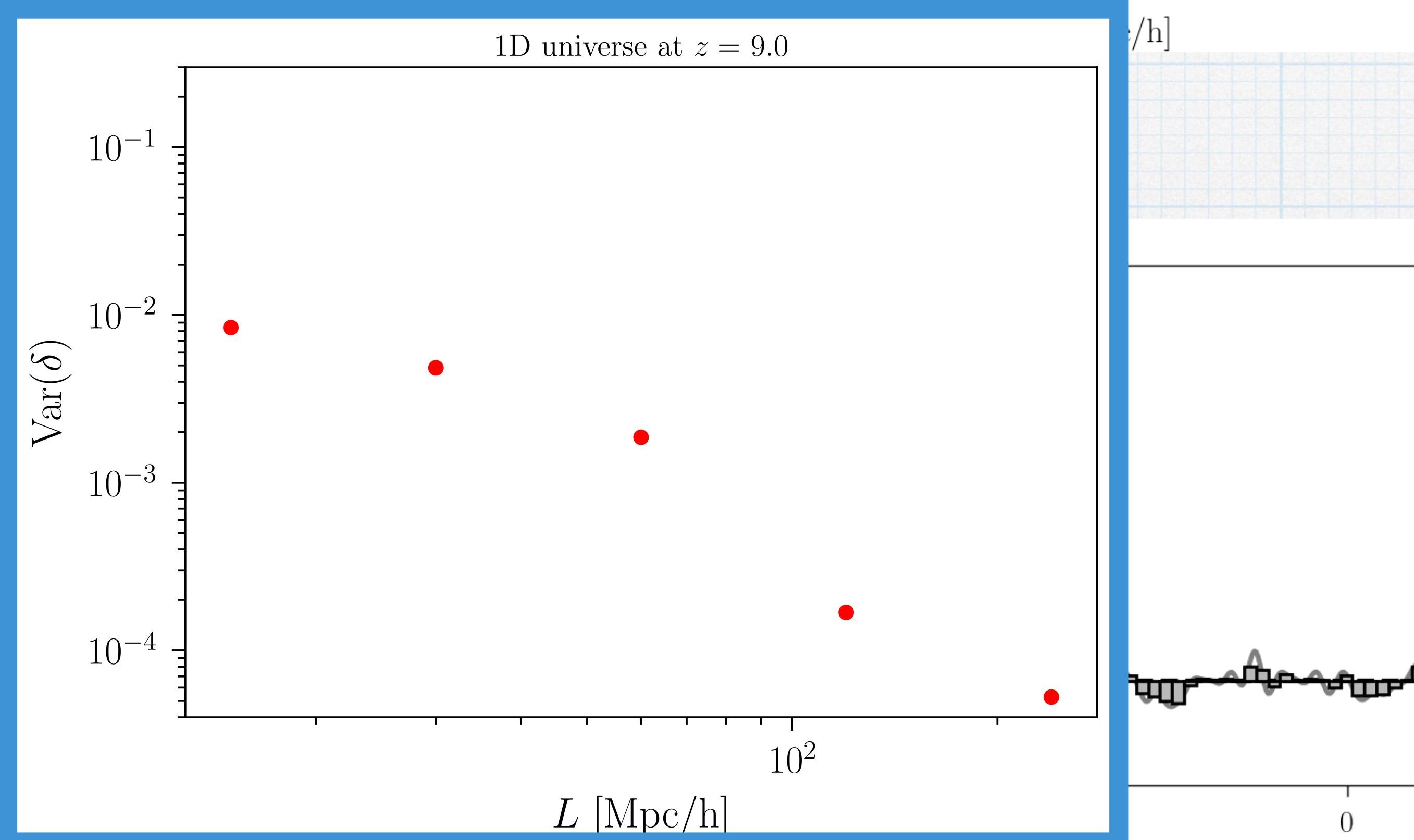
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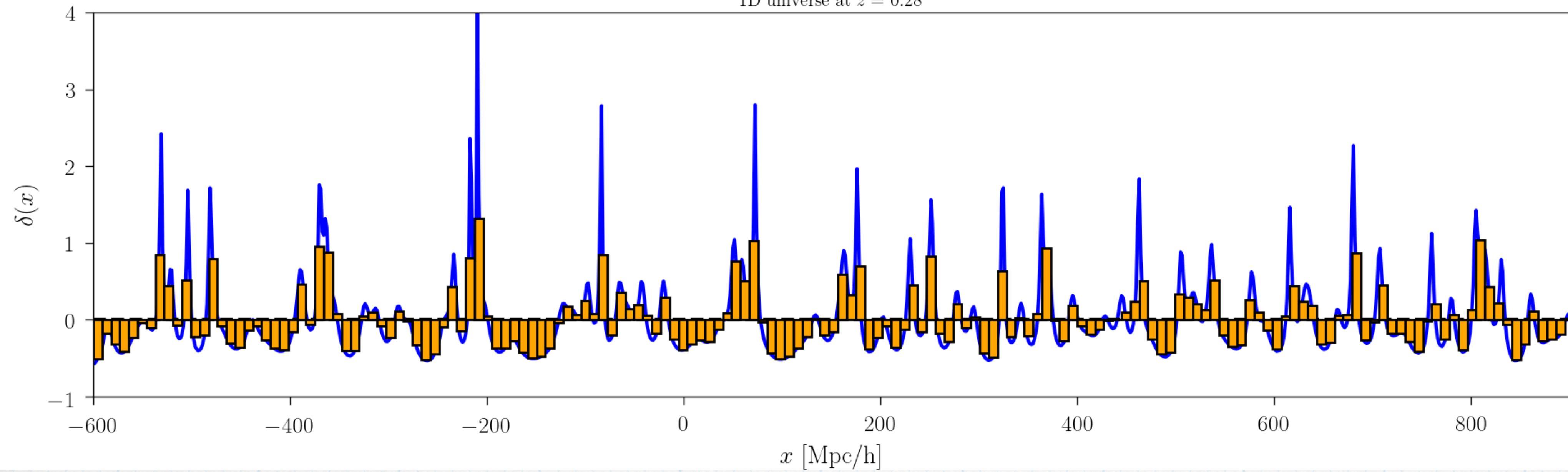
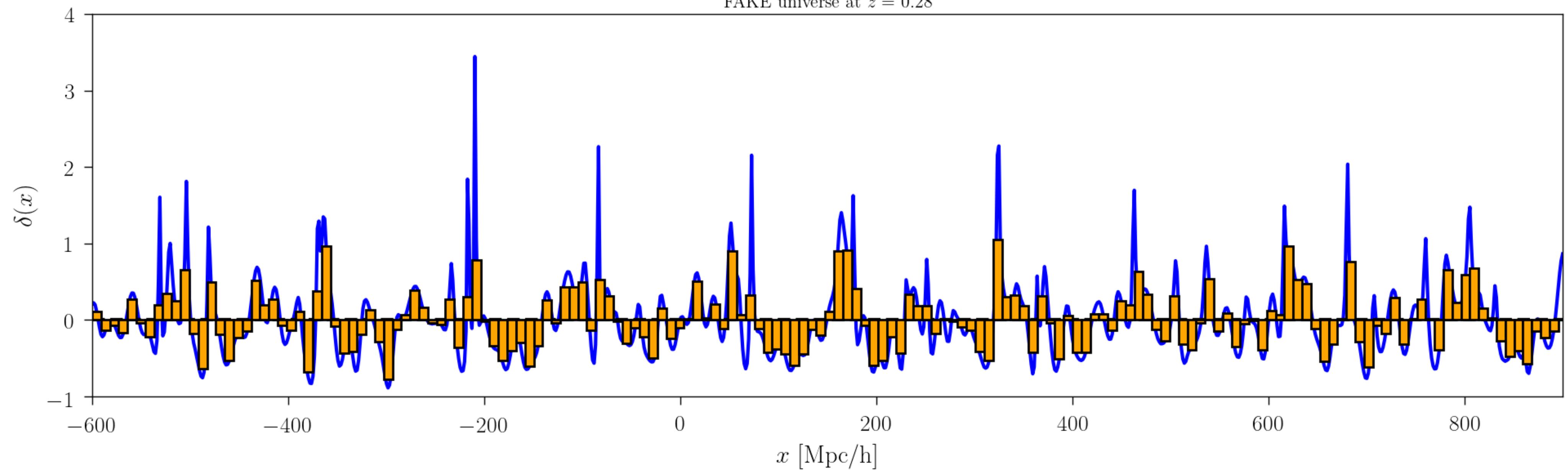


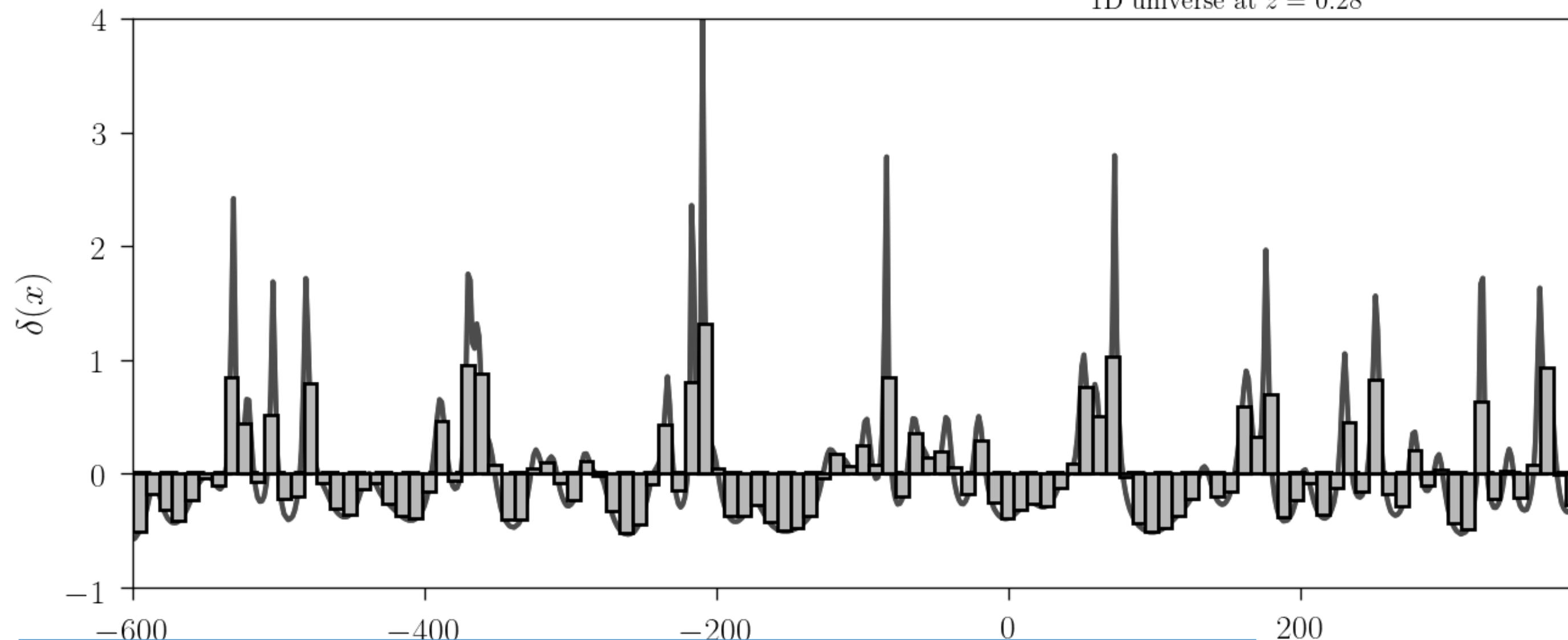
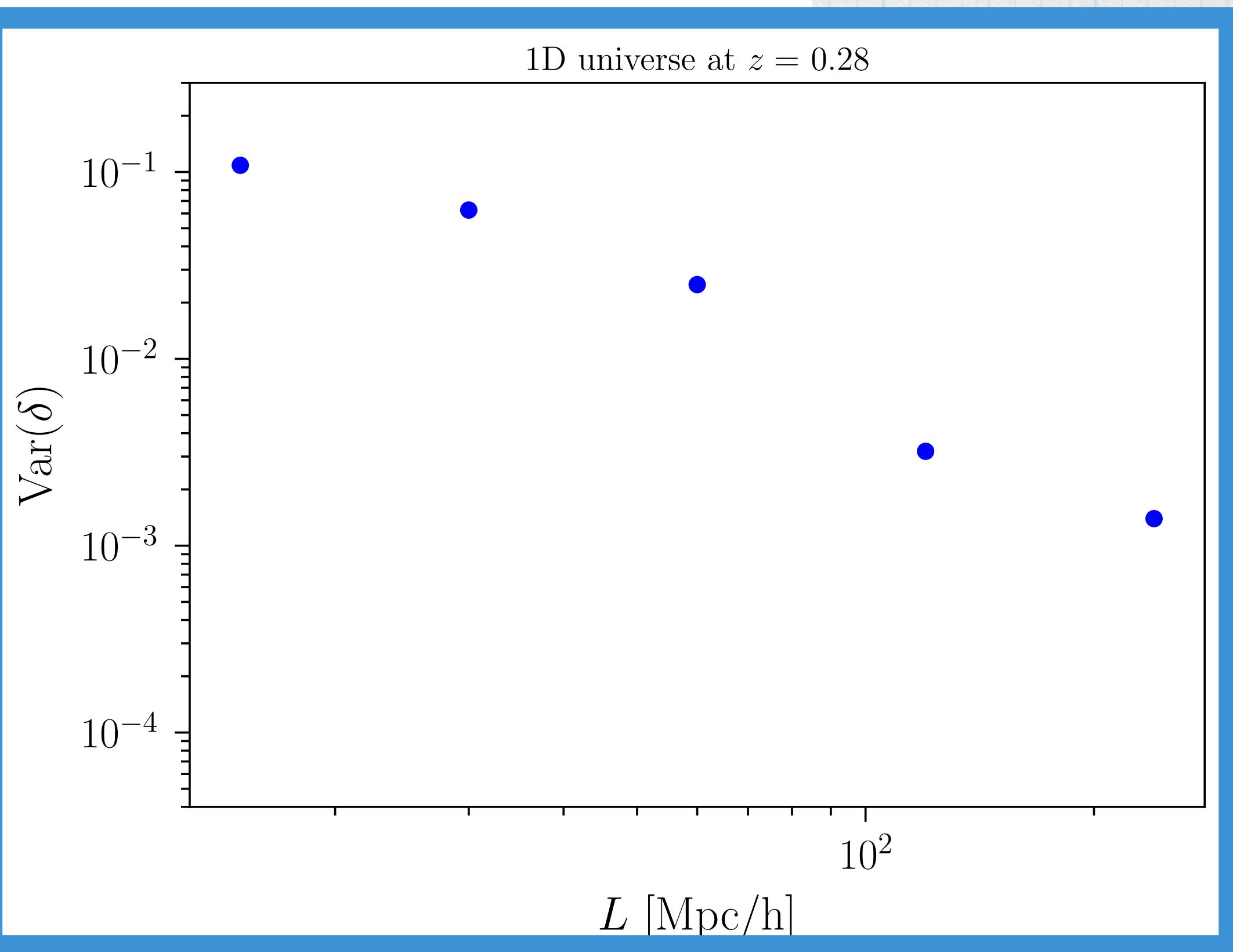
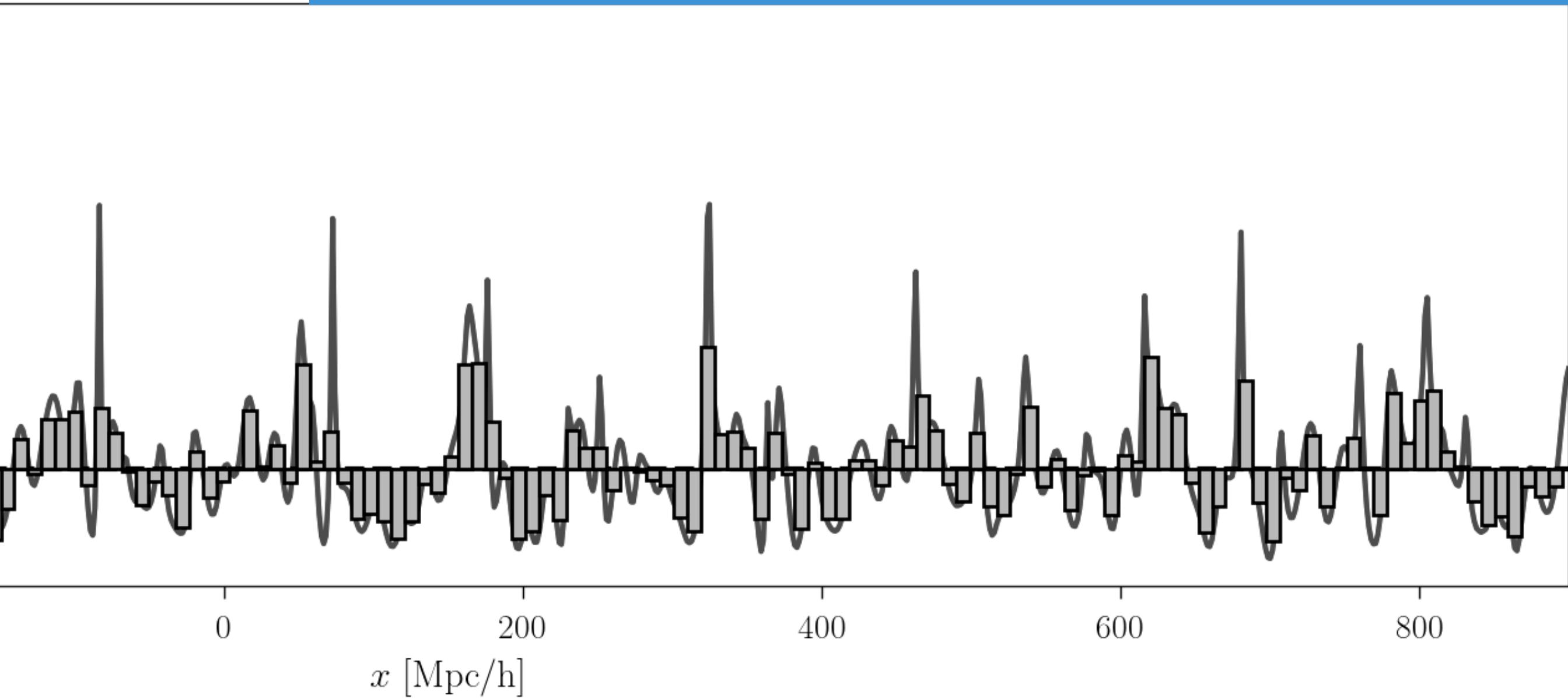
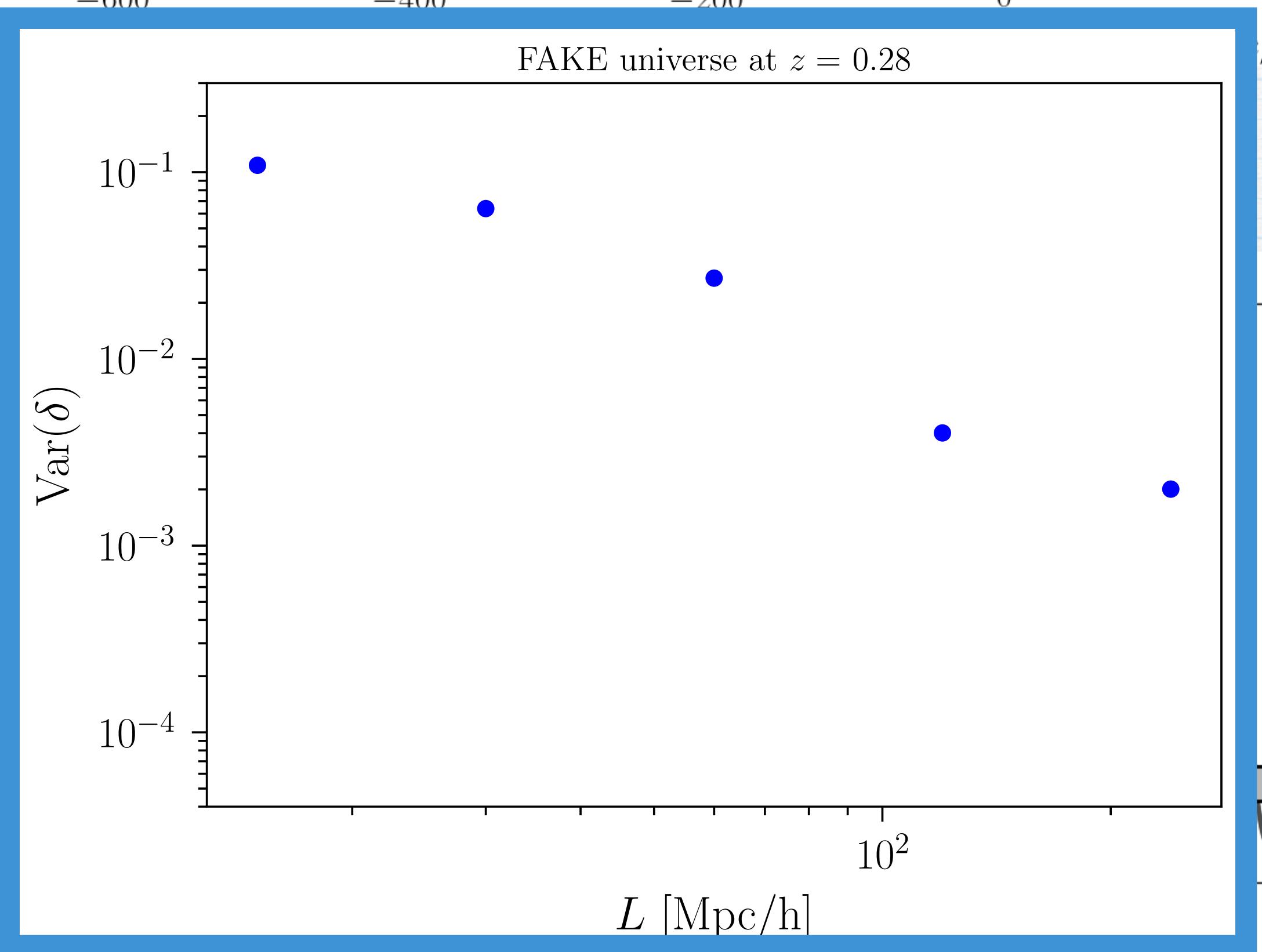
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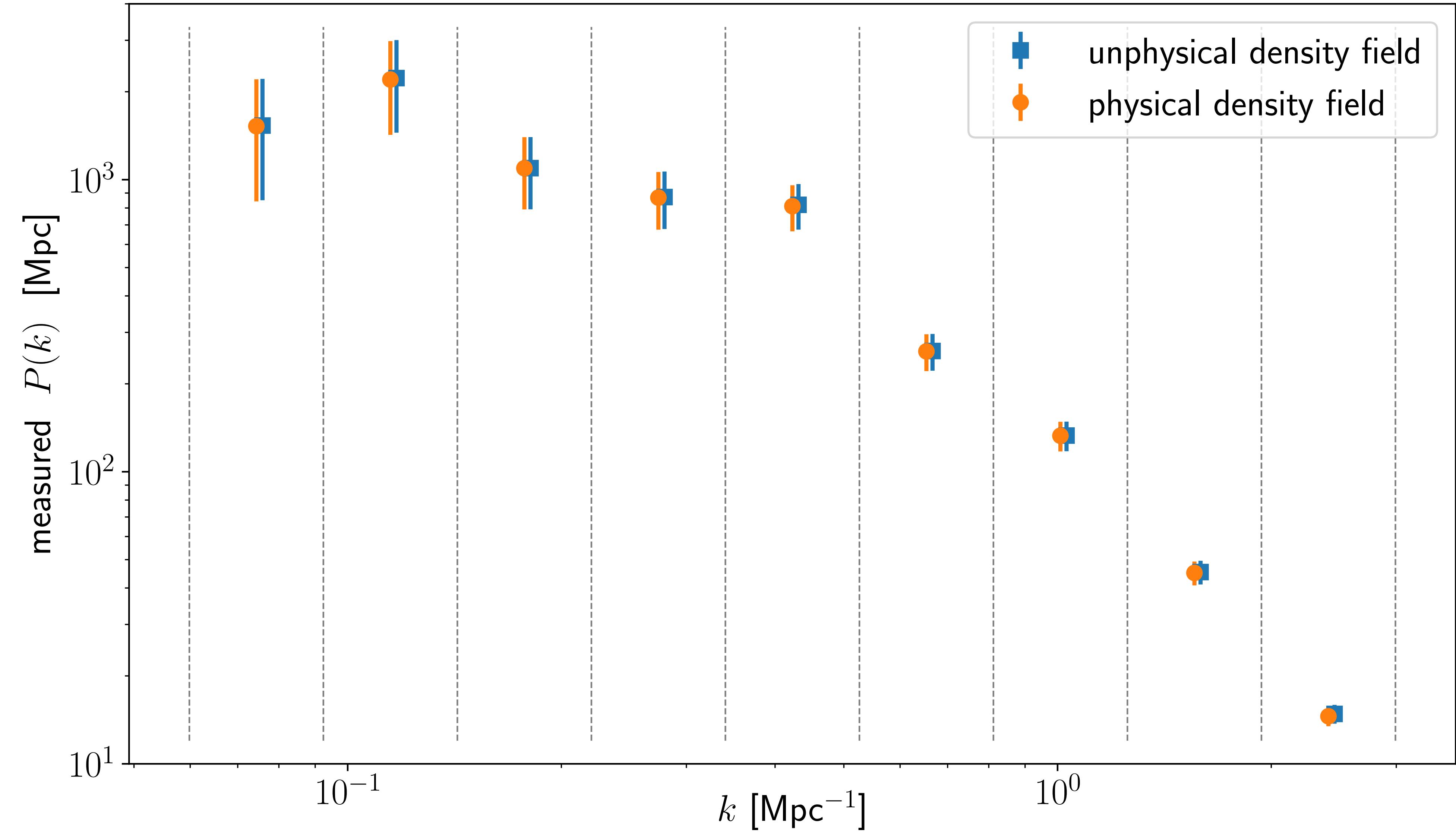


