# Quantum Euler angles and agency-dependent space-time

Giuseppe Fabiano – University of Naples «Federico II»

In collaboration with:

G. Amelino-Camelia, V. D'Esposito, D. Frattulillo, P. Hoehn, F.

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## Quantum space-time and symmetries

- Theoretical investigations in quantum gravity suggest that space-time itself should acquire «quantum features»
- Relativistic transformations also acquire quantum features, in order for the quantum space-time properties to be valid for every observer
- Observers are connected by quantum group transformations
- As a case study, we will consider the  $SU_q(2)$  quantum group, to investigate purely rotated systems.

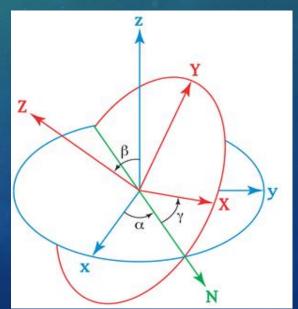
## SU(2) coordinatization and Euler Angles

 In classical and quantum mechanics, rotation transformations are governed by the group SU(2)

$$SU(2) \ni U = \begin{pmatrix} a & -c^* \\ c & a^* \end{pmatrix} \quad a, c \in \mathbb{C} : |a|^2 + |c|^2 = 1$$
$$a = e^{i\chi} \sin\left(\frac{\theta}{2}\right) \quad c = e^{i\phi} \cos\left(\frac{\theta}{2}\right)$$

SU(2) parameters and Euler Angles

$$\begin{cases} \theta = \beta \\ \chi = \frac{\alpha + \gamma}{2} \\ \phi = \frac{\pi}{2} - \frac{\alpha - \gamma}{2} \end{cases}$$



## Link between SU(2) and SO(3)

• The connection between SU(2) and classical rotations is established via the canonical homomorphism with SO(3).

$$R = \begin{pmatrix} \frac{1}{2}(a^2 - c^2 + (a^*)^2 - (c^*)^2) & \frac{i}{2}(-a^2 + c^2 + (a^*)^2 - (c^*)^2) & a^*c + c^*a \\ \frac{i}{2}(a^2 + c^2 - (a^*)^2 - (c^*)^2) & \frac{1}{2}(a^2 + c^2 + (a^*)^2 + (c^*)^2) & -i(a^*c - c^*a) \\ -(ac + c^*a^*) & i(ac - c^*a^*) & 1 - 2cc^* \end{pmatrix}$$

## $SU_q(2)$

• Parameters become the generators of  $C_q ig( SU(2) ig)$ , the algebra of complex functions on SU(2)

$$\begin{pmatrix} a & -c^* \\ c & a^* \end{pmatrix} \Rightarrow \begin{pmatrix} a & -qc^* \\ c & a^* \end{pmatrix} \qquad a, c \in C_q(SU(2))$$

endowed with a non-commutative product realized by

$$ac = qca$$
  $ac^* = qc^*a$   $cc^* = c^*c$ 

$$c^*c + a^*a = 1$$
  $aa^* - a^*a = (1 - q^2)c^*c$ 

• q is a «small» deformation parameter, larger than 0 and close to 1.

## Homomorphism between $SU_q(2)$ and $SO_q(3)$

- $C_q(SO(3)) \coloneqq C_q(SU(2)/Z_2)$ , realizing the q-analogue of the SU(2) to SO(3) homomorphism
- A 3x3 matrix representation is given by

$$R_{q} = \begin{pmatrix} \frac{1}{2}(a^{2} - qc^{2} + (a^{*})^{2} - q(c^{*})^{2}) & \frac{i}{2}(-a^{2} + qc^{2} + (a^{*})^{2} - q(c^{*})^{2}) & \frac{1}{2}(1 + q^{2})(a^{*}c + c^{*}a) \\ \frac{i}{2}(a^{2} + qc^{2} - (a^{*})^{2} - q(c^{*})^{2}) & \frac{1}{2}(a^{2} + qc^{2} + (a^{*})^{2} + q(c^{*})^{2}) & -\frac{i}{2}(1 + q^{2})(a^{*}c - c^{*}a) \\ -(ac + c^{*}a^{*}) & i(ac - c^{*}a^{*}) & 1 - (1 + q^{2})cc^{*} \end{pmatrix}$$

