





# Status of the GINGER Project

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- ✓ GINGER, t<sub>0</sub> February 2023
- ✓ A bit of history
- ✓ Fundamental physics
- ✓ Near future plan
- ✓ Conclusions





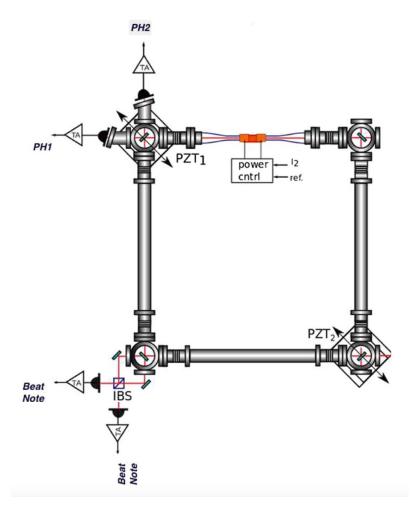
## **GINGER**



GINGER is an array of high sensitivity laser gyroscopes (Ring Laser Gyroscopes, RLG). RLG are based on a square active cavity where two counter-propagating modes circulate. The two modes are in general highly symmetric, but non reciprocity is present and the RLG signal gathers information on non reciprocity on the light propagation in different directions. Very small variations are expected in General relativity, the larger being due to the de Sitter and Lense Thirring effect on the Earth surface, Lorentz violation tests on the gravity sector in the Standard Model extended formalism, and on the non classical electromagnetism. The larger effect is certainly the Sagnac effect, which links the RLG response with the absolute rotation rate of the optical cavity. This is a large effect on the Earth surface, and at present RLG has top sensitivity to measure Earth absolute angular rotation rate in order to investigate any tiny deviations.

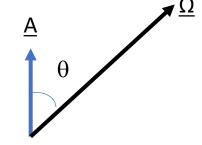
GINGER is highly interdisciplinary and able to provide data not only for fundamental physics but also for geophysics and geodesy. The validity of GINGER for fundamental physics investigation requires long time measurements and very high sensitivity, the relevant target being to reach and go behind the sensitivity of 1 part in 10° of the Earth rotation rate, target that has been already demonstrated.

### **RLG**



- a physical system composed of an active medium inside a polygonal optical cavity, such to generate two counterpropagating beams (traveling waves)
- Differences between the two directions of propagation, non reciprocity, generate small differences between the two beams, that interferometry can detect
- The Sagnac effect is due to the non reciprocity between the two waves when the cavity and the active medium rigidly rotate
- scalar product between the vector  $oldsymbol{arOmega}$  and the area vector  $oldsymbol{A}$