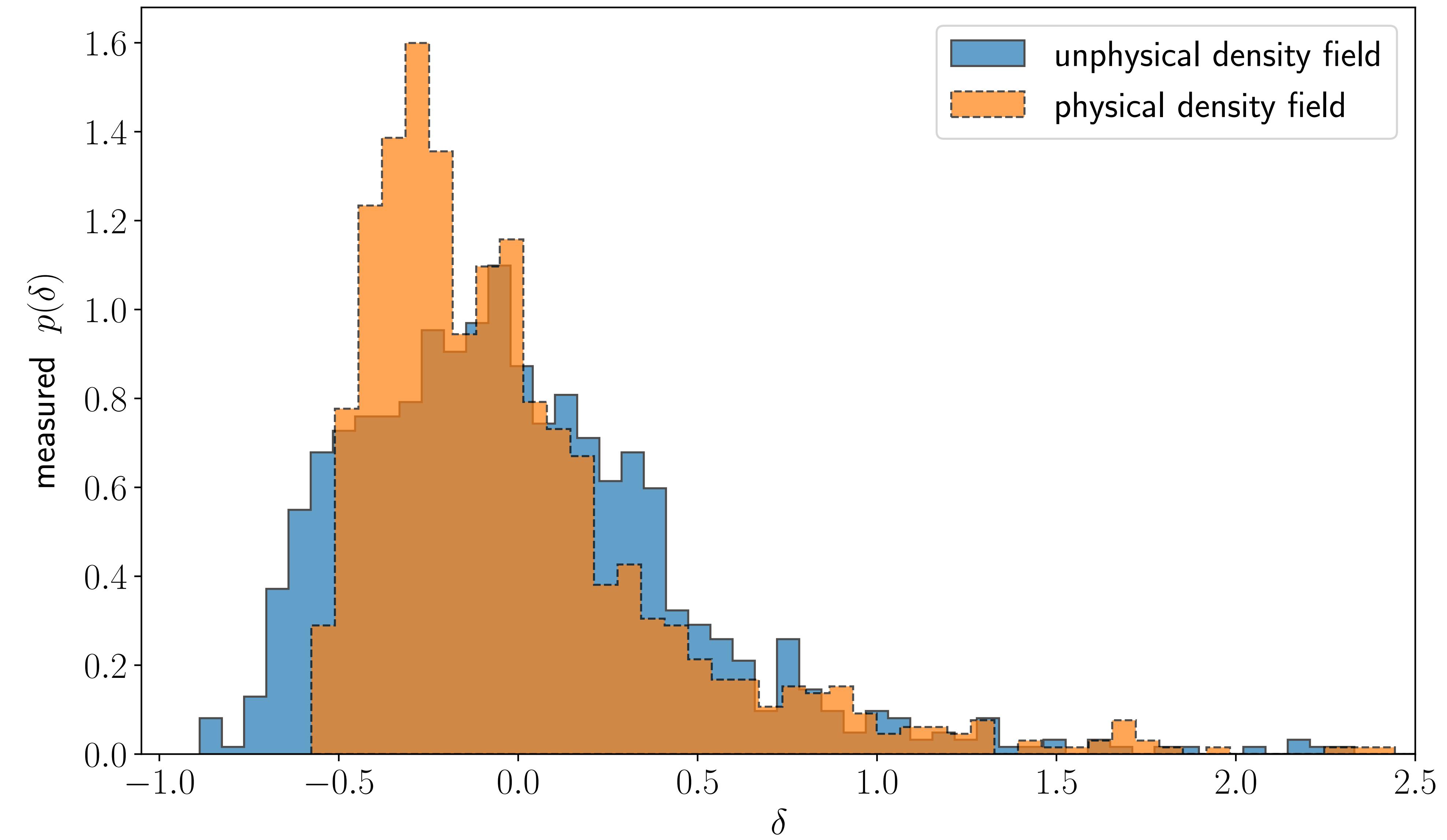
1D universe at $z = 9.0$



1D universe at $z = 0.28$ FAKE universe at $z = 0.28$ 

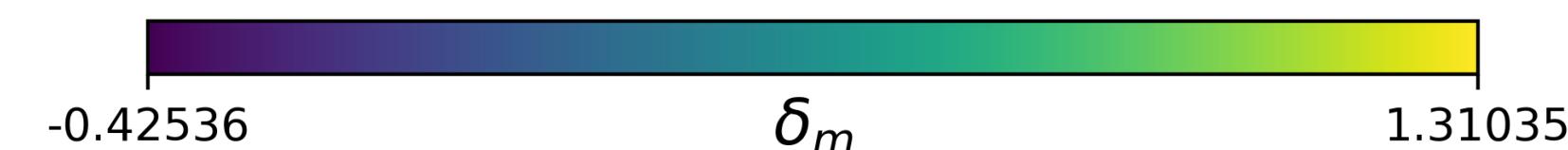
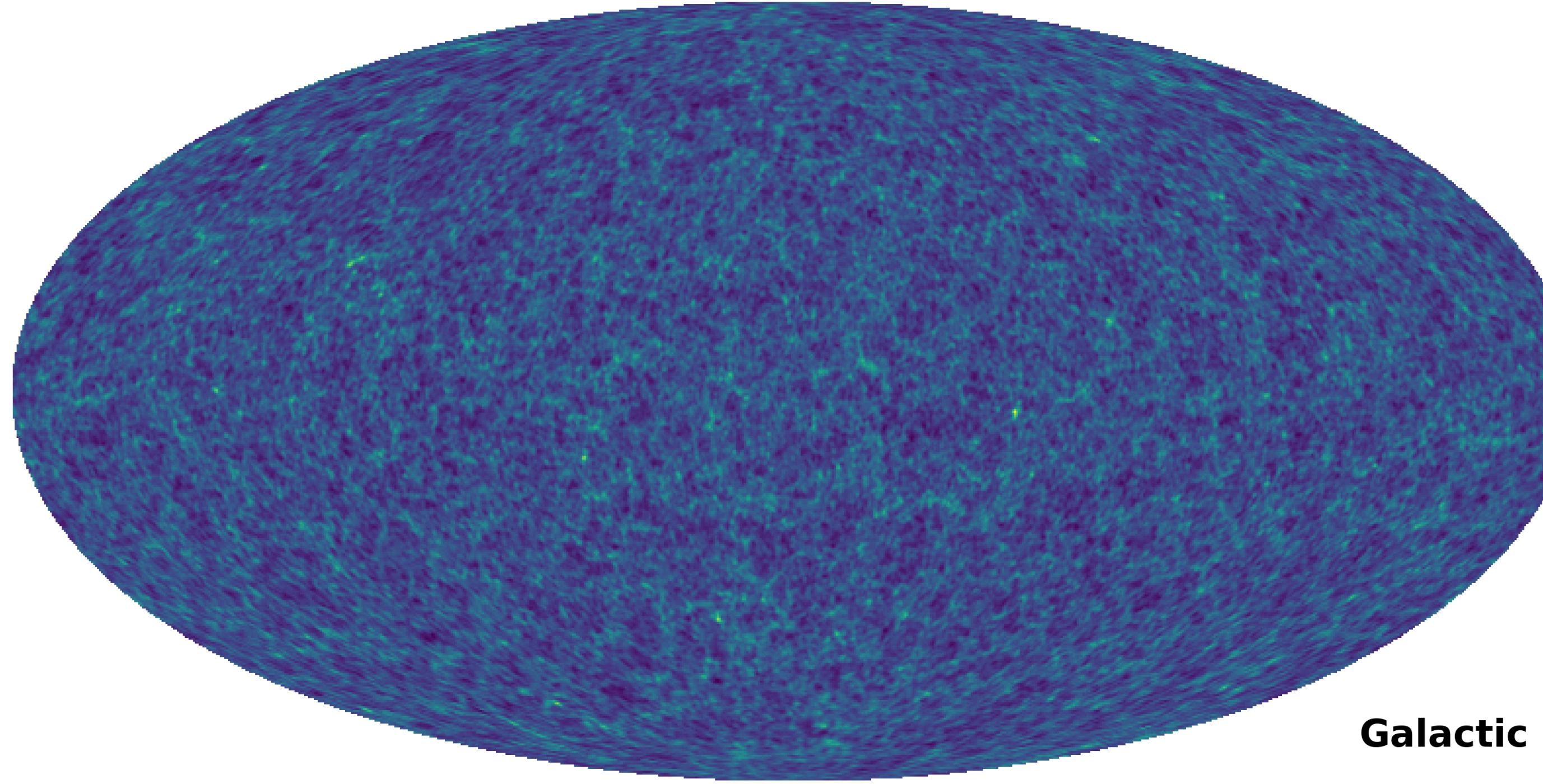
1D universe at $z = 0.28$ 1D universe at $z = 0.28$ FAKE universe at $z = 0.28$ 



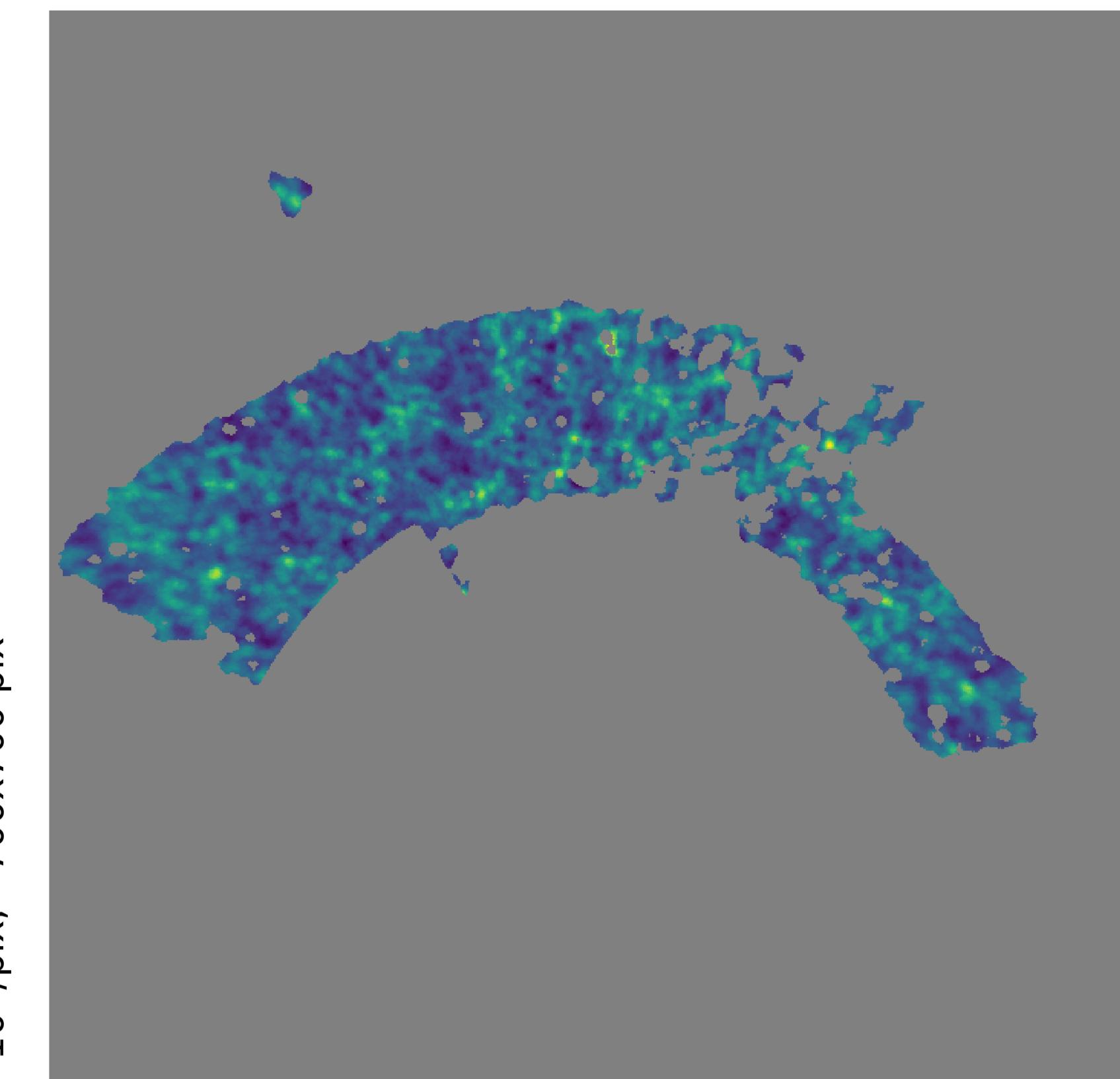


Matter density contrast in Buzzard N-body simulation:

matter density contrast in Buzzard N-body simulation



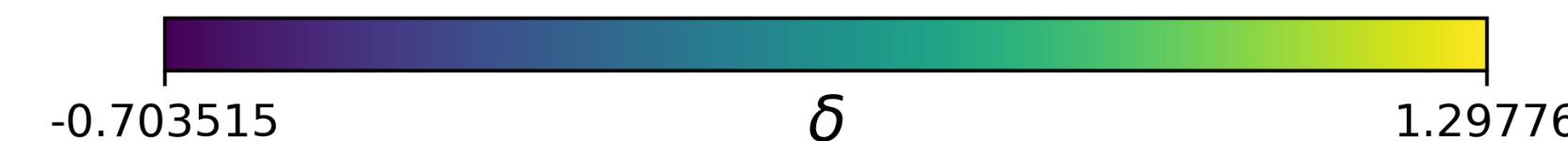
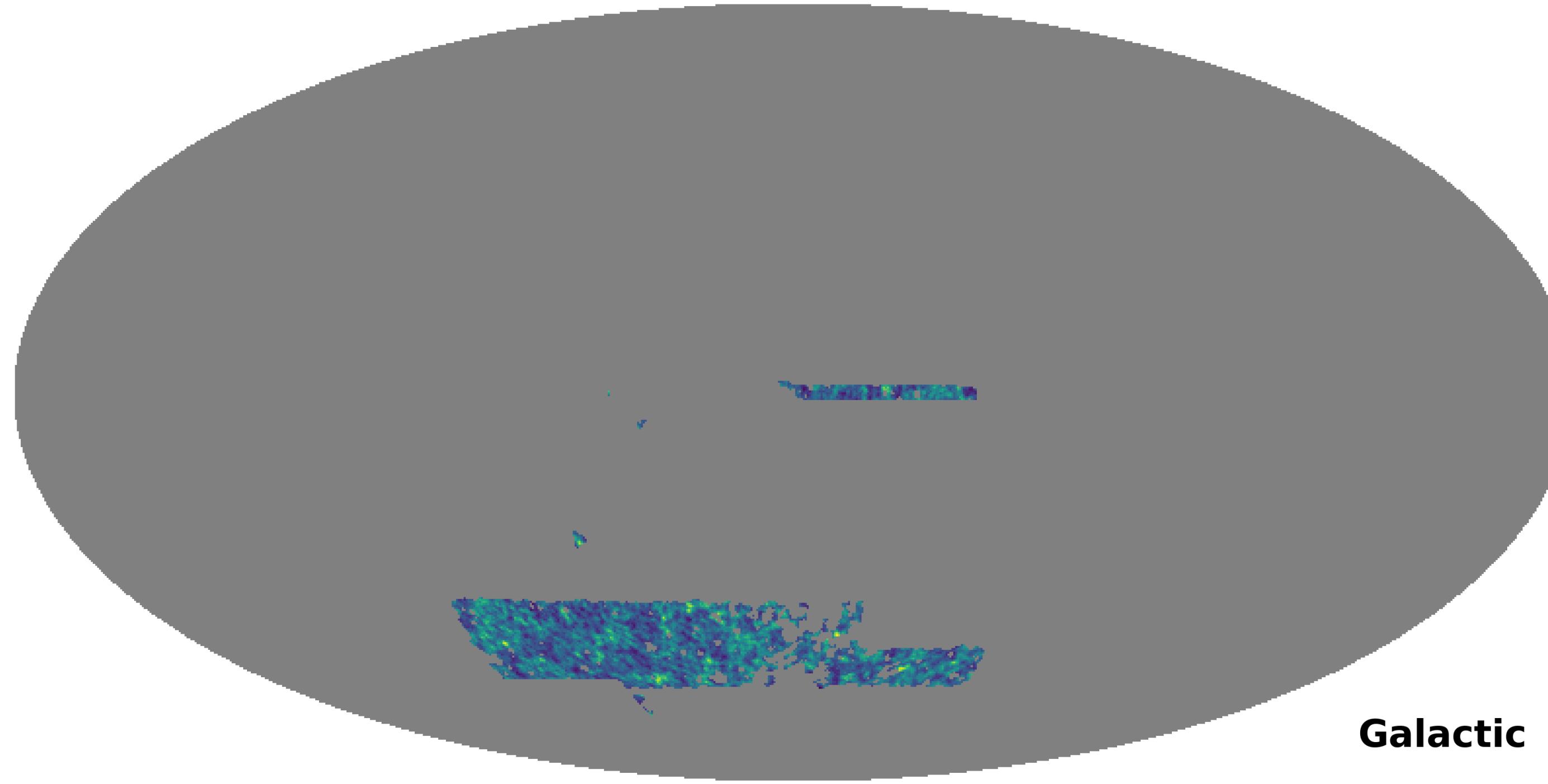
matter density contrast in Buzzard N-body simulation