• This is not a real valued matrix anymore, it contains operators

## $SU_q(2)$ representations

• The Hilbert space containing the two unique irreducible representations of the  $SU_q(2)$  algebra is  $H=H_\pi \oplus H_\rho$ , where  $H_\pi=L^2(S^1) \otimes L^2(S^1) \otimes \ell$  and  $H_\rho=L^2(S^1)$ 

• 
$$\rho(a)|\eta\rangle = e^{i\eta}|\eta\rangle;$$
  $\rho(a^*)|\eta\rangle = e^{-i\eta}|\eta\rangle;$   $\rho(c)|\eta\rangle = 0;$   $\rho(c^*)|\eta\rangle = 0;$ 

• 
$$\pi(a)|n,\delta$$
,  $\epsilon\rangle=e^{i\epsilon}\sqrt{(1-q^{2n})}|n-1,\delta$ ,  $\epsilon\rangle$ ;  $\pi(a^*)|n,\delta$ ,  $\epsilon\rangle=e^{-i\epsilon}\sqrt{(1-q^{2n+2})}|n+1,\delta$ ,  $\epsilon\rangle$ ;

• 
$$\pi(c)|n,\delta$$
,  $\epsilon\rangle = e^{i\delta}q^n|n,\delta$ ,  $\epsilon\rangle$ ; 
$$\pi(c^*)|n,\delta$$
,  $\epsilon\rangle = e^{-i\delta}q^n|n,\delta$ ,  $\epsilon\rangle$ ;

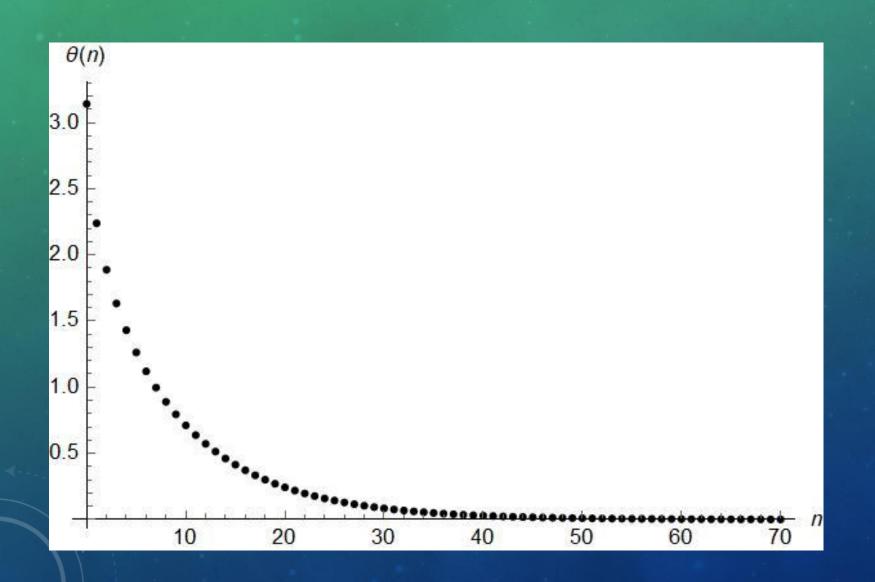
• 
$$a = e^{i\chi} \cos\left(\frac{\theta}{2}\right)$$
  $c = e^{i\phi} \sin\left(\frac{\theta}{2}\right)$  (Classical case)

## **Quantum Euler Angles (1)**

- We promote the SU(2)-Euler Angles relations to the quantum case.
- Comparing the phases of  $\alpha$  and c to their classical analogues, we identify  $\epsilon$  with  $\chi$  and  $\delta$  with  $\phi$ . They are continuous and play the same role as before.
- Exploiting the fact that c is a diagonal operator

$$q^n = Sin\left(\frac{\theta(n)}{2}\right) \leftrightarrow \theta(n) = 2Arcin(q^n)$$

## **Quantum Euler Angles (2)**



$$\theta(n) = 2Arcin(q^n)$$
q=0.99

## **Physical interpretation and Quantum rotations**

- A state  $|\psi\rangle \in H$  is representative of the relative orientation between two reference frames, A and B.
- Our interpretation is that the mean value of  $R_q$  on  $|\psi\rangle$  will give an estimate of the entries of the rotation matrix that connects A and B

$$\langle \psi | R_q | \psi \rangle_{ij}$$

 However, due to non-commutatitvity, we will have a non vanishing variance for the matrix elements, in general:

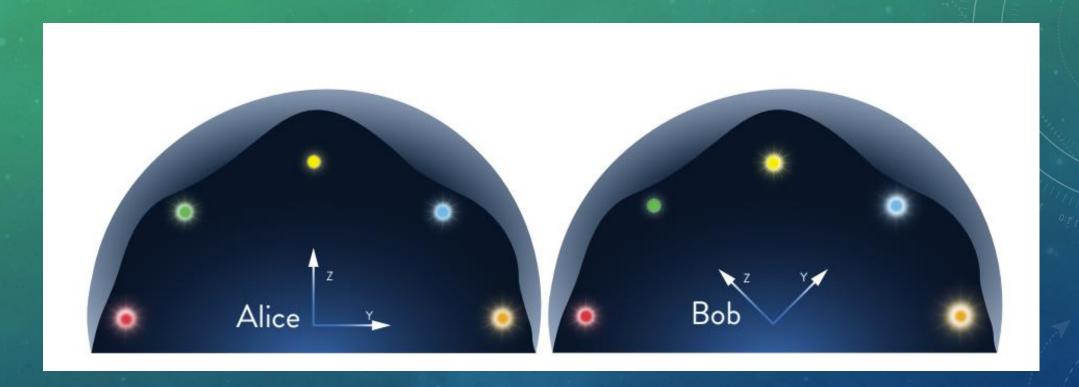
$$\Delta_{ij} = \sqrt{\langle \psi | R_q^2 | \psi \rangle_{ij} - \langle \psi | R_q | \psi \rangle_{ij}^2}$$

## **Examples of Quantum rotations**

• Basis states in representation  $\rho$  have  $\Delta_{ij}=0$  and the mean value of  $R_q$  on such states gives sharp rotations around the z-axis.

• Superpositions of basis states in representation  $\pi$  yield non-zero values of  $\Delta_{ij}$  and the mean value of  $R_q$  gives rise to deformed rotation matrices, which reduce to standard rotation matrices about an axis in the x-y plane in the commutative limit

## Same stars, different skies



• Fuzziness of space-time points depends on the choices made by the observer. In this sense, the reconstructed space-time is agency-dependent.



#### **Example: rotation around the z-axis**

• Consider a state  $|\chi\rangle$  in representation  $\rho$ . The mean value of the rotation matrix is:

$$\langle \chi | R_q | \chi \rangle_{ij} = \begin{pmatrix} \cos(2\chi) & -\sin(2\chi) & 0 \\ \sin(2\chi) & \cos(2\chi) & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

• It coincides with a standard SO(3) rotation matrix. Indeed, computing the uncertainties, we have

 $\Delta_{ij} = 0 \rightarrow \text{Sharp rotations around the z-axis}$ 

### «Physical» states construction

• To effectively describe rotations' deformations, we demand that our states of geometry  $|\psi\rangle$  satisfy

$$\langle \psi | R_q | \psi \rangle_{ij} \rightarrow (R_{ij})$$
  $\Delta_{ij} \rightarrow 0$  when  $q \rightarrow 1$ 

where  $(R_{ij})$  are the entries of a classical rotation matrix.

• Since  $(\phi, \chi)$  behave as in the classical case, we must look for states of the form

$$|\psi\rangle = \sum_{n=0}^{\infty} c_n |n, \phi, \chi\rangle$$

heavily weighted around  $\bar{n}$  and which satisfy the criteria above, to properly describe a rotation deformation of Euler angles  $(\phi, \chi, \theta(\bar{n}))$ 

### Example: rotation of $\pi$ around the x-axis

• Consider the state  $|\psi\rangle=|n;\chi;\phi\rangle=\left|0;\frac{\pi}{2};0\right\rangle$ . The relevant quantities, working at first order in (1-q)

$$\langle \psi | R_q | \psi \rangle = \begin{pmatrix} 1 - (1 - q) & 0 & 0 \\ 0 & -1 + (1 - q) & 0 \\ 0 & 0 & -1 + 2(1 - q) \end{pmatrix} + o(1 - q)$$

$$\langle \psi | \Delta R_q | \psi \rangle = \begin{pmatrix} \sqrt{2}(1-q) & \sqrt{2}(1-q) & \sqrt{2}(1-q) \\ \sqrt{2}(1-q) & \sqrt{2}(1-q) & \sqrt{2}(1-q) \\ \sqrt{2}(1-q) & \sqrt{2}(1-q) & 0 \end{pmatrix} + o(1-q)$$

 As q → 1, these correctly reproduce a rotation of π around the x-axis with null uncertainty.