$$f_s = 4\frac{A}{\lambda L}\Omega\cos\theta$$

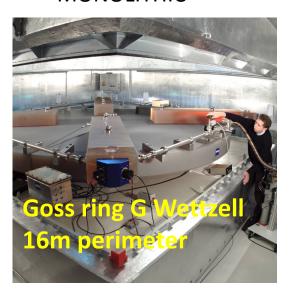




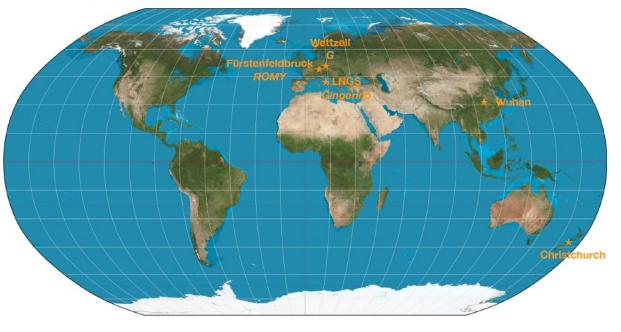


# Large frame RLGs

#### **MONOLITHIC**

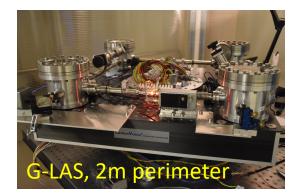


### Large-frame optical gyroscopes in the world



#### HETEROLITHIC



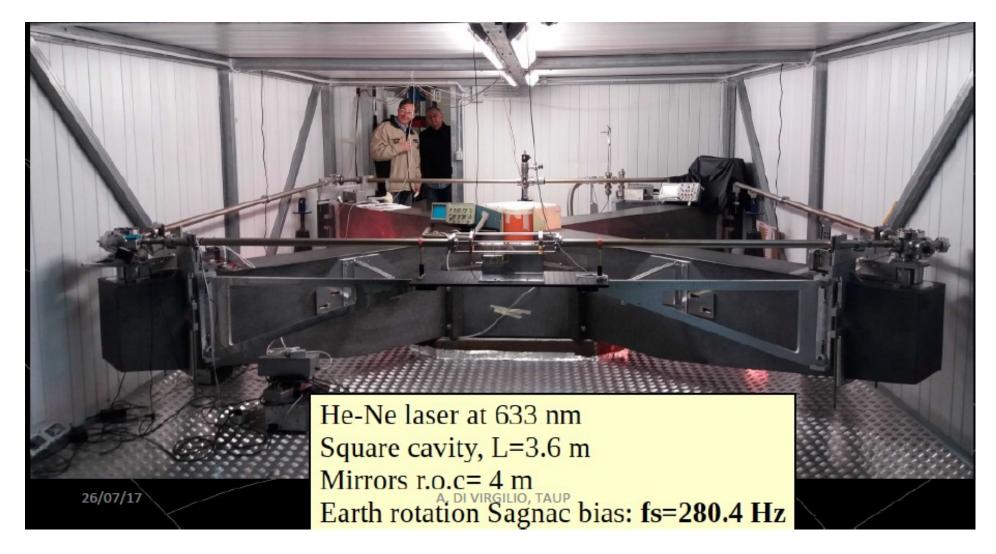






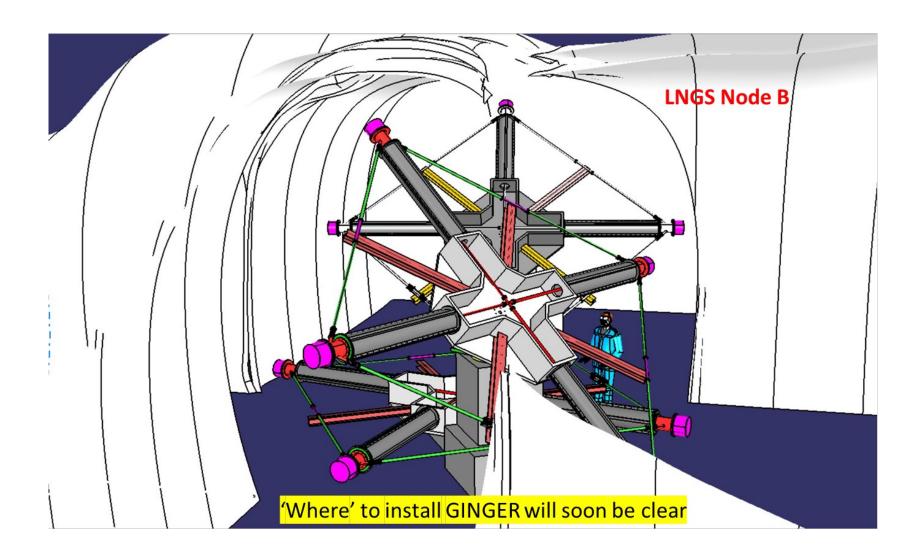


## GINGERINO









Array of 3 ring laser gyroscopes (RLG), 5m sides, the drawing is related to Node B of LNGS
This area is large enough to separate most of the electronics in a separate room.

**RLX** oriented at maximum Sagnac Signal, to measure the absolute value of the angular rotation

**RLH** with vertical area vector

**RLO** with area vector outside the meridian plane





- INFN Pisa and LNL: coordination of the Sagnac frequency reconstruction;
- Naples: optics simulation and quantum noise;
- LNL, Naples , Salerno and Turin: interface with fundamental physics;
- INGV: DAQ and remote control of the apparatus;
- INGV: interface with geophysics analysis;
- UNIVAQ: mechanics simulation and test.