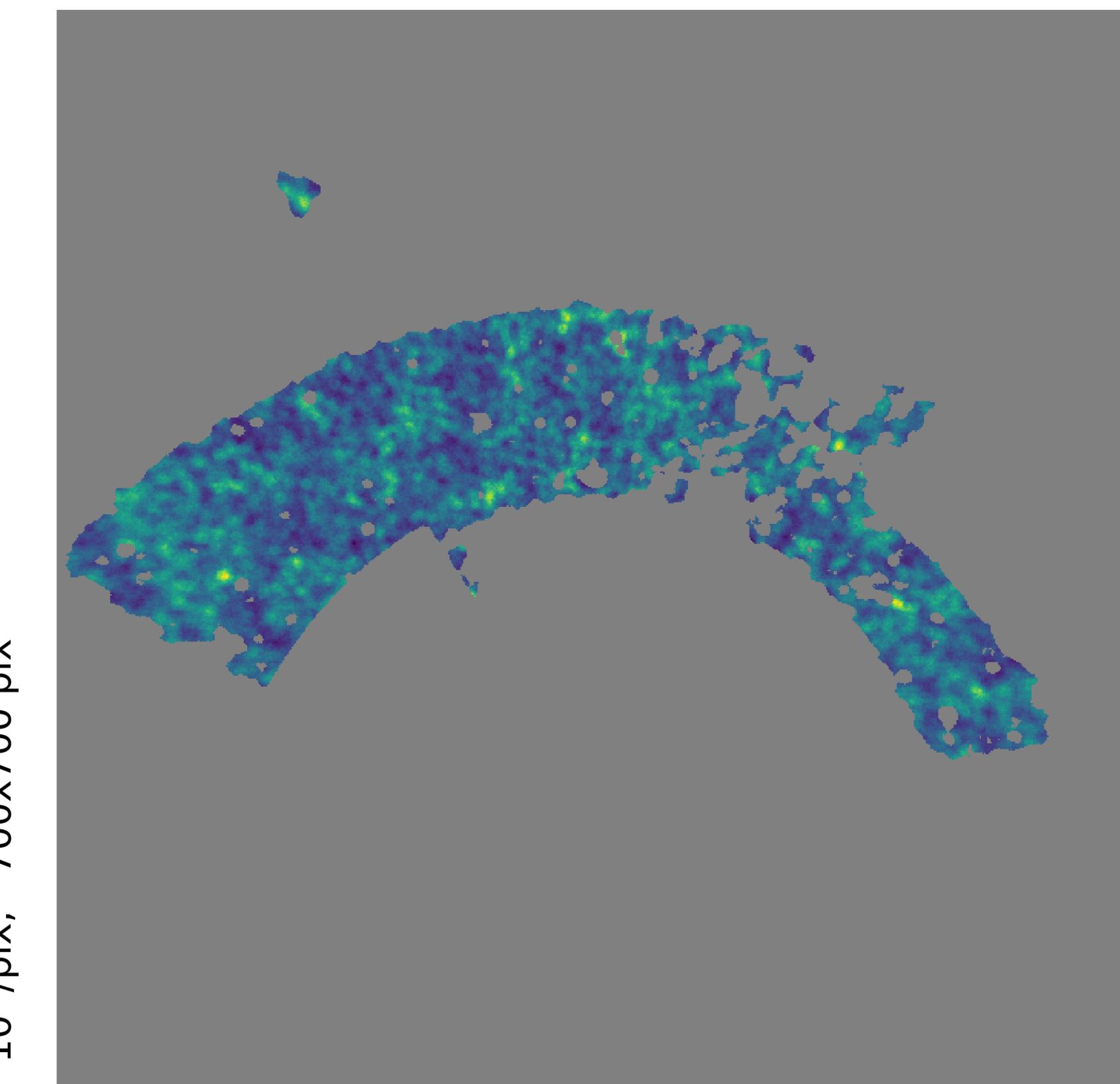
Y1 Buzzard Flock, DeRose++ (2019)

Galaxy density contrast in Buzzard N-body simulation:

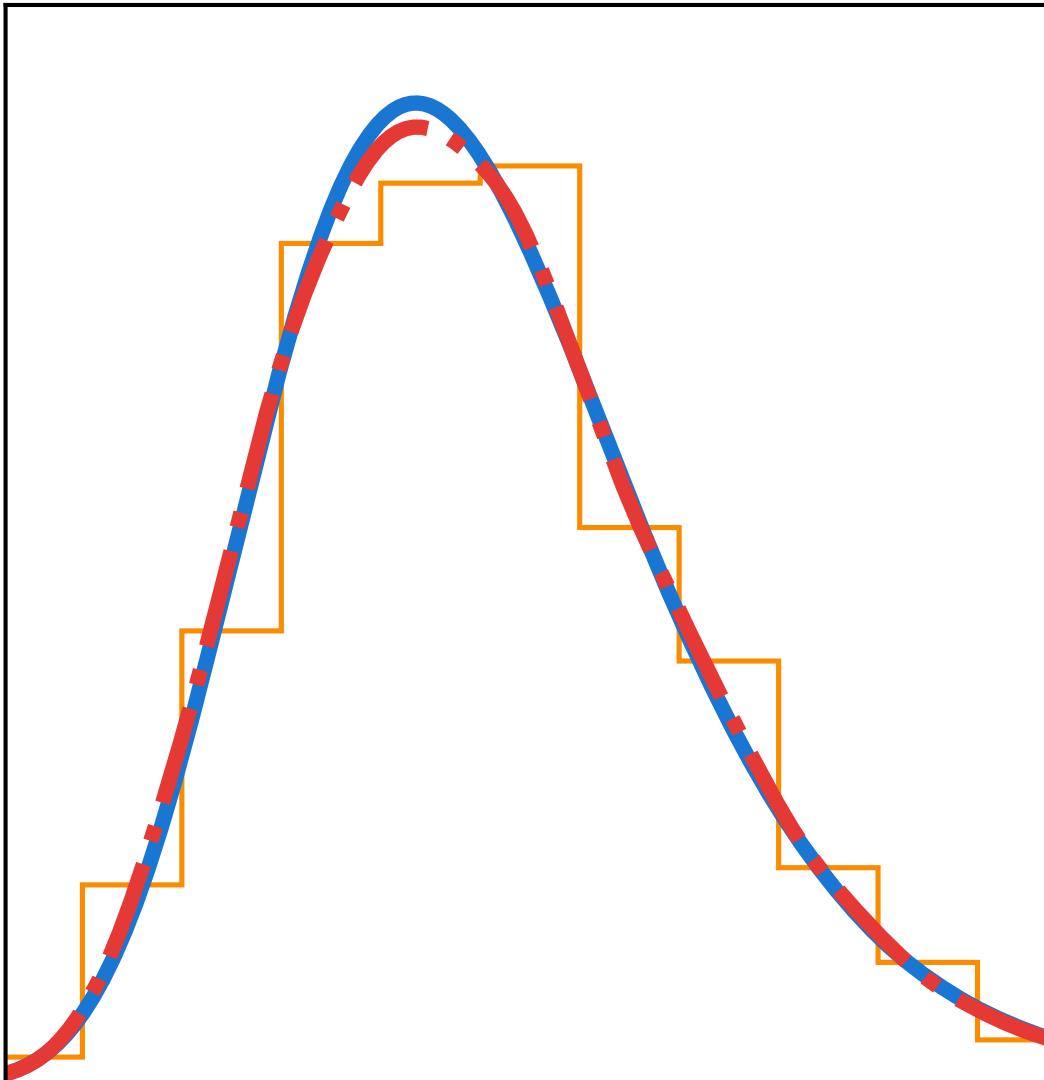
density contrast of red galaxies in Buzzard N-body simulation



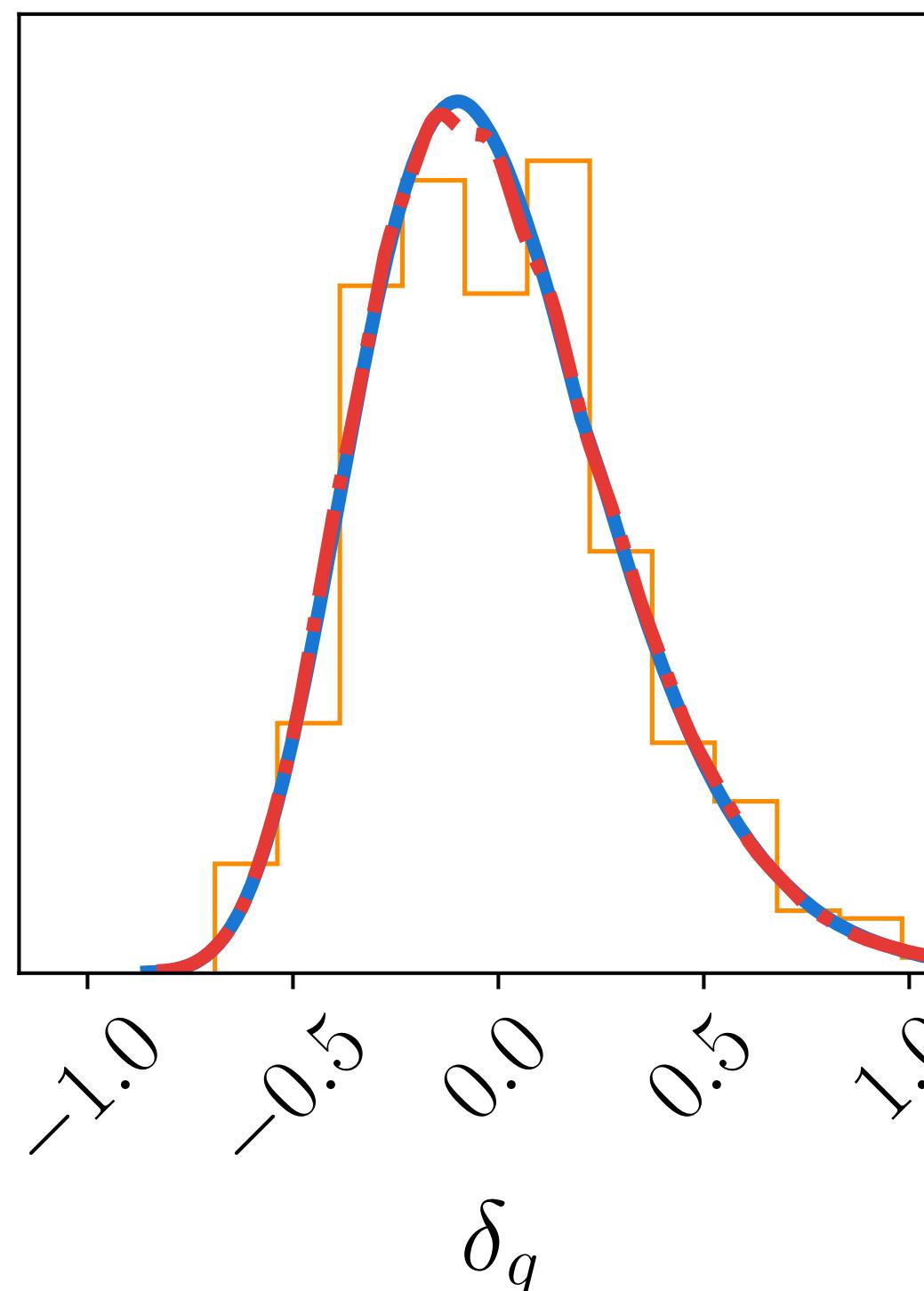
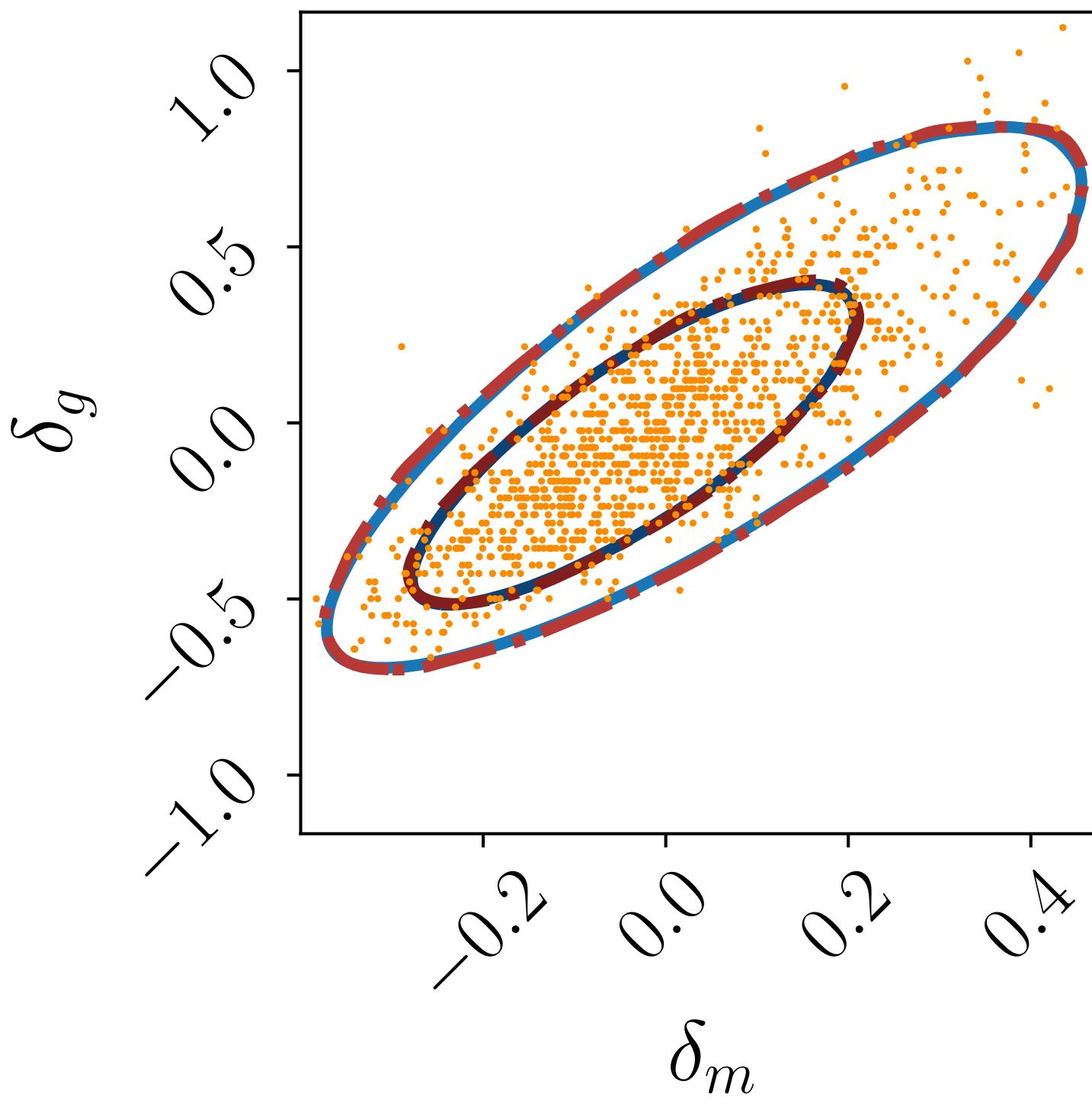
density contrast of red galaxies in Buzzard N-body simulation



Y1 Buzzard Flock, DeRose++ (2019)



$p(\delta_m, \delta_g)$
 buzzard (contours)
 buzzard (rand. subset)



galaxies are biased & stochastic
tracers of matter density

2-point statistics only extracts
three numbers from this!

→ way to do better: analyse full shape
of the PDF

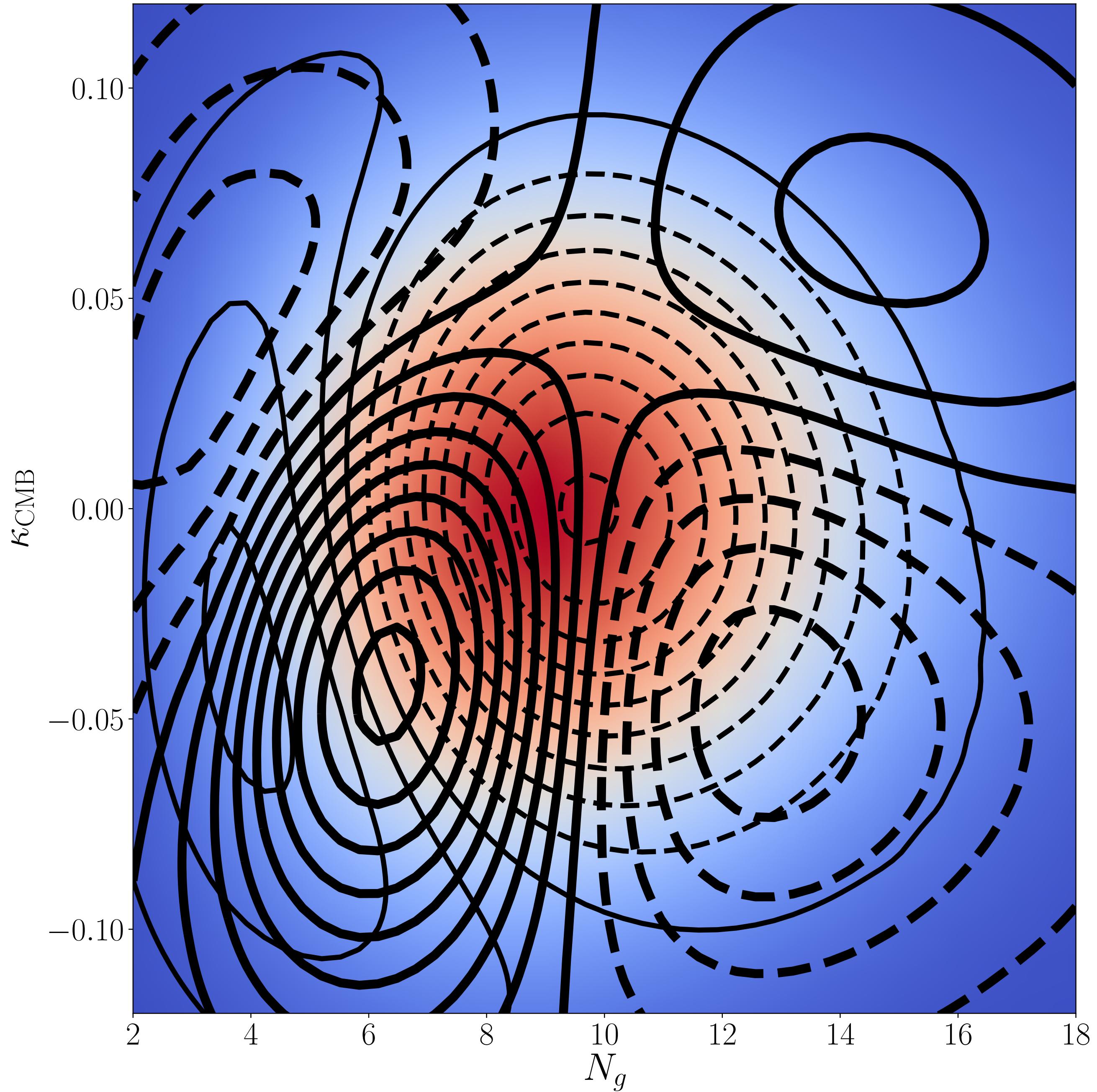
Blue: model predictions
 calculated in **Friedrich++(2018)**
Red: data from Buzzard N-body
 sims (DeRose, Wechsler++2019)

$$\delta_m = (\rho_m - \bar{\rho}_m)/\bar{\rho}_m$$

$$\delta_g = (\rho_g - \bar{\rho}_g)/\bar{\rho}_g$$

The full shape of the density PDF: A new dimension for studies of the large-scale structure

- Why the PDF?
- **PDF vs. $P(k)$ — the details**
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PDF / cumulants

- Cumulant generating function (CGF):

$$\varphi_R(\lambda) = \sum_n \langle \delta_R^n \rangle_c \frac{\lambda^n}{n!}$$

δ_R = density contrast smoothed by radius R

$\langle \delta_R^n \rangle_c$ = Cumulant (part of the nth moment that vanishes for Gaussian field)

2-point statistics

- Power spectrum:

$$\langle \delta_{\mathbf{k}} \delta_{\mathbf{q}} \rangle \sim P(k) \delta_D(\mathbf{k} + \mathbf{q})$$

PDF / cumulants

- Actual observable:

$$\text{PDF } p(\delta_R)$$

related to CGF via Laplace transform

$$e^{\varphi_R(\lambda)} = \int d\delta_R \ p(\delta_R) \ e^{\lambda\delta_R}$$

2-point statistics

- Alternative observable:

$$\text{2-point function } \xi(r) \equiv \langle \delta(\mathbf{x})\delta(\mathbf{x} + \mathbf{r}) \rangle$$

related to power spectrum via Fourier transform:

$$P(|\mathbf{k}|) \sim \int d^3r \ \xi(|\mathbf{r}|) \ e^{-i\mathbf{r}\mathbf{k}}$$

PDF / cumulants

- Leading order perturbation theory
Friedrich++ 2020, arxiv.org/abs/1912.06621 :