INGV



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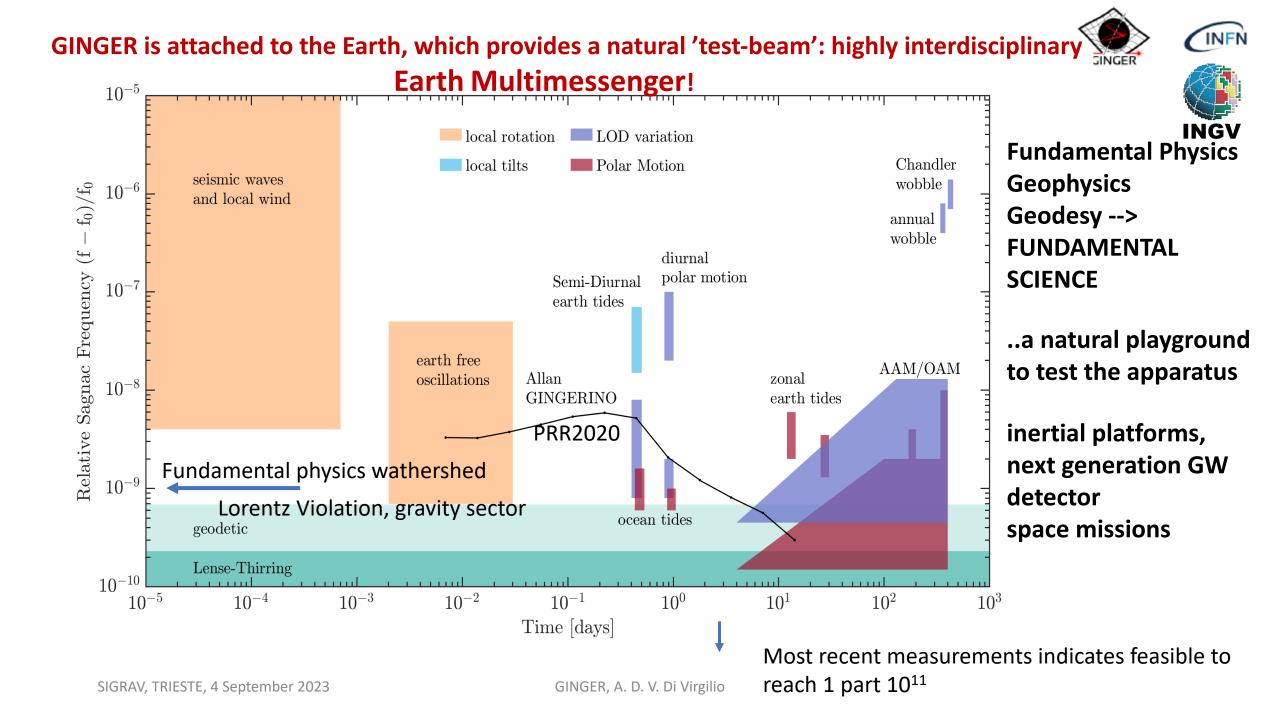
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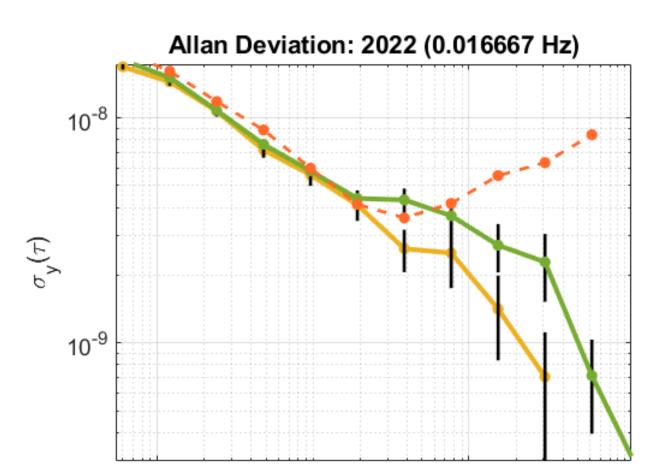
• To observe the environment in which we are immerse brings the advantage to compare our data with independent apparatus, very important to validate the analysis procedure

 Comparison GINGERINO and GNSS data: https://arxiv.org/abs/2308.01277





# G crosses 1 part 10<sup>9</sup> boundary



https://doi.org/10.1140/epjc/s10052-022-10798-9

10<sup>2</sup>

au [sec]