$\varphi_R(\lambda) \approx$ minimum of function

$$s_\lambda(\delta_i, j) = -\lambda F(\delta_i) + j\delta_i + \varphi_{R(1+F(\delta_i))^{1/3}, init}(j)$$

Spherical or
cylindrical collapse of
an initial fluctuation

“step 1: Legendre transform initial CGF”
“step 2: Legendre trans. back, but wrt transformed variable

See also:

Valageas (2002), Bernardeau et al. (2015)

Uhlemann et al. (2017), Friedrich et al. (2018)

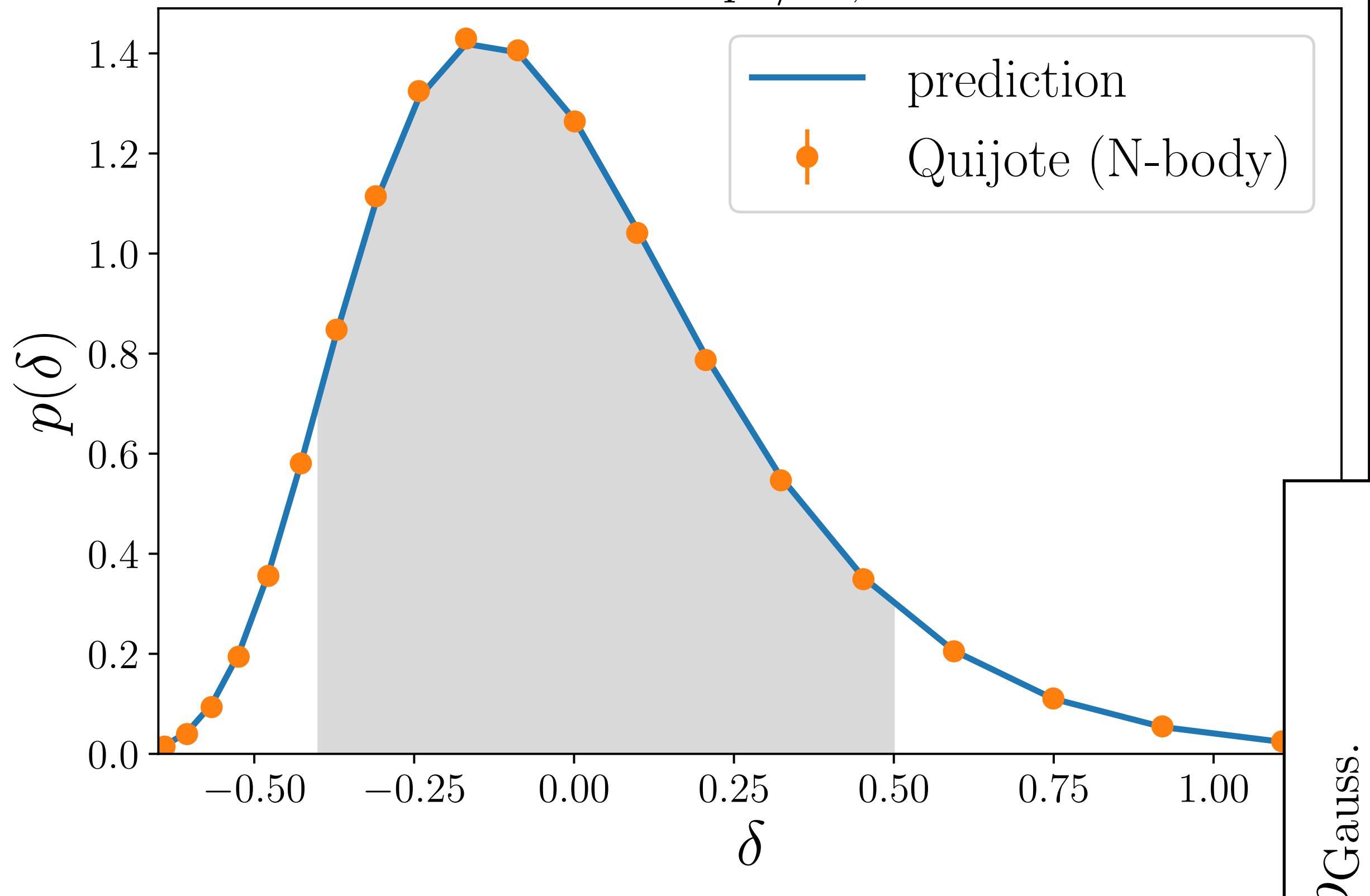
Ivanov et al. (2019) ... and many more!

2-point statistics

- Leading order perturbation theory

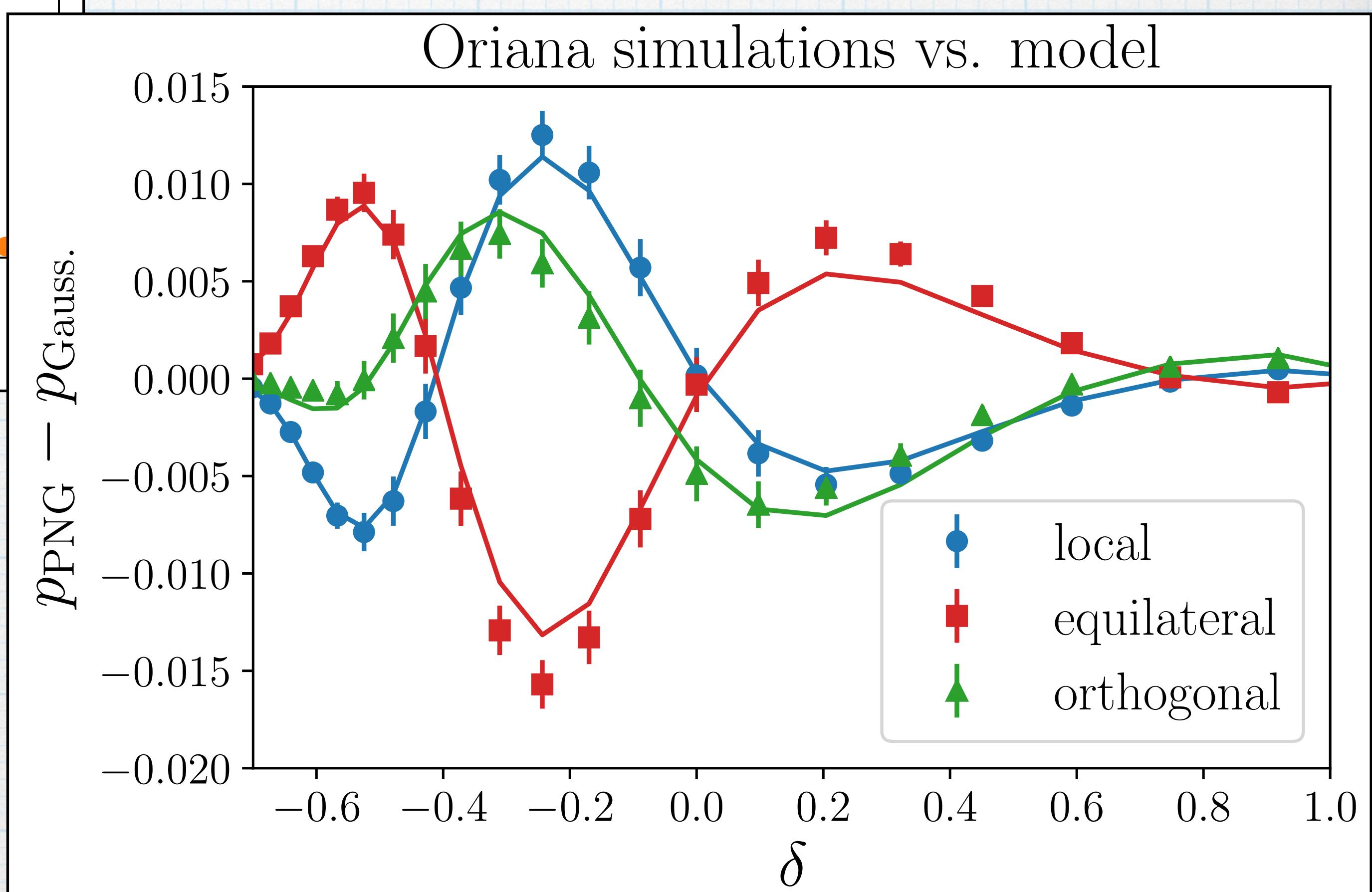
$$P(k, t) = P(k, t_i) \frac{D^2(t)}{D^2(t_i)}$$

$R = 15 \text{ Mpc}/h, z = 1$



Theory prediction
(Friedrich et al. 2020)
compared to the PDF measured in
the Quijote N-body simulations
(Villaescusa-Navarro et al. 2020)

Prediction for impact of primordial
non-Gaussianity in the PDF
(Friedrich et al. 2020)
compared to Oriana N-body
simulations **(Mao et al. 2014)**

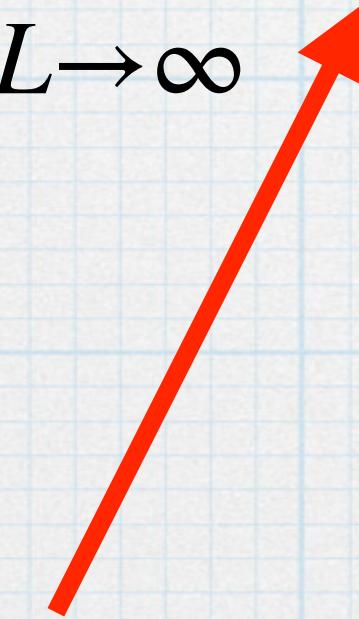


PDF / cumulants

- line-of-sight projections:

$$\varphi_{q,\theta}(\lambda) \approx$$

$$\int dw \lim_{L \rightarrow \infty} \frac{\varphi_{\text{cy},w\theta,L}(q(w)L\lambda, w)}{L}$$



**CGF of cylindrical
filter**

2-point statistics

- line-of-sight projections:

$$P_q(\ell) \approx$$

$$\int dw \left(\frac{q(w)}{w} \right)^2 P_{3D} \left(\frac{\ell + \frac{1}{2}}{w}, \eta_0 - w \right)$$

PDF / cumulants

- Joint PDF of galaxies and matter:

$$p(\delta_m, \delta_g) = p(\delta_m) p(\delta_g | \delta_m)$$

Sofar mostly:

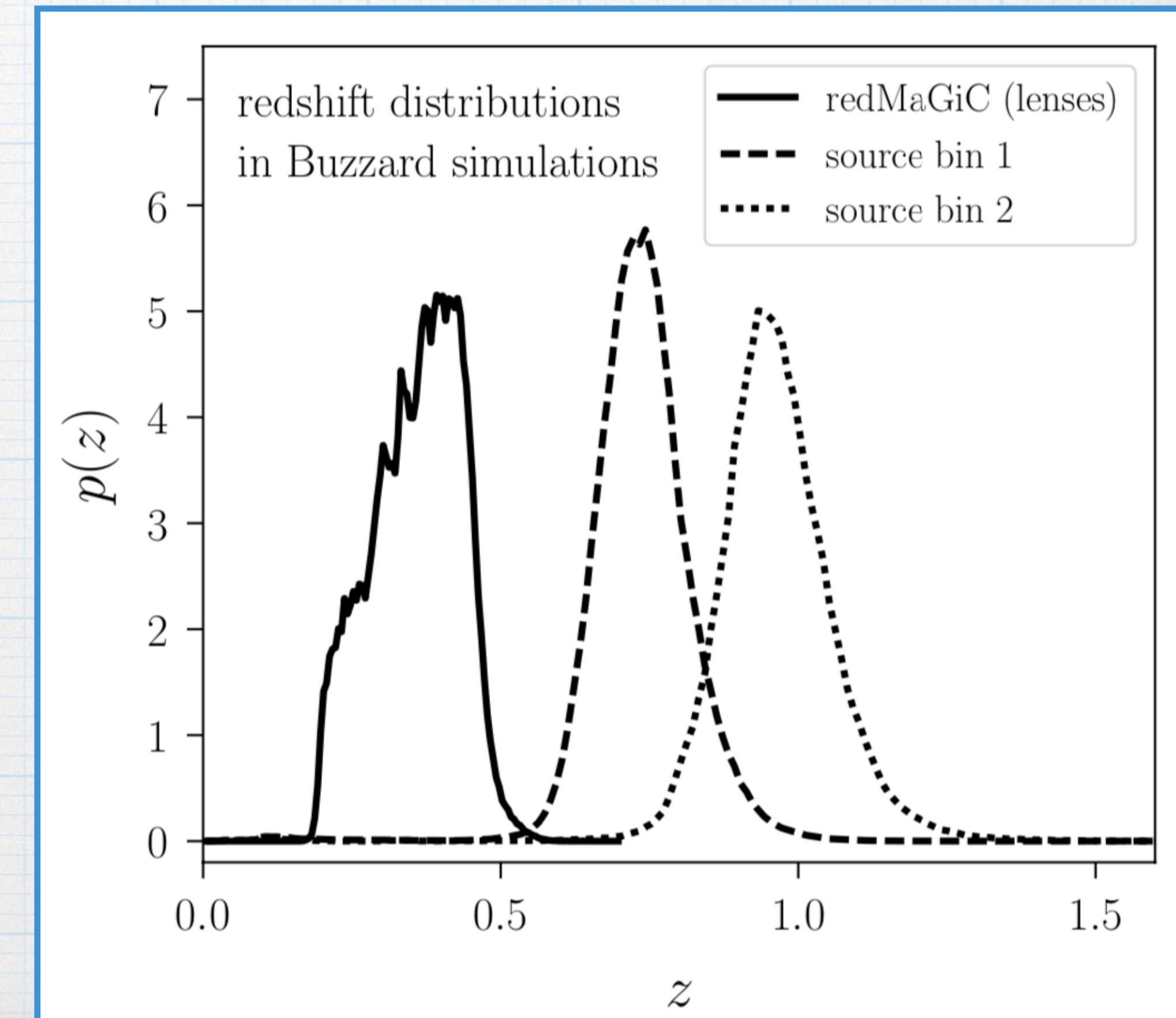
$$p(\delta_g | \delta_m) = \text{Poisson}(\langle \delta_g | \delta_m \rangle)$$

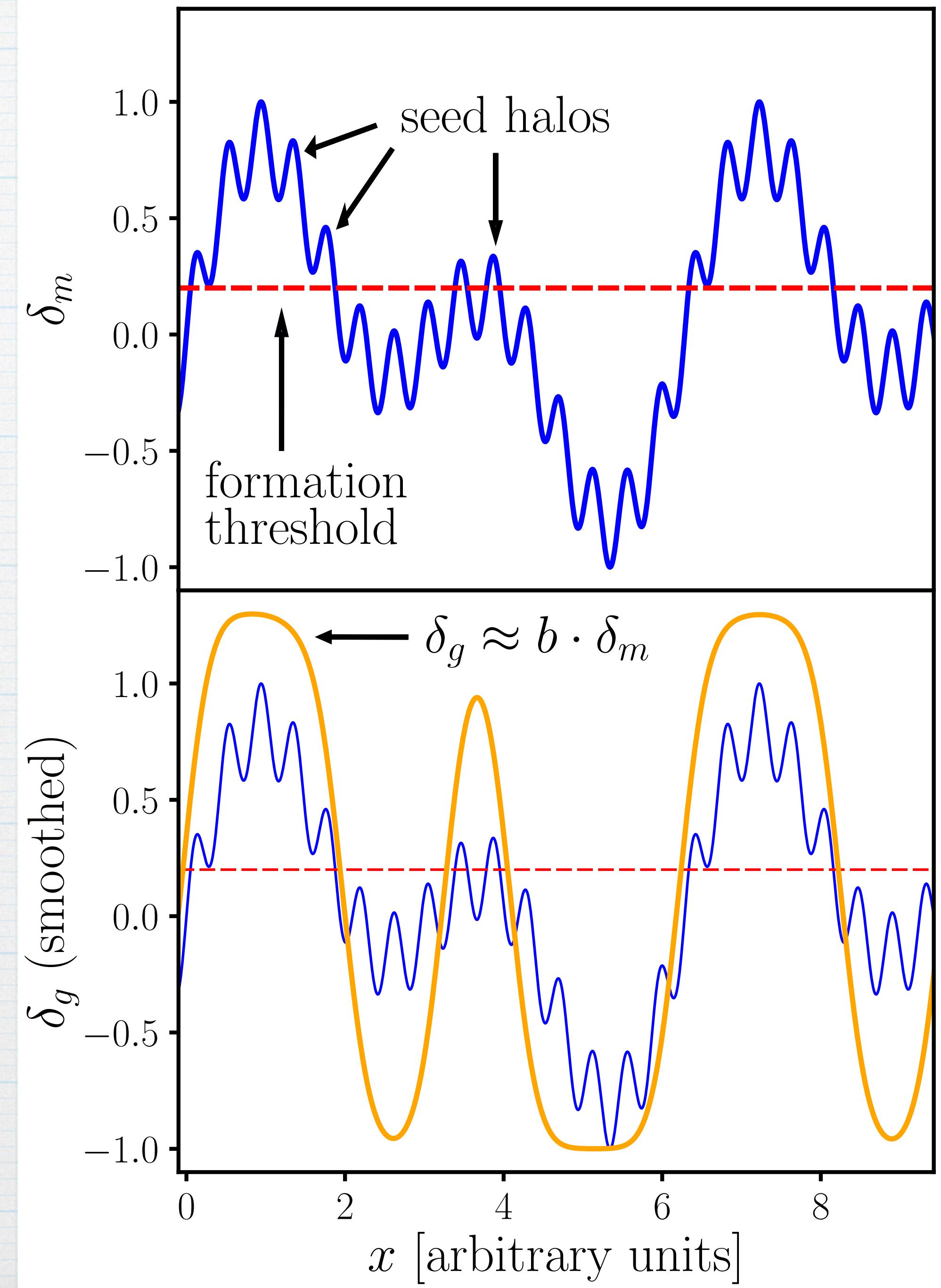
$$\langle \delta_g | \delta_m \rangle \approx b_1^E \delta_m + \frac{b_2^E}{2} (\delta_m^2 - \langle \delta_m^2 \rangle)$$