10<sup>4</sup>

## Impact on fundamental physics

Large discussion can be found in GINGER
 <a href="https://arxiv.org/pdf/2209.09328.pdf">https://arxiv.org/pdf/2209.09328.pdf</a>, which will soon appear in MEMOCS

upgrades can be found in the presentations of the IV GRM workshop in Pisa, 14-16 June:

https://agenda.infn.it/event/33962/timetable/?layout=room#2023061 4.detailed

Local measurement, not averaged, gravity map not necessary

## **Perspectives**

In  $f(R, R^{\mu\nu}R_{\mu\nu}, \phi)$  gravity, GP-B and LARES satellites provide

$$m_V > 7.3 \times 10^{-7} m^{-1}$$
  $m_Y > 1.2 \times 10^{-6} m^{-1}$ 

constraint on  $m_{
m v}$  by GINGER

constraints on  $a_1$ ,  $a_2$  by GINGER

In Horava-Lifshitz gravity, the weak-field limit provide

$$c \, \delta \tau = \frac{4A\Omega_E}{c} \left[ \cos(\theta + \alpha) - \left( 1 + \frac{G}{G_N} a_1 - \frac{a_2}{a_1} \right) \frac{GM}{c^2 R} \sin \theta \sin \alpha \right]$$
$$- \frac{GI_E}{c^2 R^3} \left( 2\cos \theta \cos \alpha + \sin \theta \sin \alpha \right)$$

## **Perspectives**

$$\Omega_S = \frac{4A}{P\lambda} \Omega_E \left[ \cos(\theta + \alpha) - \left( 1 + \frac{G}{G_N} a_1 - \frac{a_2}{a_1} \right) \frac{GM}{c^2 R} \sin \alpha \sin \theta - \frac{GI_E}{c^2 R^3} \left( 2\cos \theta \cos \alpha + \sin \theta \sin \alpha \right) \right]$$

- Fixing a<sub>1</sub> and a<sub>2</sub> by GINGER allows to retain or reject viable theories
- GINGER could select effective models for Quantum Gravity in the weak field limit
- With respect to satellite experiments, results can be tuned and reproduced.

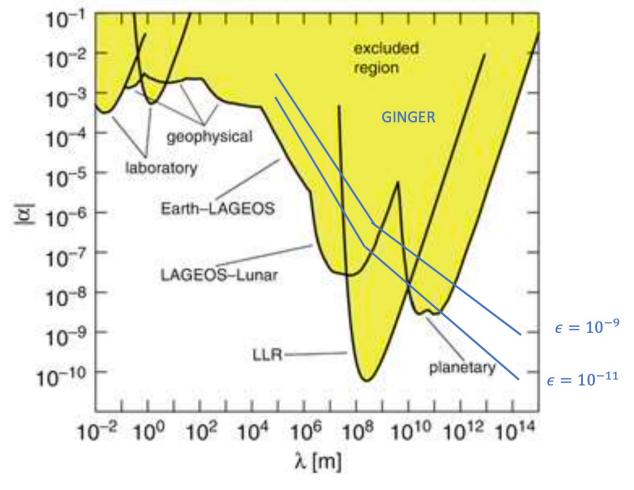
S. Capozziello, C. Altucci, F. Bajardi, A.Di Virgilio et al... Euro. Phys. J. Plus 136 (2021) 5

A.Di Virgilio, U. Giacomelli, A. Simonelli et al... Euro. Phys. J. C 81 (2021) 457

A. Porzio, C. Altucci, S. Capozziello, R. Velotta, et al... PoS Corfù 2017 (2018) 181

# Yukawa-like corrections (GINGER) GINGER α λ exclusion plot

fixing the compatibility ranges of Yukawa-like parameters can be a fundamental tool to discriminate among viable gravitational models.



# GINGER is part of UGGS







(Underground Geophysics at Gran Sasso)

- High sensitivity seismometers
- Monitoring of acquifer wather pressure
- Gravimeter
- GNSS antennas on the Gran Sasso area
- GINGER





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r r r t t

In 10 years we have published more than 60 papers, spanning from theory, completely new RLG data analysis and experimental set up, noise investigation and geophysics, turning the RLG from a not acceptable instrument because laser non linearity to fundamental tool for angular rotation investigation.