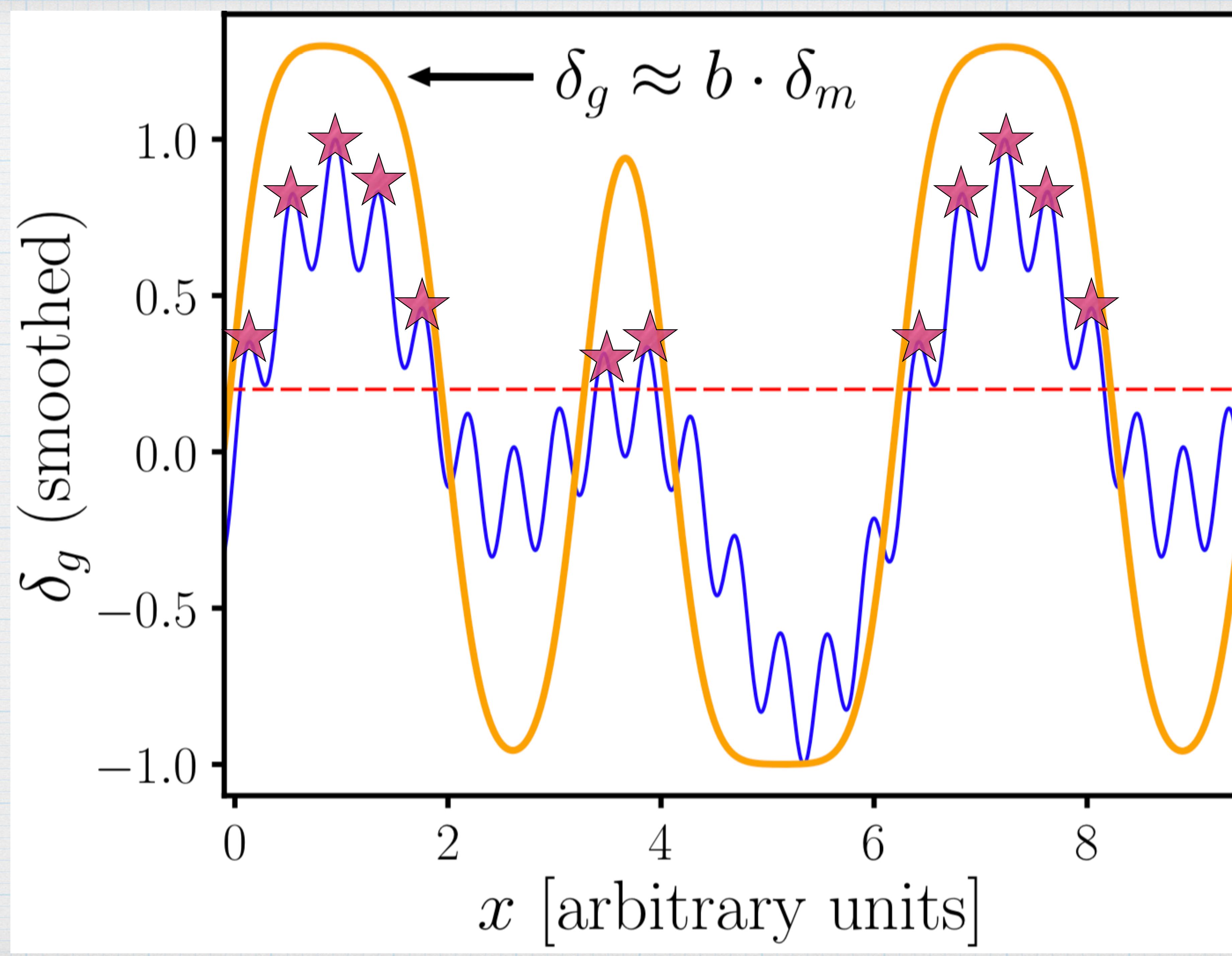
2-point statistics

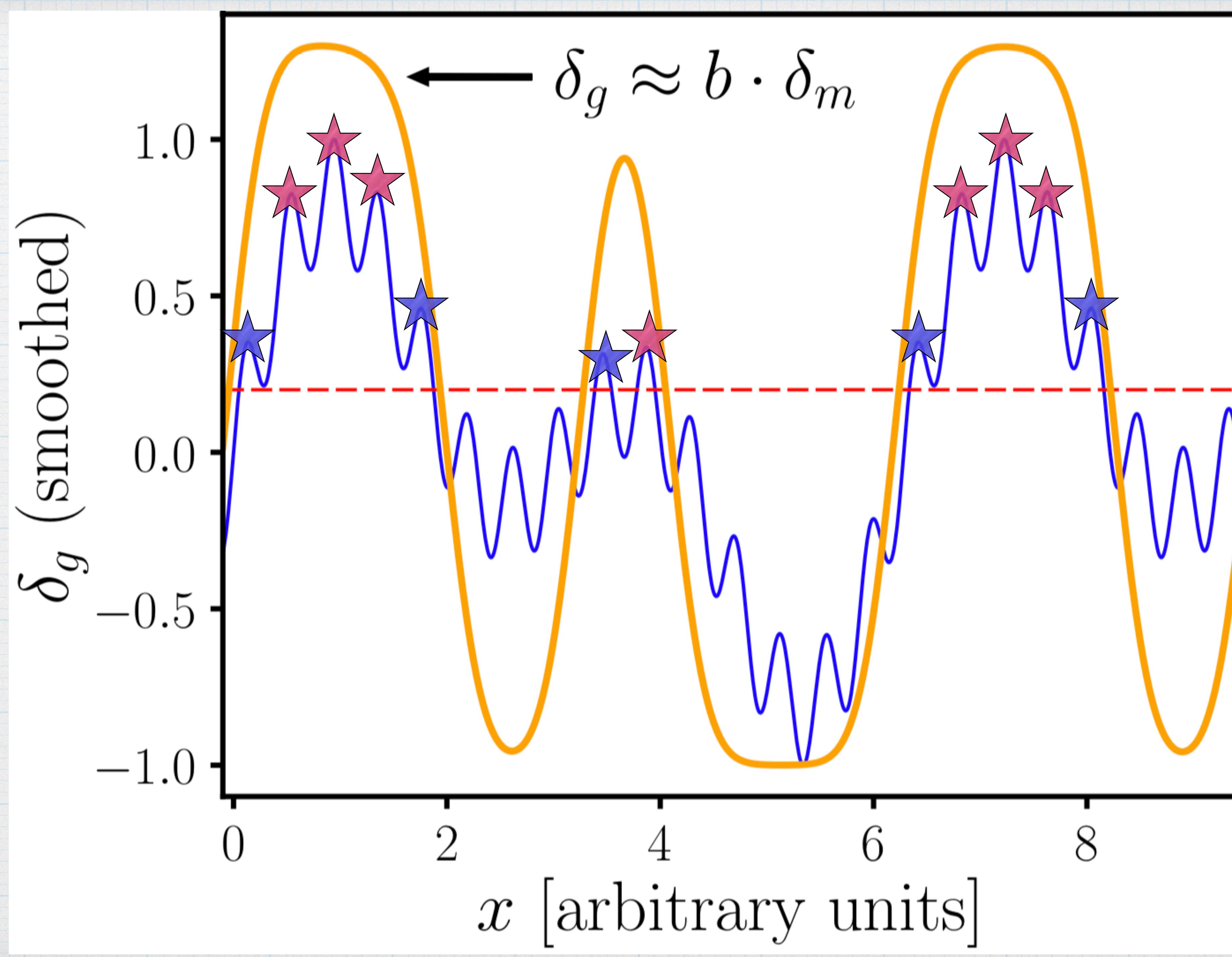
- Galaxy bias & shot-noise:

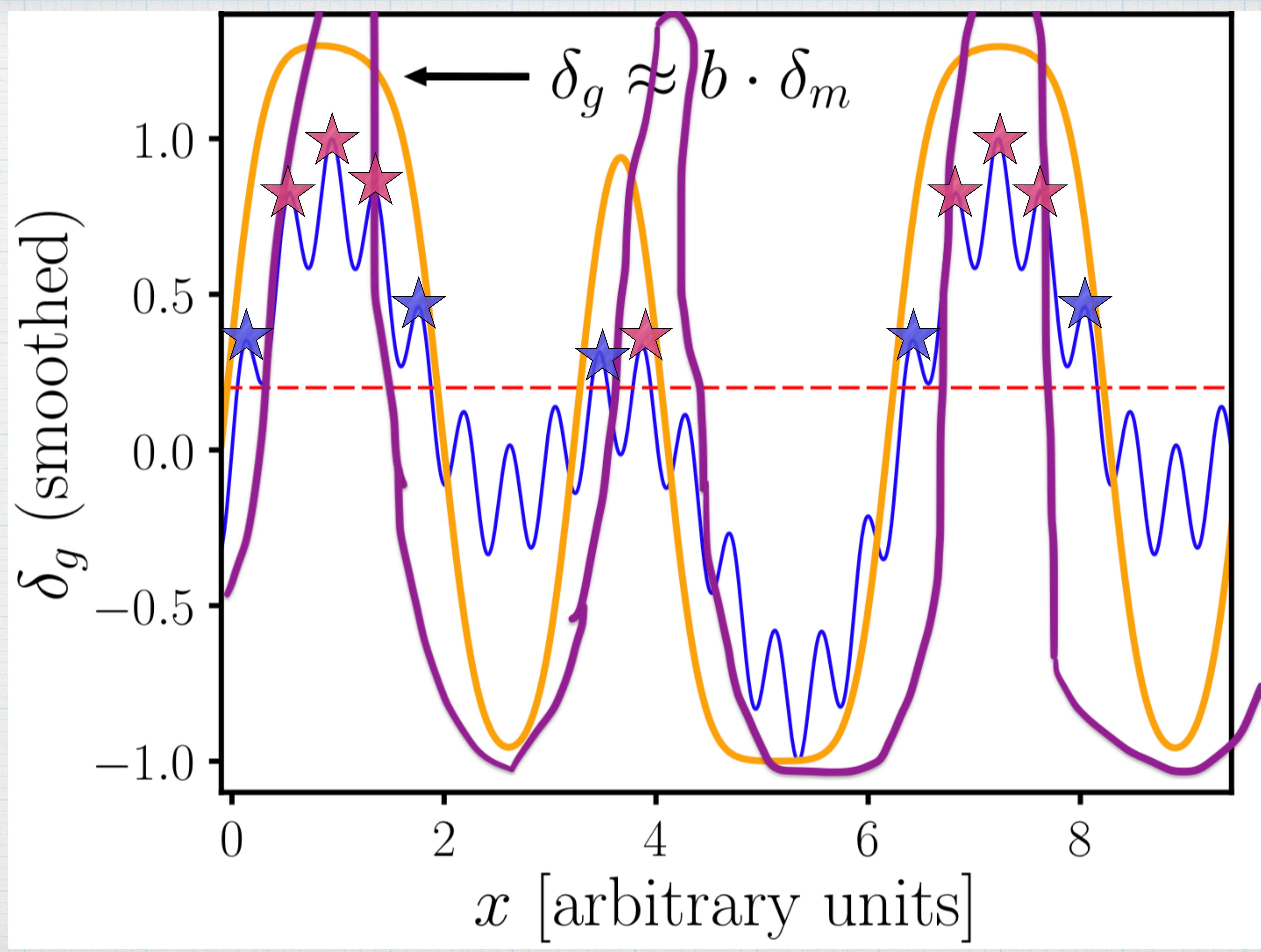
$$P_g(k) \approx b^2 P_m(k) + \frac{1}{n_g}$$











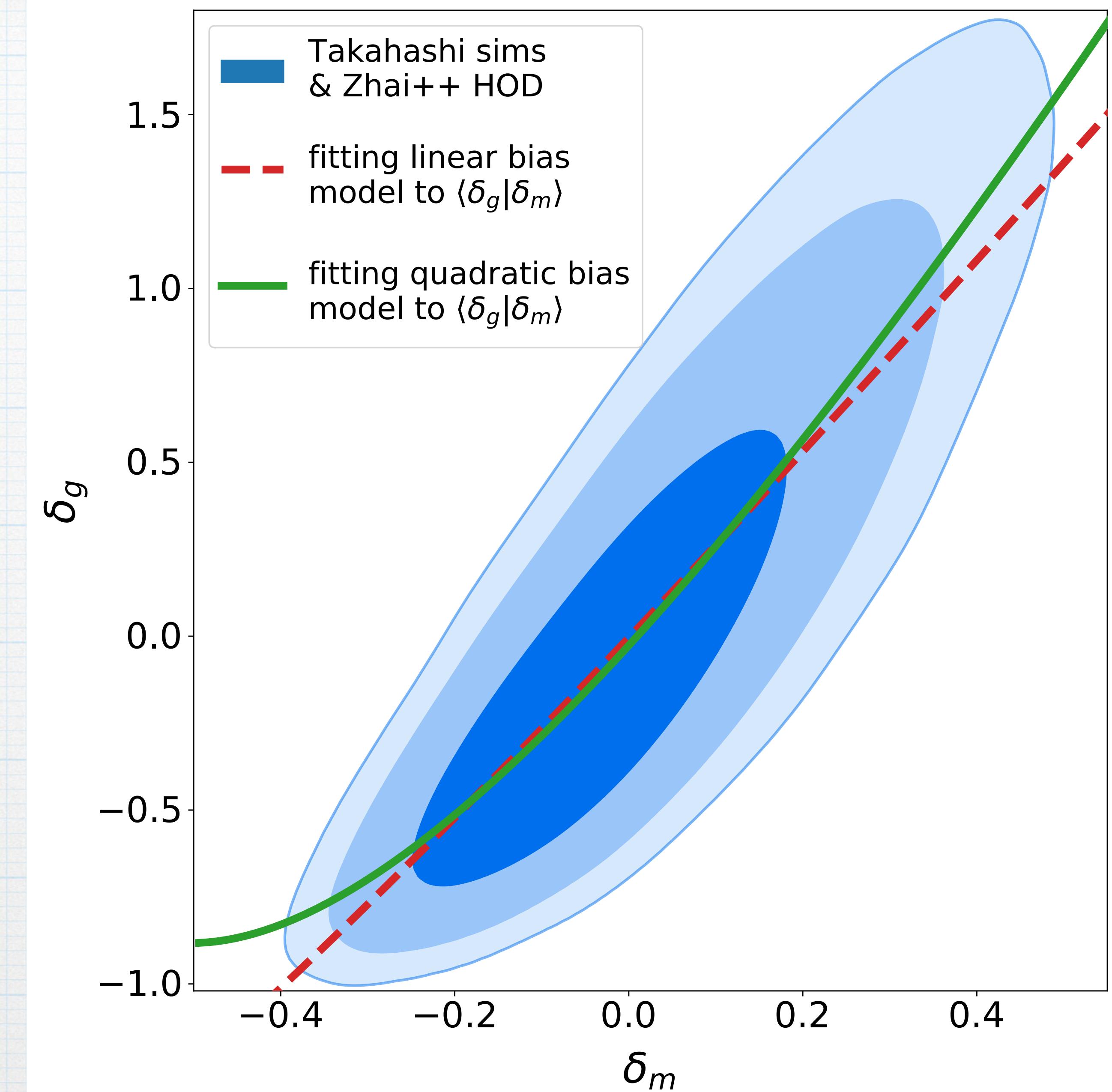
- Sofar mostly:

$$p(\delta_g | \delta_m) = \text{Poisson}(\langle \delta_g | \delta_m \rangle)$$

$$\langle \delta_g | \delta_m \rangle \approx b_1^E \delta_m + \frac{b_2^E}{2} (\delta_m^2 - \langle \delta_m^2 \rangle)$$

- We have implemented:

- Lagrangian bias model
(expansion in terms of initial δ)
- non-Poisson noise



- Joint PDF of galaxies and matter:

$$p(\delta_m, \delta_g) = p(\delta_m) p(\delta_g | \delta_m)$$

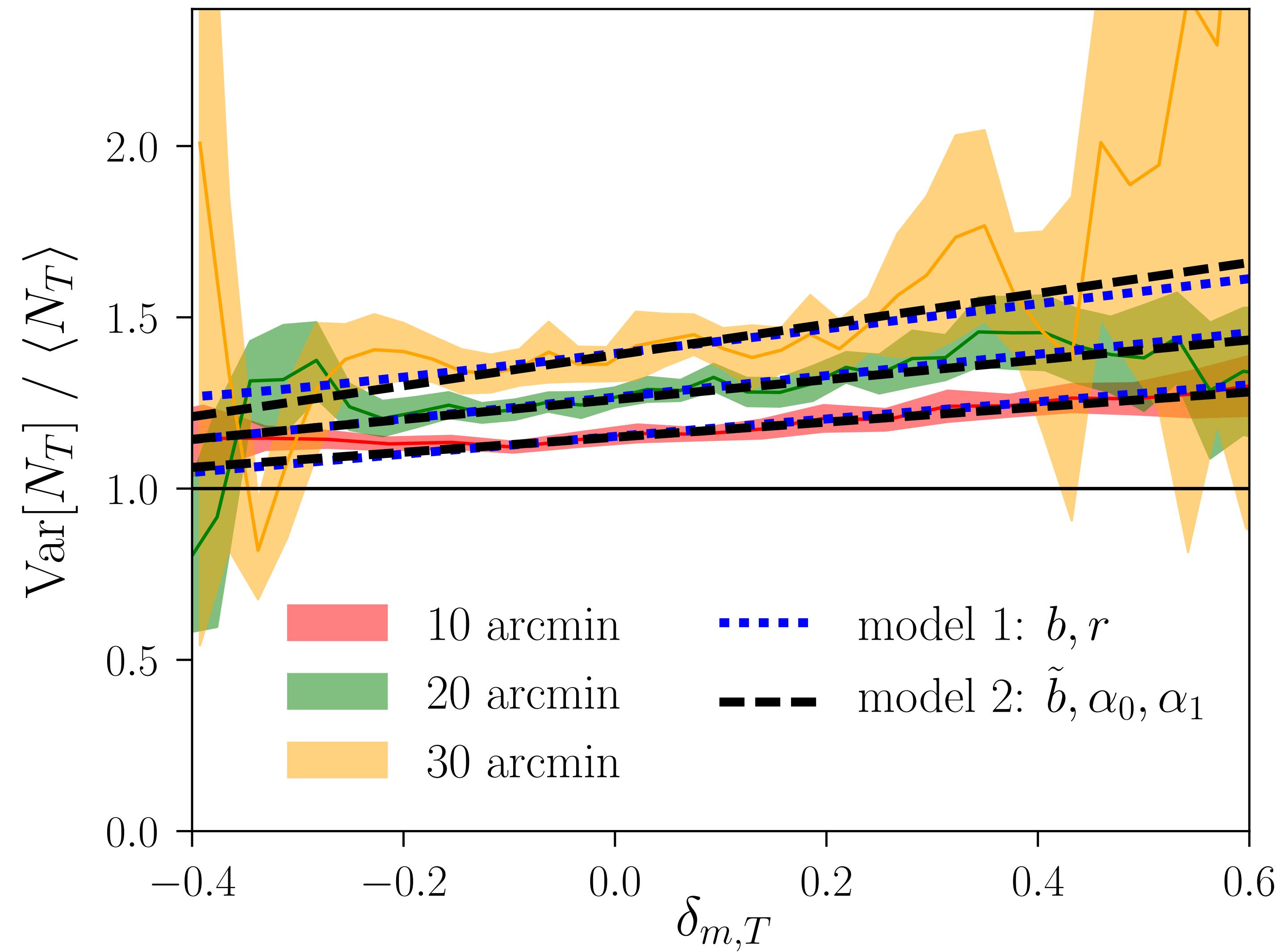
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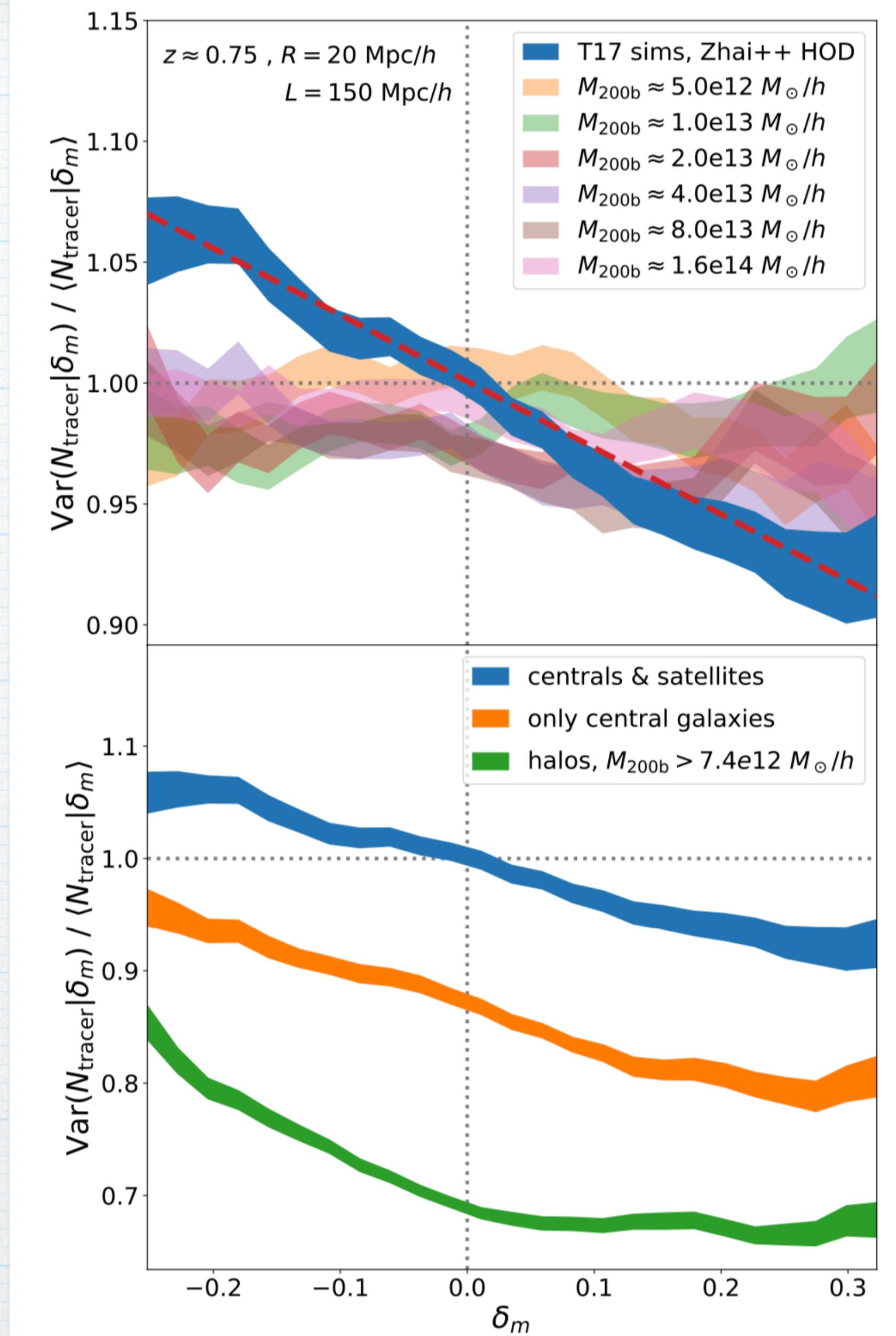
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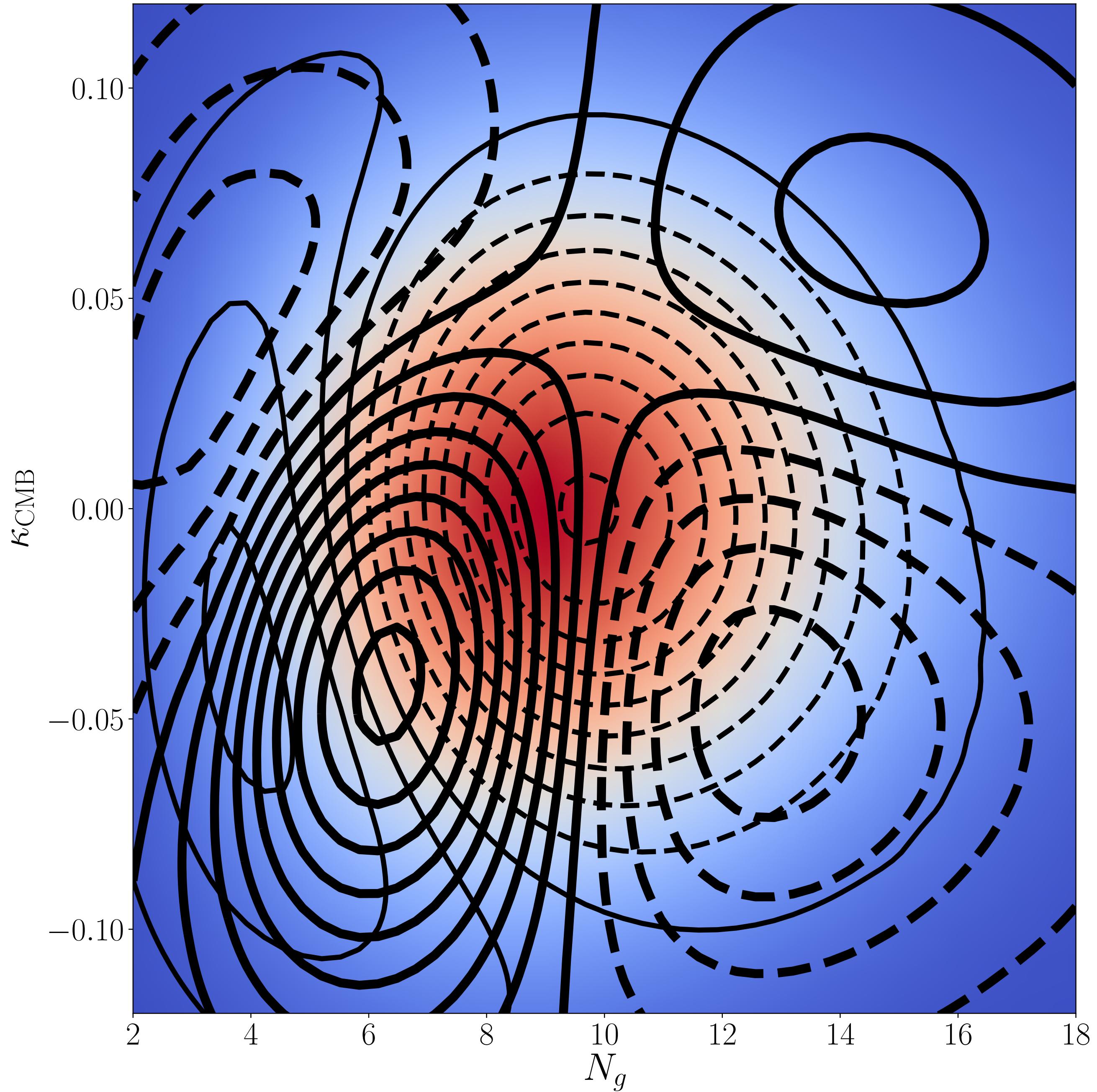
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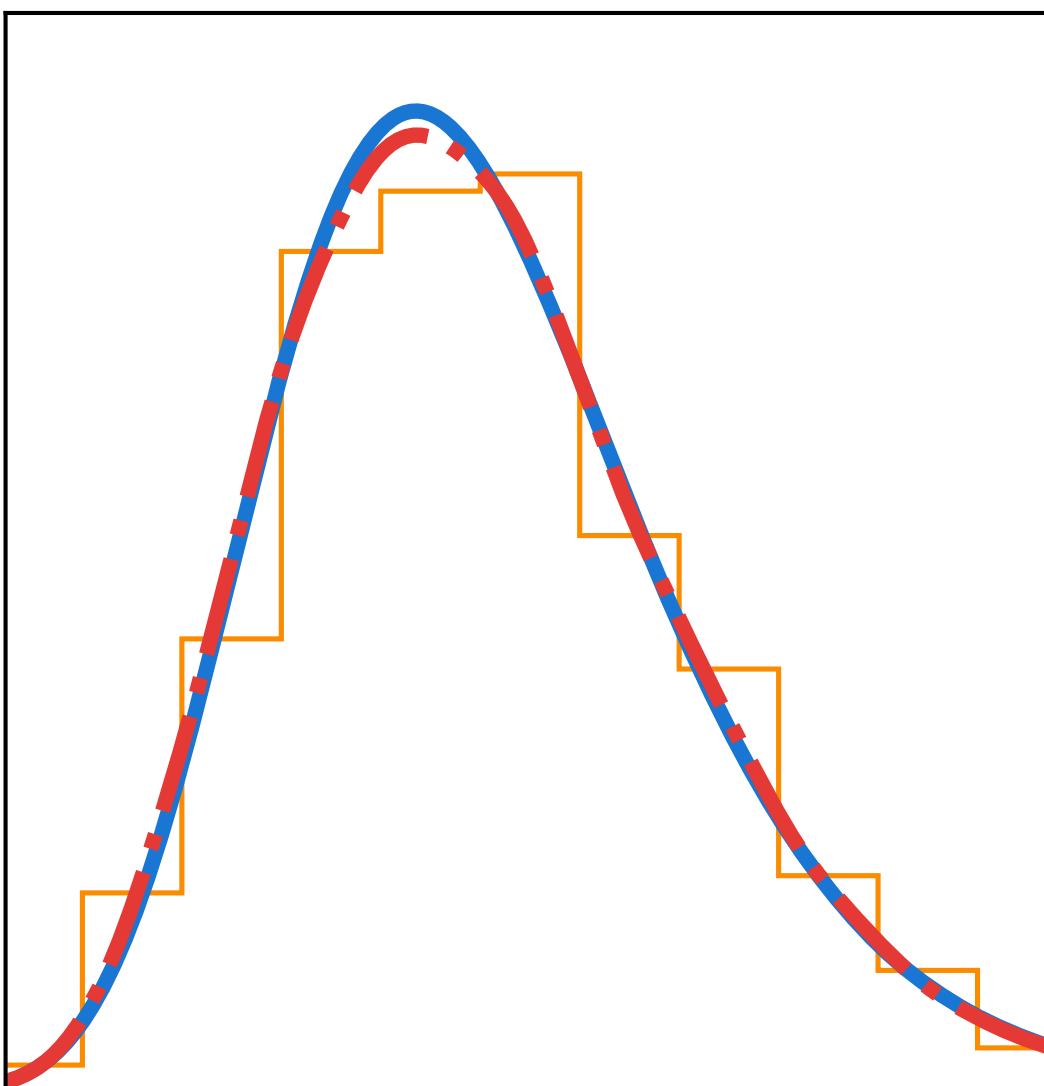
- Lagrangian bias model
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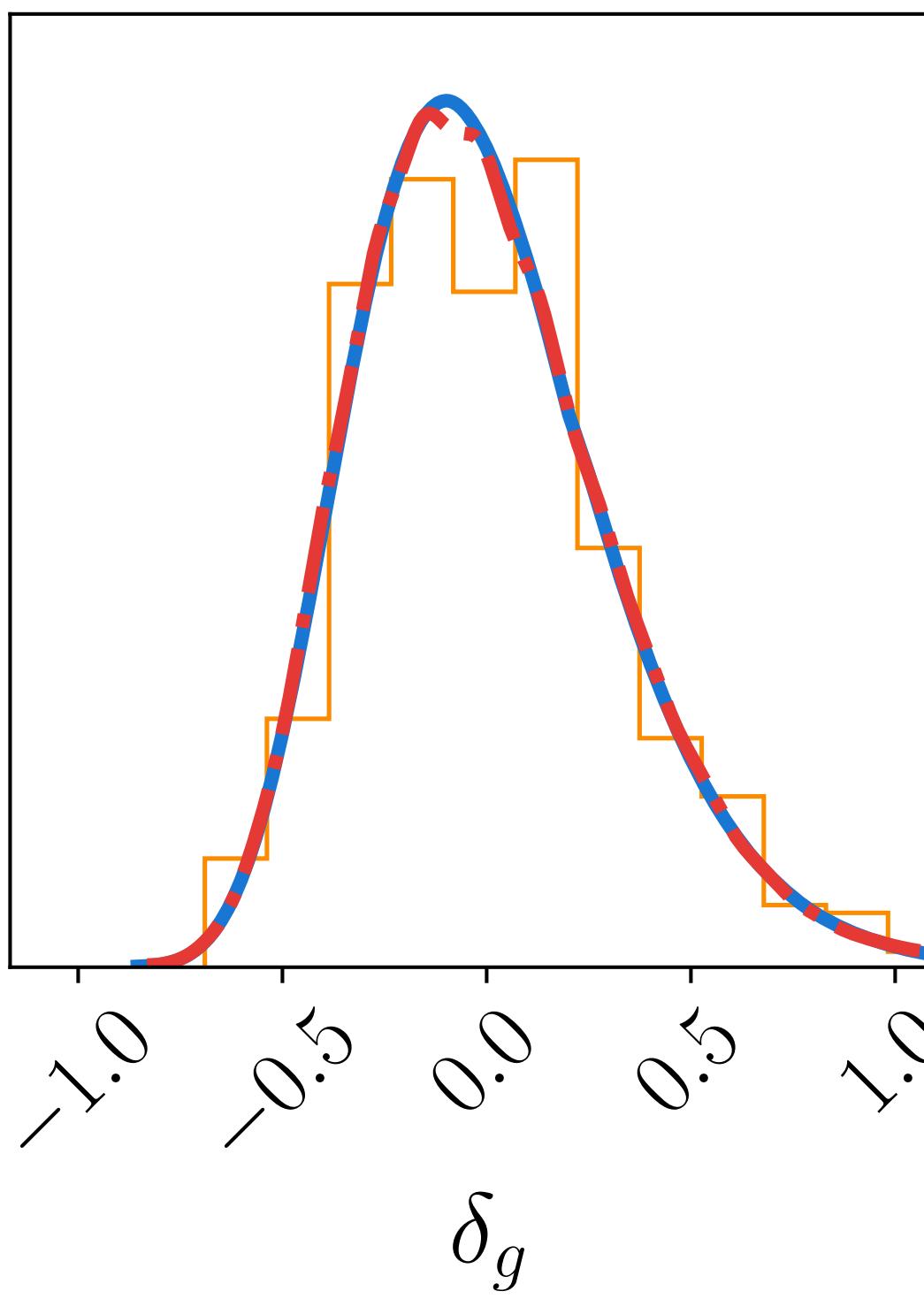
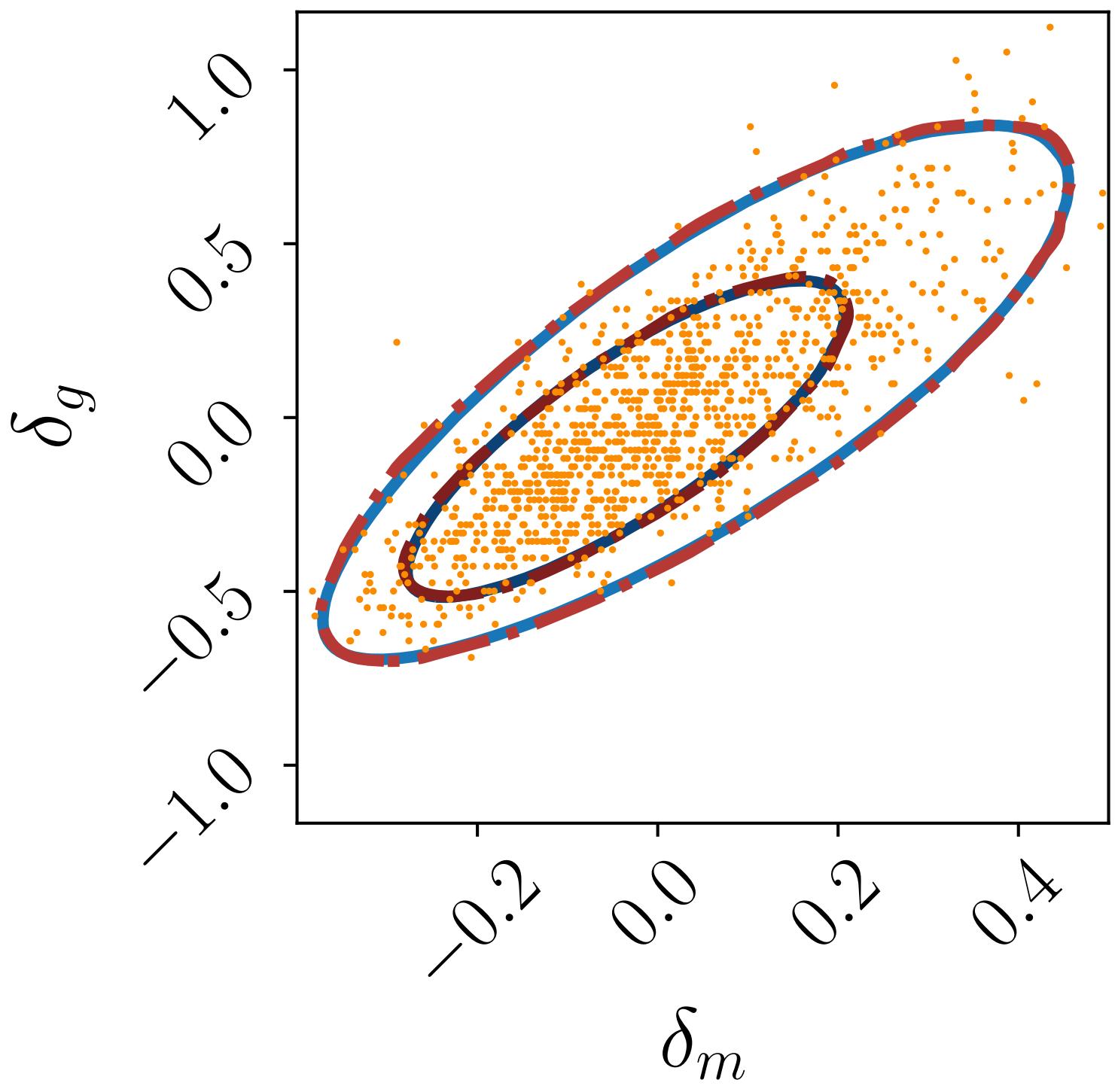
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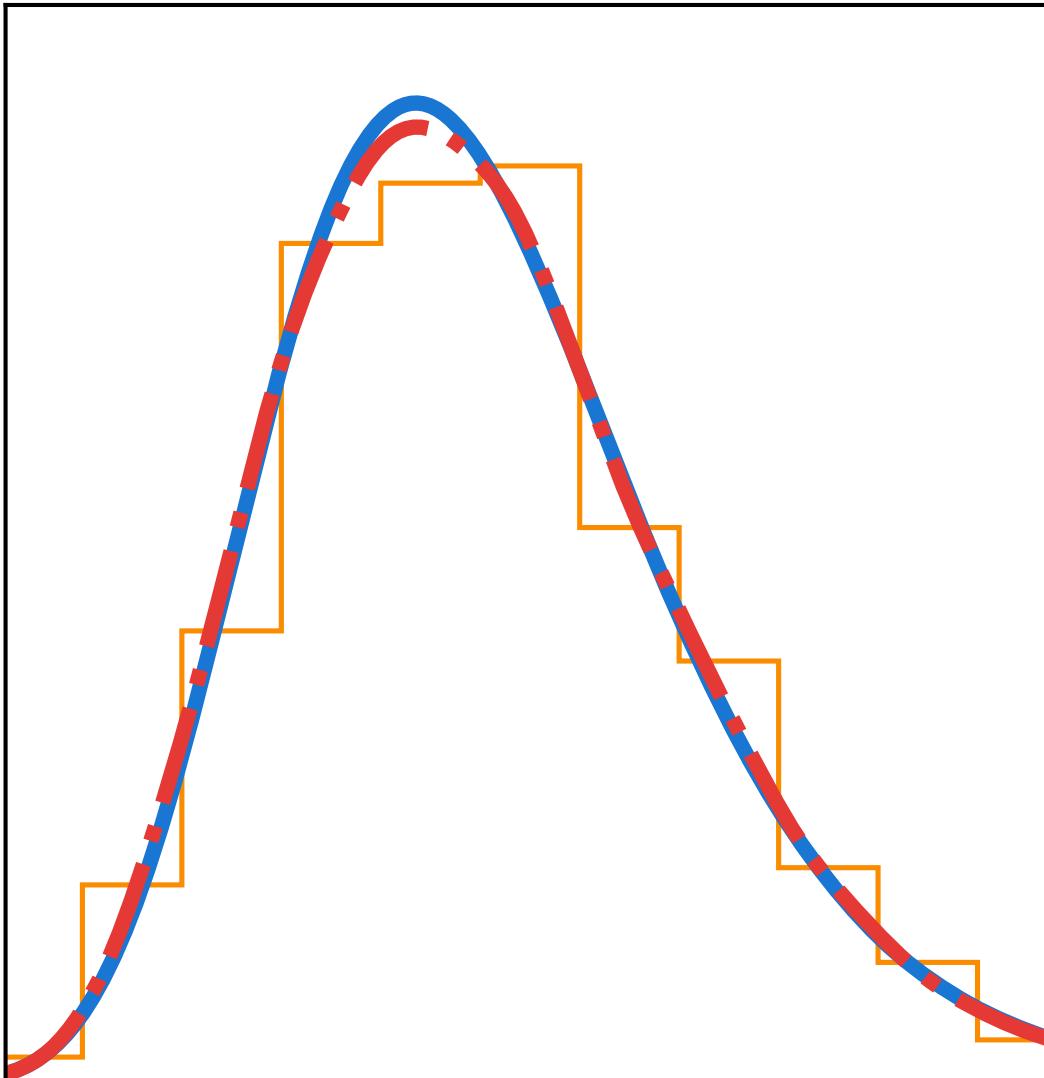
buzzard (contours)
buzzard (rand. subset)



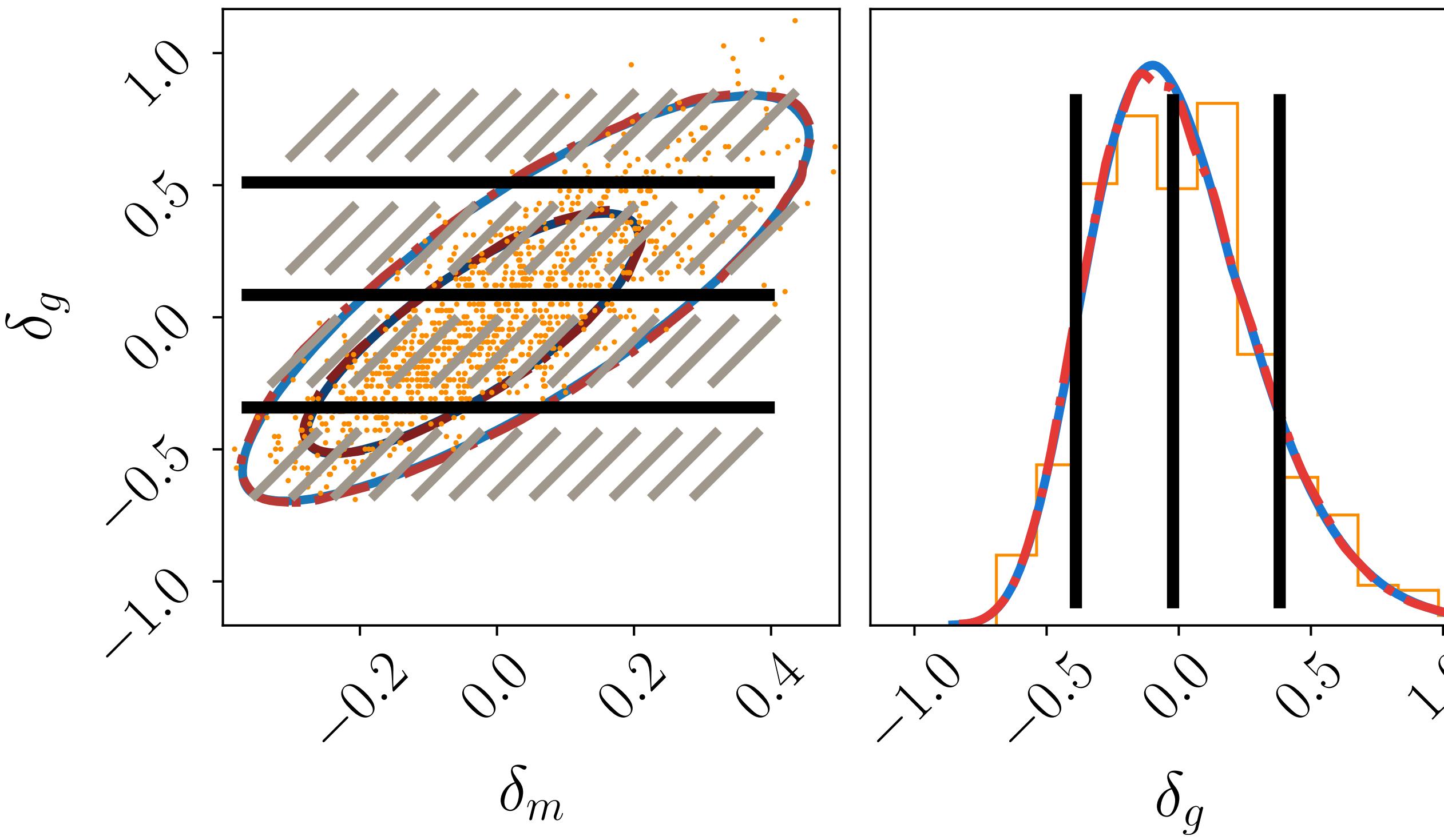
Blue - model predictions
calculated in **Friedrich++ (2018)**

Red - data from Buzzard N-body
sims (DeRose, Wechsler++2019)

But:
does this work in real data?



$p(\delta_m, \delta_g)$
 buzzard (contours)
 buzzard (rand. subset)



Blue - model predictions
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But:
does this work in real data?

**We already (partly) did this with
DES year-1 data!**

Density split statistics uses
gravitational lensing to scan
through galaxy density PDF
“quantile-by-quantile” to connect it
to underlying matter density PDF

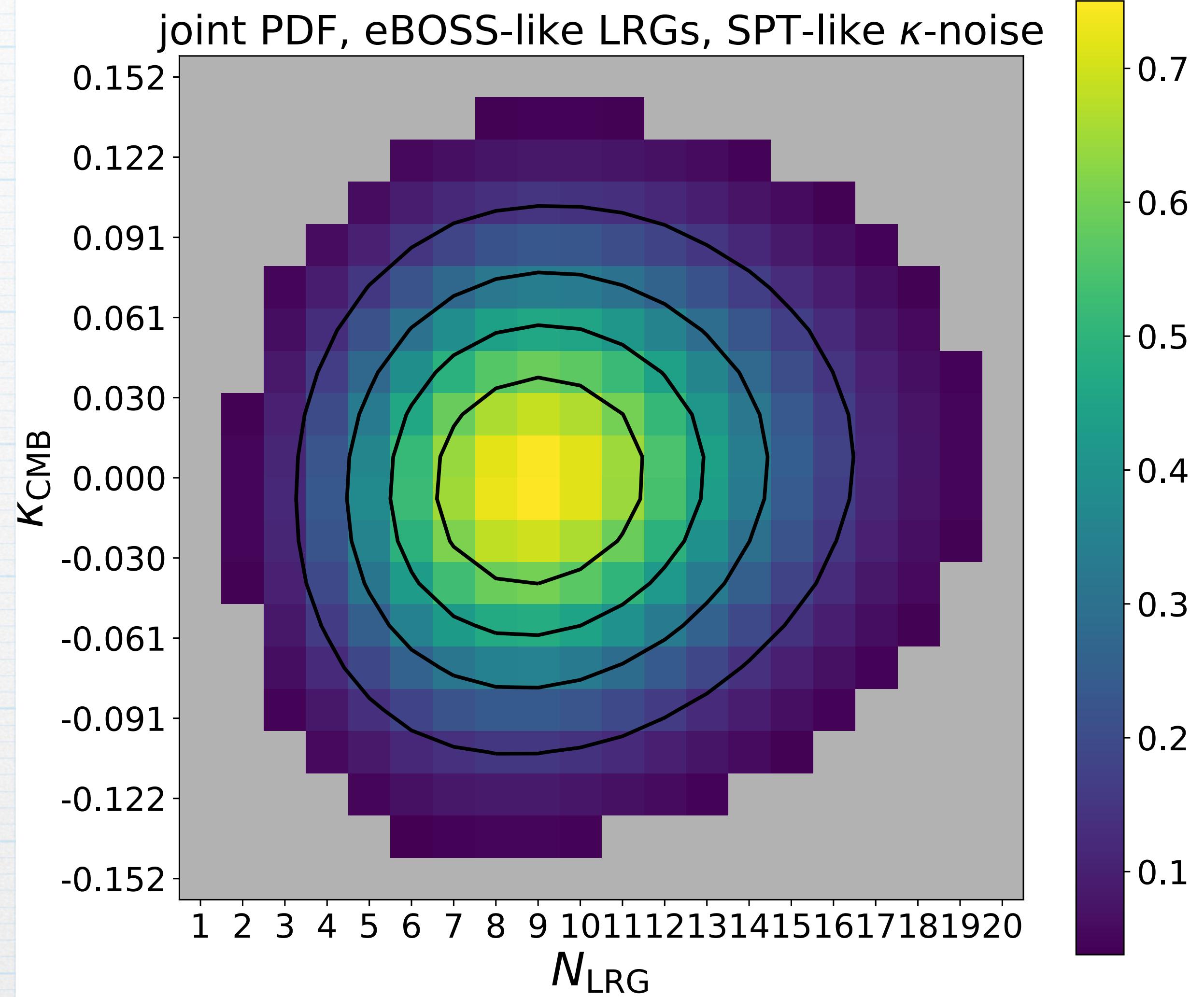
Friedrich et al. (2018)
Gruen et al. (2018)

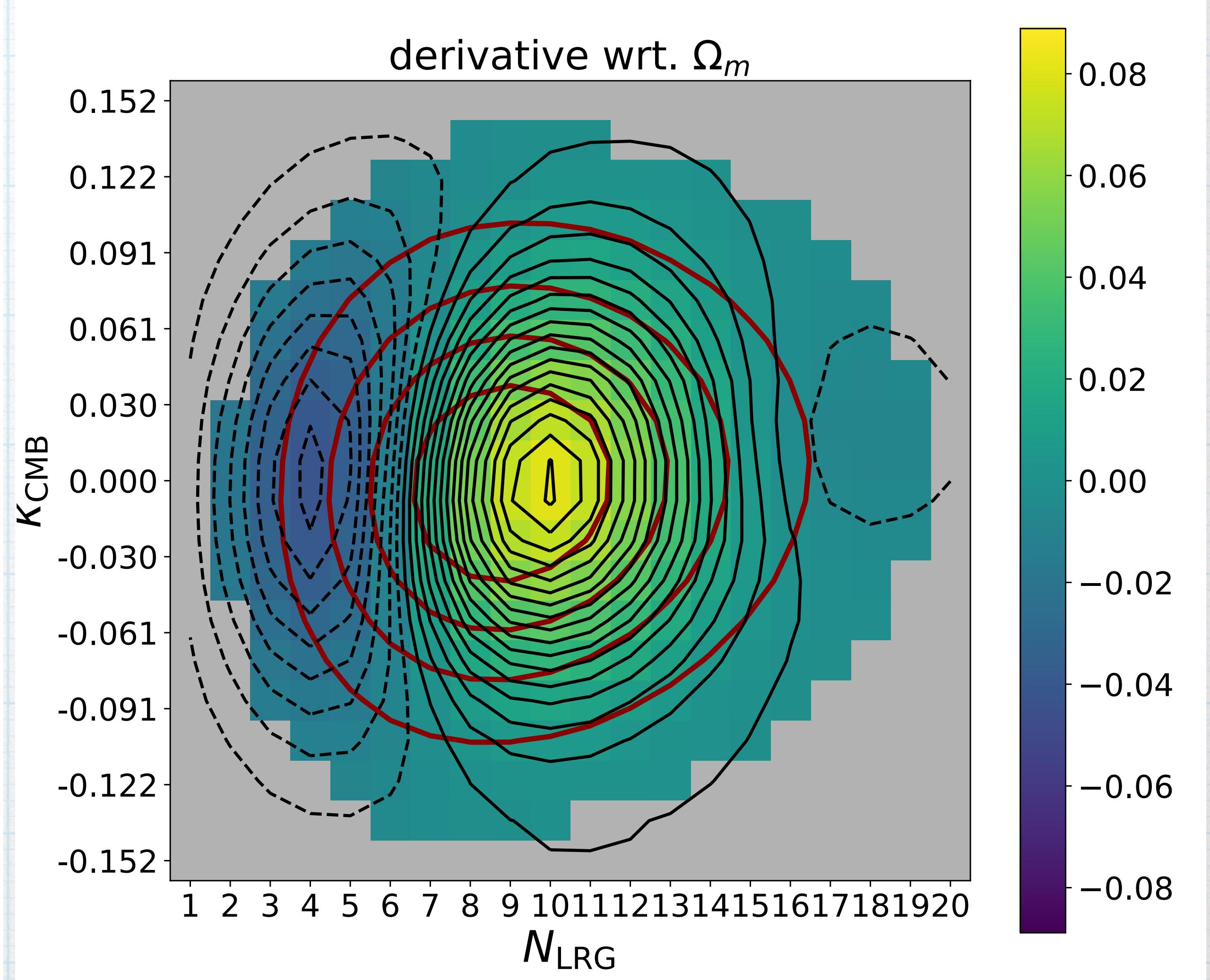
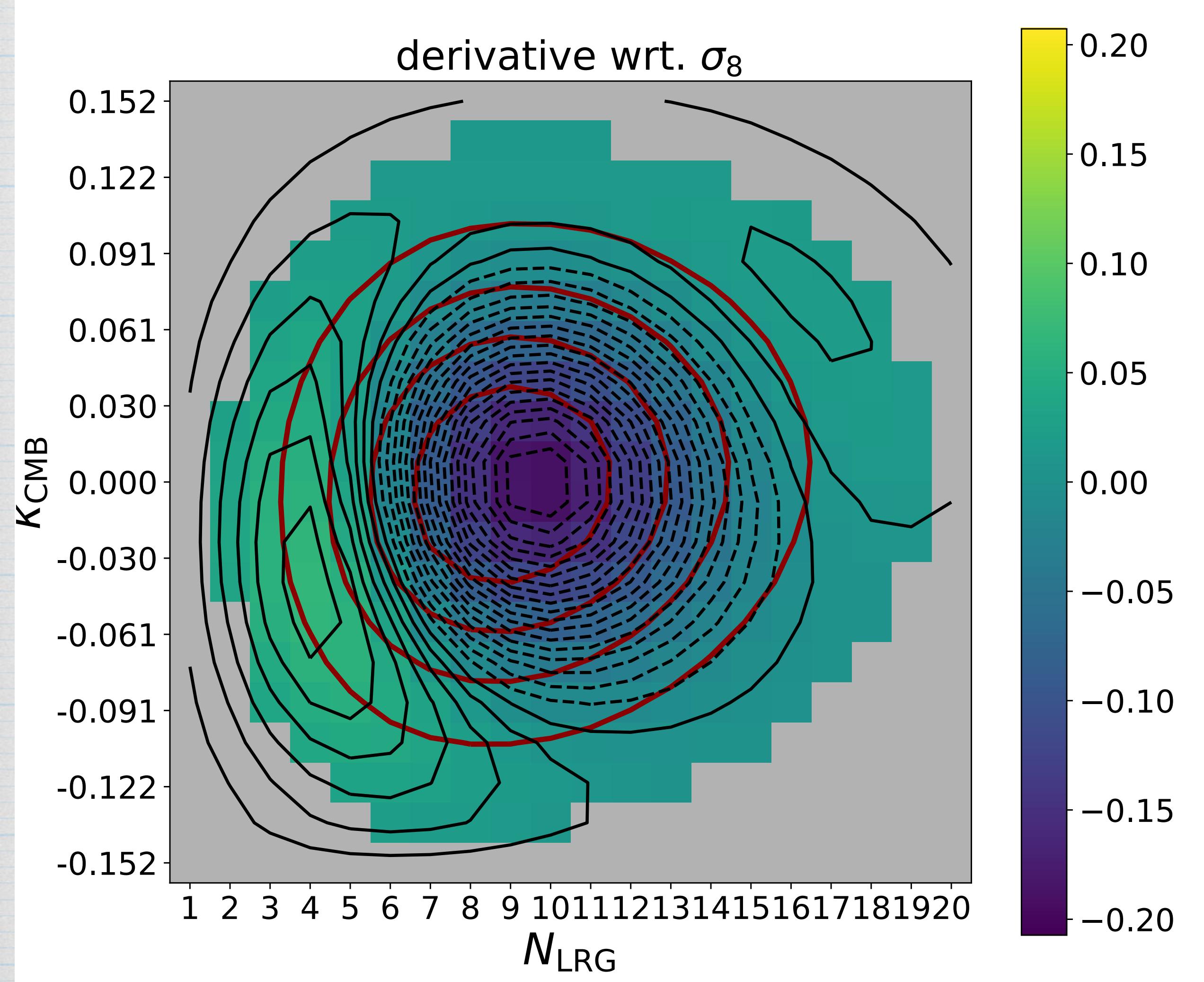
What do cosmological parameters do to the PDF?

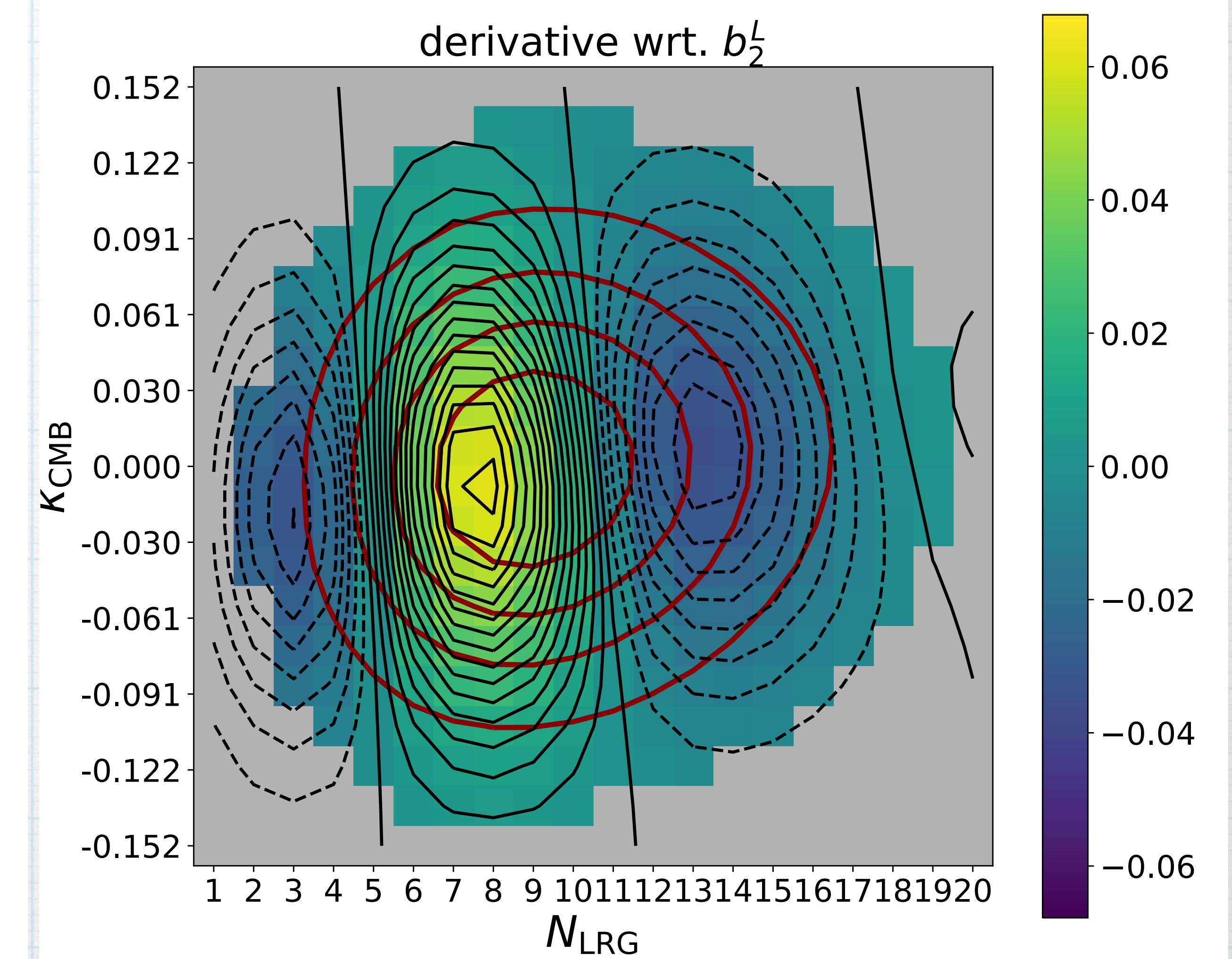
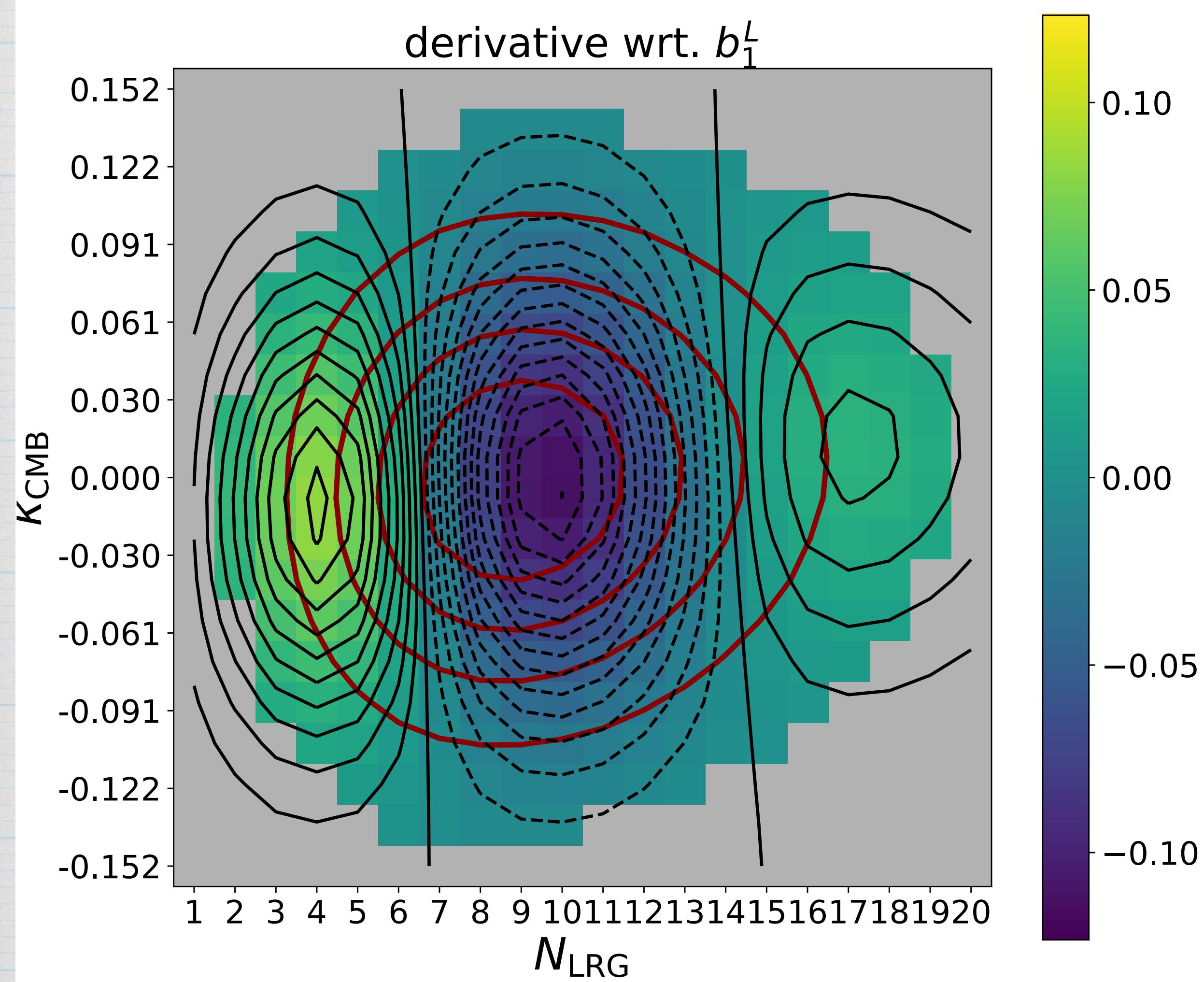
in the following assume:
(from Friedrich++ in prep)