# GINGER is based on a single device, different from most experiments

- Focus on lower frequency, as a consequence it needs: sensitivity, accuracy and <u>stability</u>
- based on highly symmetric interferometer: it can operate free running
- It is based on frequency reconstruction—> not affected by electronic 1/f noise and has huge dynamic range
- The laser cavity plays a crucial role, it is rigid; effective limitations are deformations, that can cause 'rotation' of the optical cavity  $\rightarrow$  difficult to study with fine elements analysis, avoid thermal gradients and use uniform materials; structural analysis of the mechanical system for second generation retrofitting

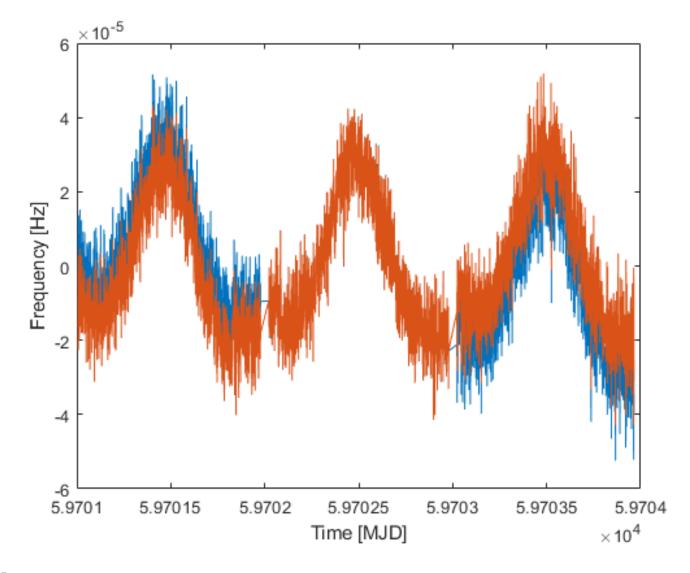
# RLG: ideal instruments for low frequency investication (0-1Hz)

# ...angular rotation are ideal for low frequency investigation...

Semidiurnal tides are usually dominating the low frequency signals, but the diurnal polar motion dominates the RLG signal

This is the signal of G Wettzell, based on a perfectly rigid cavity.

We have to obtain the same with hetero lithic RLG structure.



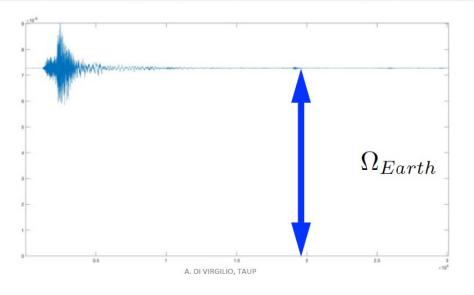
Overcoming 1 part in  $10^9$  of earth angular rotation rate measurement with the G Wettzell data A. D. V. Di Virgilio¹, G. Terreni¹, A. Basti¹², N. Beverini², G. Carelli², D. Ciampini², F. Fuso², E. Maccioni¹², P. Marsili², J. Kodet¹ and K. U. Schreiber³



- Electromagnetic radiation at the equilibrium with atomic gass in a high Q optical cavity
- The two counterpropagating modes are 'almost' equal -> the typical noise sources disappear as common mode
- It is based on frequency measurements -> very robust method and huge dynamic range

GINGERINO CAN DETECT VERY HIGH ANGULAR ROTATION SIGNALS

The Visso M 5.9 earthquake, probably the largest seismic rotational signal ever recoded







## Key point: Earth Rotation rate subtraction

$$f = \frac{4 A}{p \lambda} \Omega \cos \vartheta \rightarrow \frac{f - f_0}{f_0} < 10^{-9} (10^{-11})$$

- $\frac{4 A}{v \lambda}$  geometrical scale factor  $\rightarrow$  parameter **X0**
- in  $\theta(t) = \theta_0 + \delta\theta(t)$ ....error  $\delta\theta < 2 \ 10^{-9} \ rad (2 \ 10^{-12} \ rad !) -> parameter <math>\theta 0$

### GINGER design is such that X0 and $\theta$ 0 acts at second order

• Di Virgilio, A.D.V., Belfi, J., Ni, WT. et al. GINGER: A feasibility study. Eur. Phys. J. Plus 132, 157 (2017). https://doi.org/10.1140/epjp/i2017-11452-6







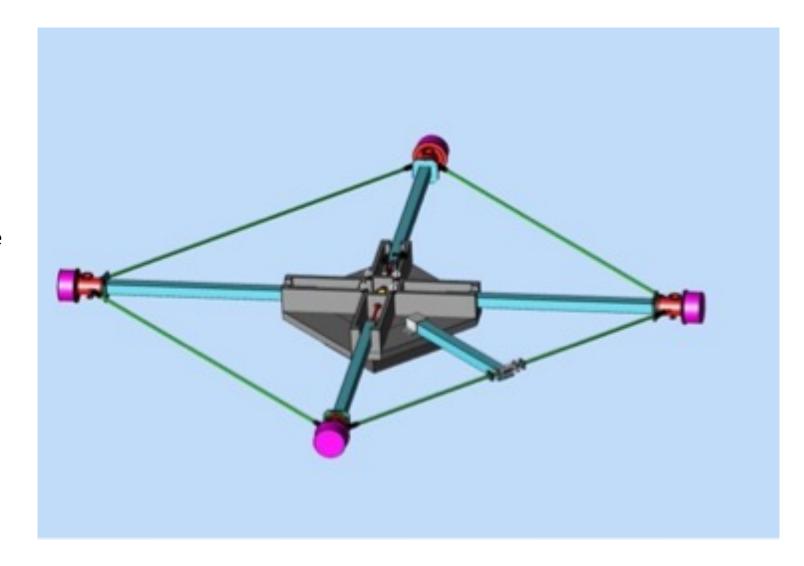
## HL RLG our general rules

- Whole apparatus contained in a thermal bath
- Natural termal stability of the cave around fractions of degrees for long time basis. If necessary actively improve thermal stability
- Attached to the floor with a single rigid monument in granite. Granite is a reasonably homogeneous material which can be well machined.
- Isolate the mirrors from external disturbances and reduce couplings among cavity mirrors
- Components machined with high precision in order to have the center of the mirrors on the vertex of a regular square with fractions of mm accuracy.

### RLG side 4m

GP3 is the model of the HL RLG developed for GINGER. Mechanical parts in titanium to reduce weight in the perimetral part

The structure looks rather slim, but we have verified that it is rigid enough thanks to the SiC properties. Static arrows are at the level of tens of microns



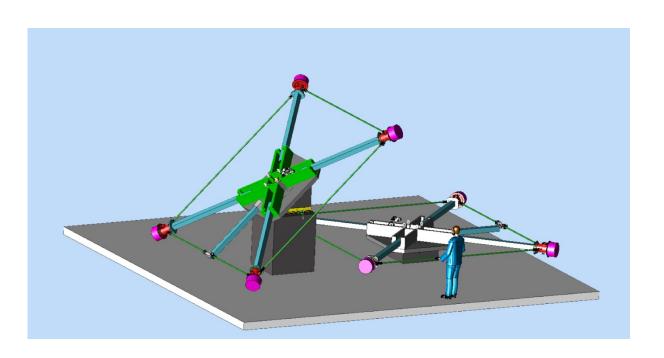






# GINGER: RLX & RLH

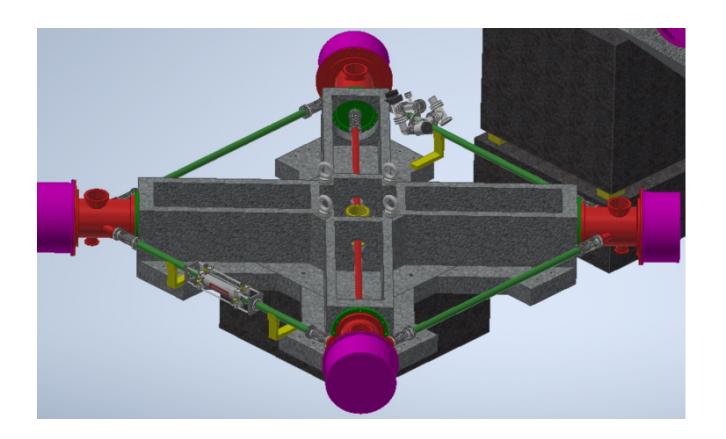
- Installation at LNGS 4 m side
- It would provid the orientation angle reconstruction and the test of the mechanics, to be compared with GINGERINO and GP2 data
- Lack of redundancy, very relevant for accuracy, important for fundamental physics



## TRIO (Transportable Ring Laser Observatory, GP3 model)

TRIO is a 1.5 m side RLG wich uses the mechanical scheme of GINGER. It is under construction, expected to be completed in fall.

It will provide a suitable test of the GP3 mechanics for GINGER.