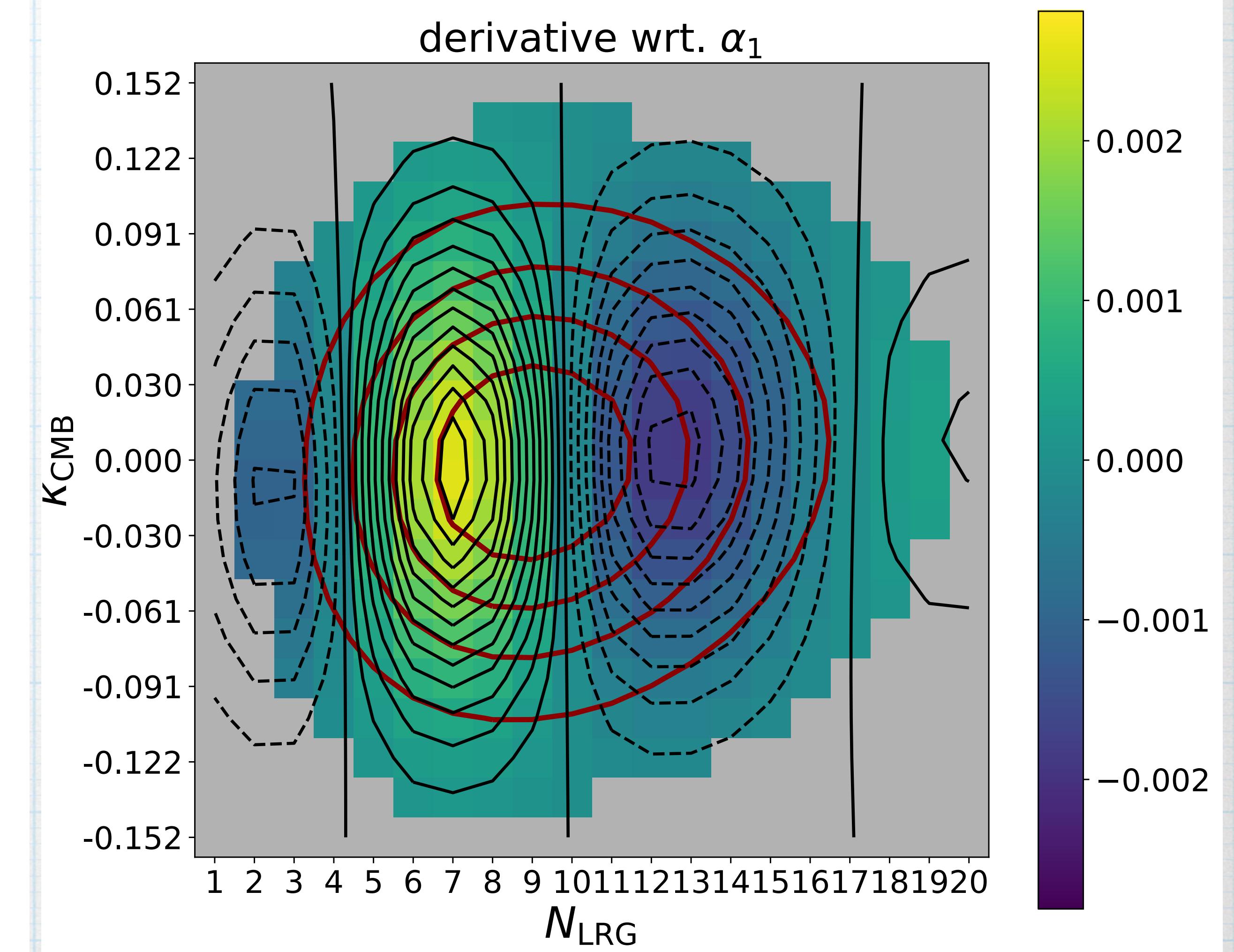
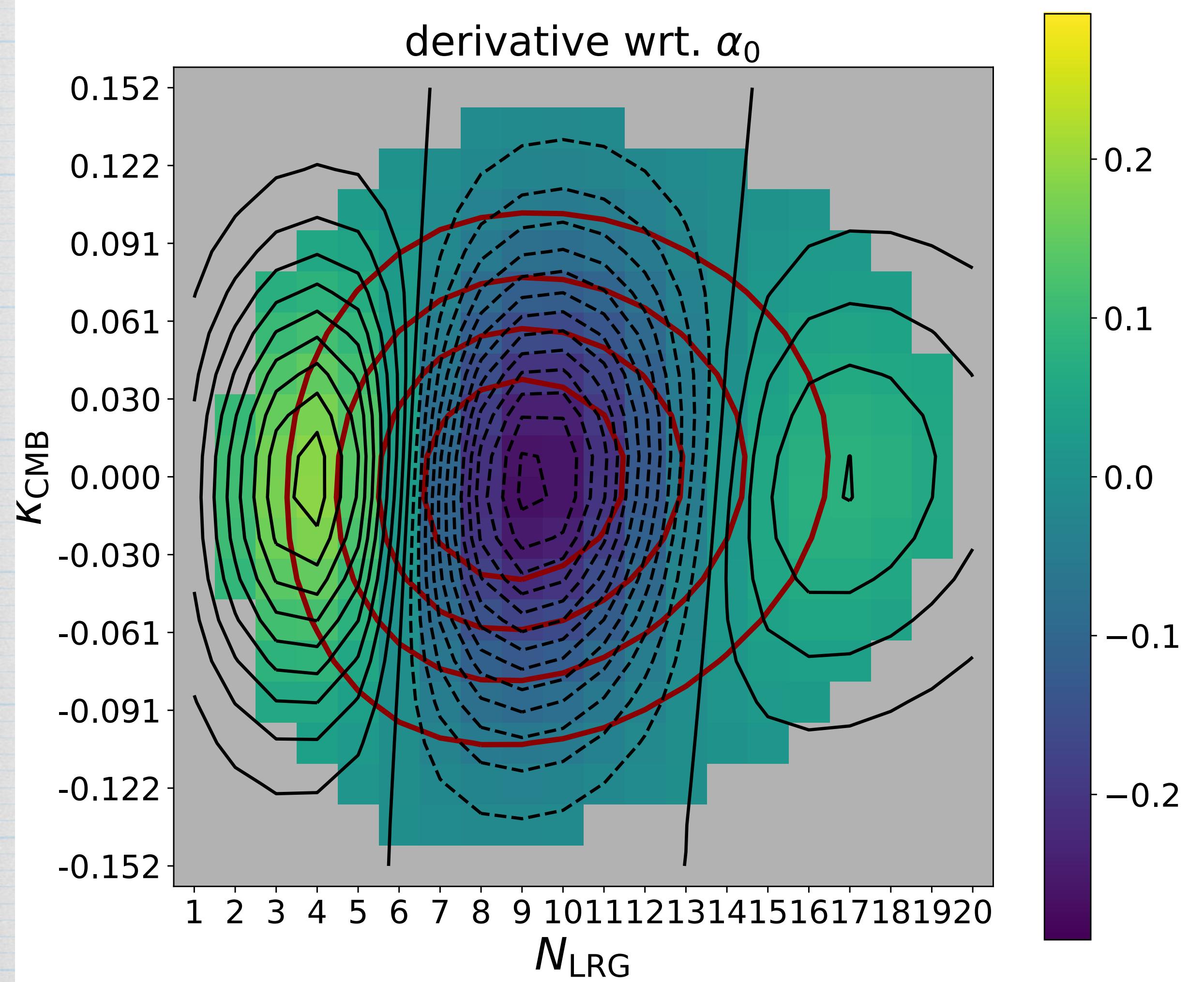
- LRGs (luminous red galaxies) within $0.6 < z < 0.9$
- CMB lensing with SPT-like noise
- 20 arcmin smoothing scale
- 5000 square degrees on the sky

→ measurable in real data!







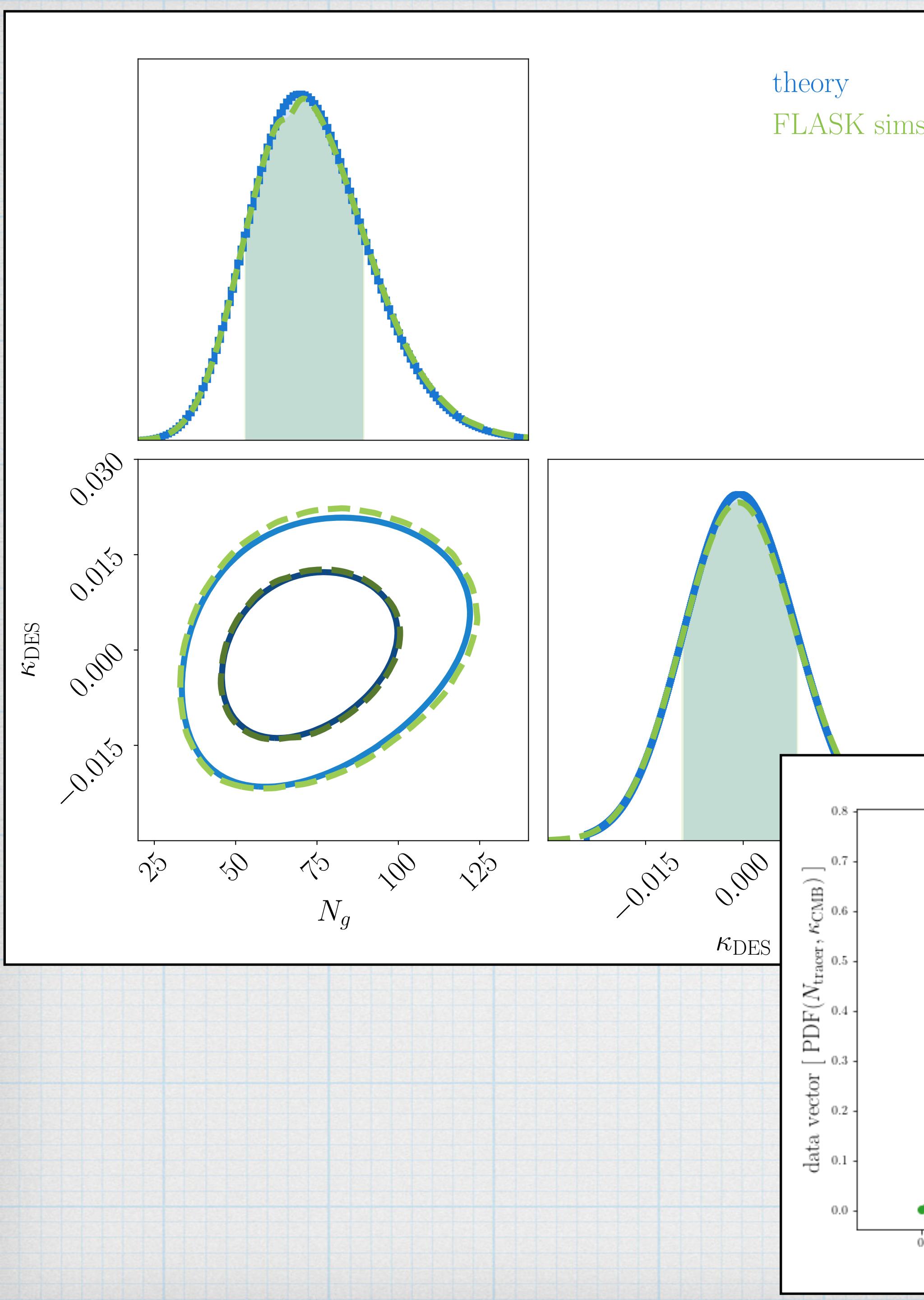


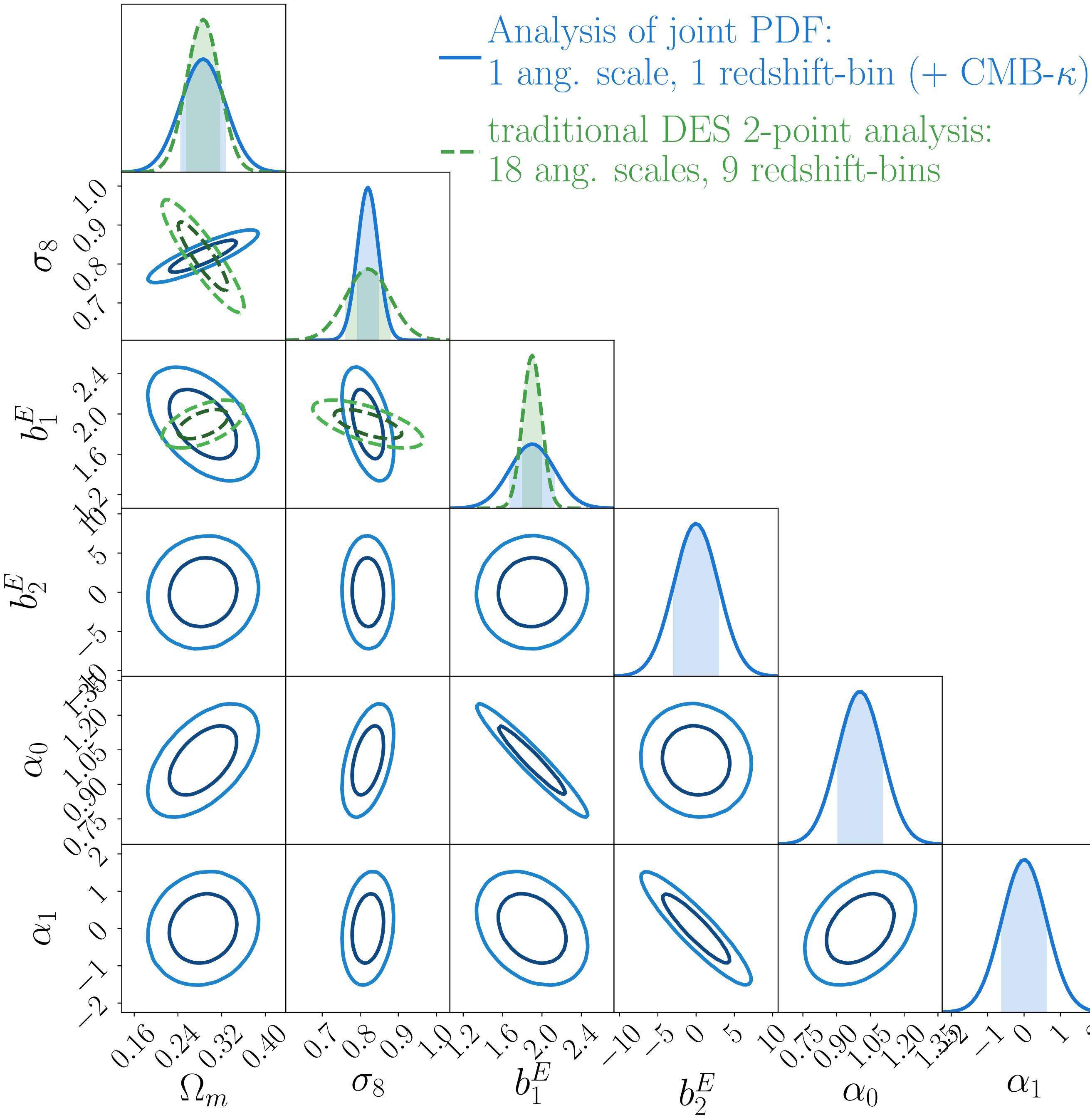
Covariance from log-normal Simulations

- have 700 FLASK simulations (Xavier et al. 2016)
- Combining several log-normal fields to capture non-Gaussian correlation structure



Work by Anik Halder,
LMU Munich



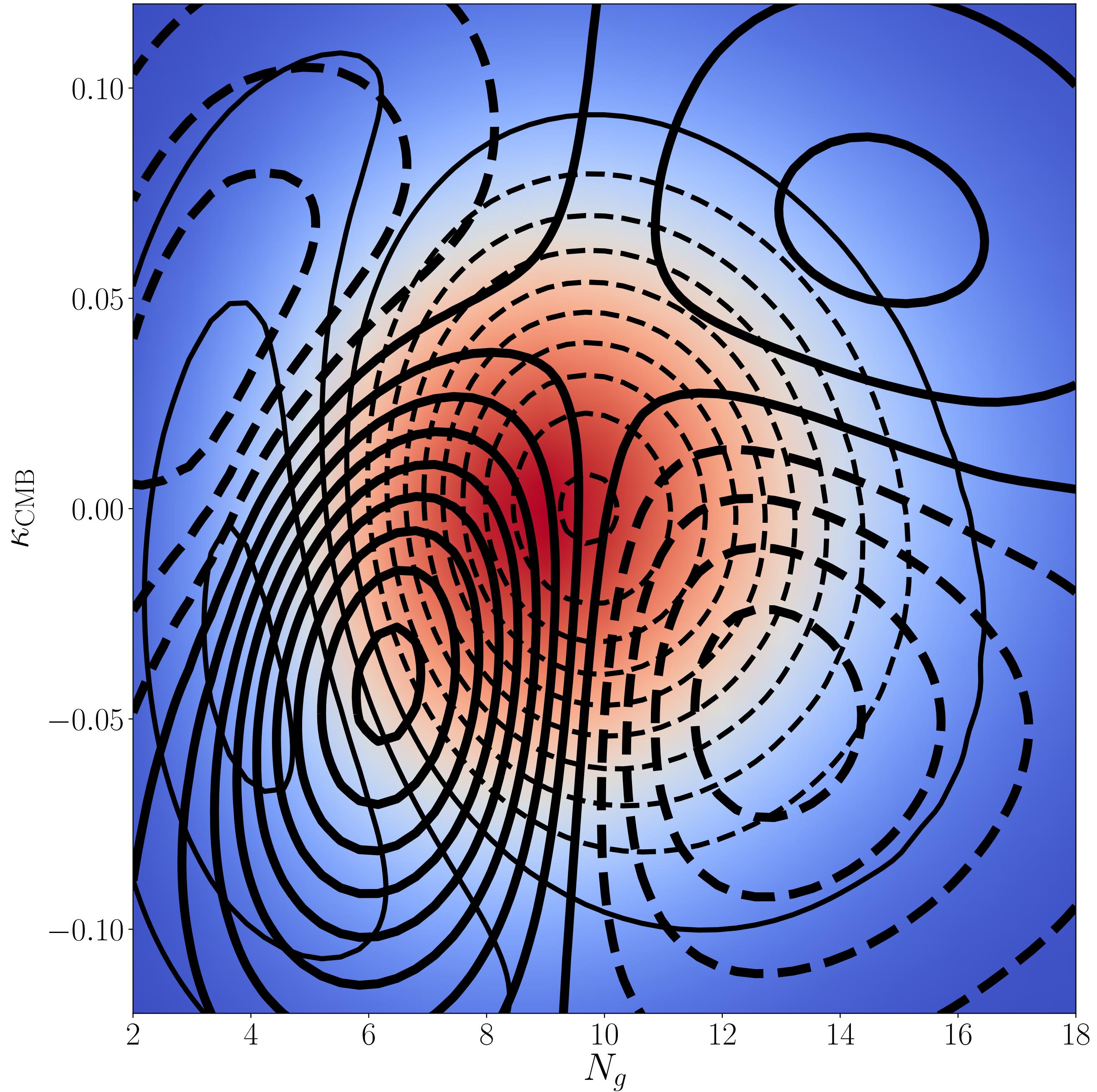


Remember:

at a single scale 2-point statistics only extracts 3 numbers & could never measure 6 properties of the cosmos at once

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The core of PDF theory

Remember that

$$e^{\varphi_R(\lambda)} = \int d\delta_R \ p(\delta_R) \ e^{\lambda\delta_R} \equiv \boxed{\langle e^{\lambda\delta_R} \rangle}$$
$$= \int \mathcal{D}\delta_{\text{lin}} \ e^{\lambda\delta_R[\delta_{\text{lin}}]} \ \mathcal{P}[\delta_{\text{lin}}]$$

The core of PDF theory

Remember that

$$e^{\varphi_R(\lambda)} = \int d\delta_R \ p(\delta_R) \ e^{\lambda\delta_R} \equiv \boxed{\langle e^{\lambda\delta_R} \rangle}$$

$$= \int \mathcal{D}\delta_{\text{lin}} \ e^{\lambda\delta_R[\delta_{\text{lin}}]} \ \mathcal{P}[\delta_{\text{lin}}]$$

$$\sim \int \mathcal{D}\delta_{\text{lin}} \ \mathcal{D}J_{\text{lin}} \ e^{-S_\lambda[\delta_{\text{lin}}, J_{\text{lin}}]}$$

with $S_\lambda[\delta_{\text{lin}}, J_{\text{lin}}] \equiv -\lambda\delta_R[\delta_{\text{lin}}] + iJ_{\text{lin}} \cdot \delta_{\text{lin}} - \Phi[iJ_{\text{lin}}]$.

The core of PDF theory

Remember that

$$e^\varphi$$

$$\Phi[J] \equiv \sum_n \int \prod_{i=1}^n d^3x_i \xi_{n,\text{lin}}(\mathbf{x}_1, \dots, \mathbf{x}_n) \frac{J(\mathbf{x}_1) \dots J(\mathbf{x}_n)}{n!}$$

$$\sim \int \mathcal{D}\delta_{\text{lin}} \mathcal{D}J_{\text{lin}} e^{-S_\lambda[\delta_{\text{lin}}, J_{\text{lin}}]}$$

$$\text{with } S_\lambda[\delta_{\text{lin}}, J_{\text{lin}}] \equiv -\lambda \delta_R[\delta_{\text{lin}}] + i J_{\text{lin}} \cdot \delta_{\text{lin}} - \boxed{\Phi[i J_{\text{lin}}]} .$$

The core of PDF theory

Remember that

$$e^{\varphi_R(\lambda)} = \int d\delta_R \ p(\delta_R) \ e^{\lambda\delta_R} \equiv \langle e^{\lambda\delta_R} \rangle$$

$$= \int \mathcal{D}\delta_{\text{lin}} \ e^{\lambda\delta_R[\delta_{\text{lin}}]} \ \mathcal{P}[\delta_{\text{lin}}]$$

$$\sim \int \mathcal{D}\delta_{\text{lin}} \ \mathcal{D}J_{\text{lin}} \ e^{-S_\lambda[\delta_{\text{lin}}, J_{\text{lin}}]}$$

with $S_\lambda[\delta_{\text{lin}}, J_{\text{lin}}] \equiv -\lambda \boxed{\delta_R[\delta_{\text{lin}}]} + iJ_{\text{lin}} \cdot \delta_{\text{lin}} - \Phi[iJ_{\text{lin}}]$.

The core of PDF theory

Approximate integral at saddle point configuration

- i.e. for **fields that minimise action!**

$$\Rightarrow \varphi_R(\lambda) \approx -S_\lambda[\delta_{\text{lin}}^*, J_{\text{lin}}^*]$$

The saddle point is spherically symmetric!

$$\Rightarrow \delta_R[\delta_{\text{lin}}] \equiv \text{spherical collapse}$$

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functional determinant of
+ Hessian of action 😱
(cf. Valageas 2005,
Ivanov++ 2019)

What we are
really doing...

$$\tilde{\varphi}_R(\lambda) = \sum_n \frac{\langle \delta_R^n \rangle_c}{\langle \delta_R^2 \rangle_c^{n-1}} \frac{\lambda^n}{n!} = \sum_n S_n \frac{\lambda^n}{n!}$$

$$\tilde{\varphi}_R(\lambda) = \tilde{\varphi}_R(\lambda \cdot \langle \delta_R^2 \rangle_c) / \langle \delta_R^2 \rangle_c$$

You can do these calculations yourself with **CosMomentum**:

<https://github.com/OliverFHD/CosMomentum>

Also, section 4.6 of <https://arxiv.org/abs/1912.06621> describes
step-by-step details of numerical implementation!

Conclusions / Outlook

- Can open up complete “plane of perturbations” for LSS studies
- complete LCMD analysis possible even with complicated bias model and simple analysis setup
- Much more!
(see e.g. primordial non-Gaussianity in **Friedrich ++ 2020** or neutrino physics from PDF in **Uhlemann, Friedrich et al. 2020** ; **Boyle, Uhlemann, Friedrich et al 2021**)
- Most of these tools are already publicly available in the **CosMomentum** package:
<https://github.com/OliverFHD/CosMomentum>
 - 3D PDFs and cumulant generating functions
 - PDF of gravitational lensing convergence
 - 2D and 3D halo counts-in-cells PDFs