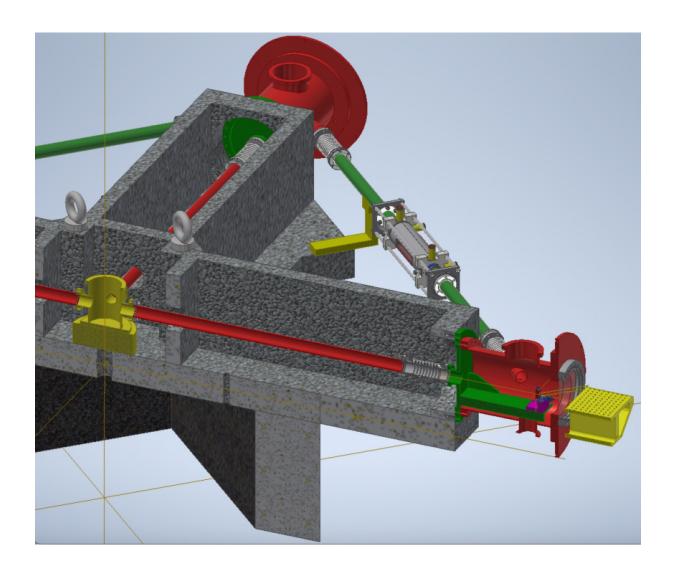
### GP3

The same scheme will be used for GINGER, adding spacers in SiC in order to extenbd up to 4m the size of the square

commercially available components as:

High stability mirror holders equipped with actuators, UHV

Decouple the mirrors from the perimeter pipes using bellows as dumpers



TRIO -> test of GP3

TRIO will provide a suitable test of the GP3 concept.

TRIO will be tested in Pisa and we will compare the results with GP2 typical behaviour









- Attach the monument to the floor, alignment of RLX at 100μrad level
- Install mechanical parts, high vacuum required to avoid gas contamination
- Install mirrors and fill the tank with special gas mixture
- Align the cavity
- Start the laser
- Install read out for the beat note and for the monobeams
- Install photodiode to monitor the plasma
- Take and store data
- Analyse the data

**RED: optics experts activity** 







# Spreading out competences during installation

- Care is required for mirrors handling, installation and cavity alignment
- Online tools to control the apparatus have been developed and tested with GP2 and GINGERINO

# WE ARE LOOKING FOR YOUNG RESEARCHERS WILLING TO BECOME RLG EXPERTS





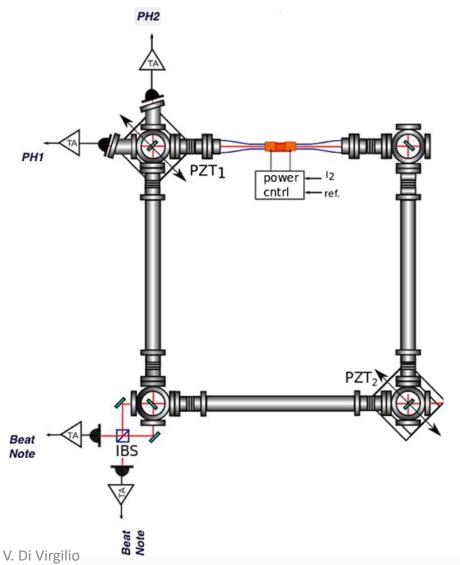




## Limiting noise measurement

The beat note is recorded independently at the two utput ports of the cube beam splitter.

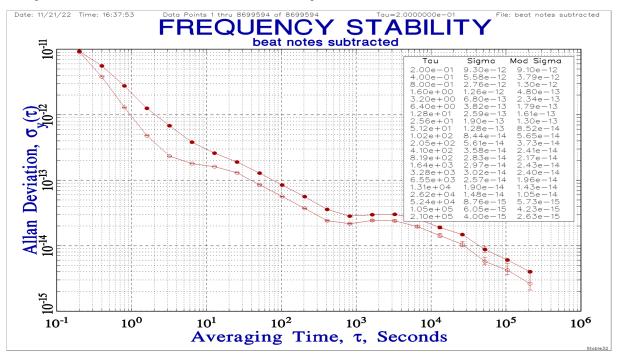
They provide two independent measurement of the RLG output







- This result confirms the validity of 1 part in 10<sup>11</sup> of the Earth rotation rate
- It is important to remark that the 1 part in 10<sup>9</sup> of the Earth rotation would be in any case new and important for fundamental science









## conclusions

- The GINGER Project has started at the beginning of 2023, in collaboration with INGV
- Shortly we will discuss with LNGS the location and the installation details
- Details of the drawings will be reviewed
- Main orders will be placed asap
- The plan is to built it in 2024 and start taking data in 2025